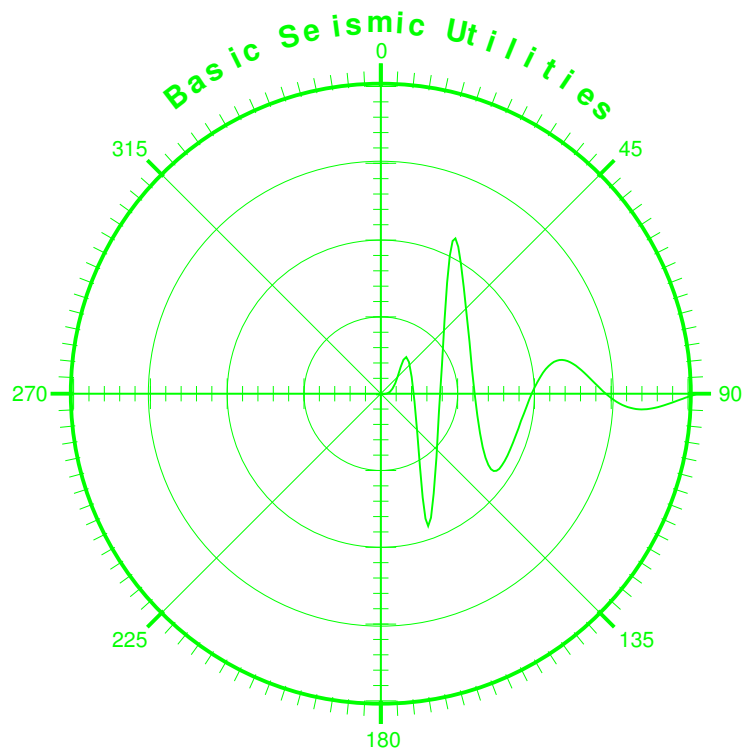


# Running Basic Seismic Utilities (BSU)

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Dr. P. Michaels, PE  
<pmsolid@cablene.net>  
<paulmichaels@boisestate.edu>

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Version 3.0.2



*Michaels Engineering Geophysics  
Boise, Idaho*

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Codes licensed to the public domain included in BSU are xdrfloa.c (IBM License, 13) and Fortran functions rand.f and runif.f from CMLIB, provided to the public from NIST. <http://gams.nist.gov/serve.cgi/Packages>

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## 1 Acknowledgements

This software is an updated version of the revised release that was done in April 2018. The user guide for version 3.0.1 is still largely valid, and this document fills the need to guide a user to which utilities might be of use to them. Building on earlier versions, the goal has been portability. That is, while some new programs have been added, much of the software has been carried forward, and thus is now available for use on a variety of operating systems. Thus, I remain indebted to the work of many others in the development of this package. I would like to thank Enders A. Robinson and the Holden-Day Inc., Liquidation Trust (1259 S.W. 14th Street, Boca Raton, FL 33486, Phone: 561.750-9229 Fax: 561.394.6809) for license to include and distribute under the GNU license subroutines found in Dr. Robinson's 1967 book [Robinson \(1967\)](#), [Multichannel time series analysis with digital computer programs](#). This book is currently out of print, but contains a wealth of algorithms, several of which I have found useful and included in the BSU Fortran77/gfortran subroutine library (sublib4.a). This has saved me considerable time.

In other cases, subroutines taken from the book [Numerical Recipes Press et al. \(1989\)](#) had to be replaced (the publisher did not give permission to distribute). While this is an excellent book, and very instructional for those interested in the theory of the algorithms, future authors of software should know that the algorithms given in that book are NOT GNU. Replacement software was found in the *GNU Scientific Library (GSL)*, and in the *CMLIB*.

For plotting, I remain indebted to the developers of *PLPLOT*. *PLPLOT* credits have grown to be too many to list. However, there are a number of instances where I ran into dependency problems with some operating systems, particularly the Microsoft family. So I have added *GNU PLOT* alternatives.

Where there was a need to solve for eigenvalues, or invert a matrix, I have relied on *LAPACK*. This excellent package is well worth installing, and I acknowledge the contributions of the many authors of *LAPACK* and *BLAS*.

Lastly, the author acknowledges financial support over the years from various clients. Financial support included that provided by grant DAAH04-96-1-0318 from the U.S. Army Research Office. Views and conclusions contained herein are those of the author and should not be interpreted as necessarily representing the official policies or endorsements, either expressed or implied, of the Army Research Office or the U.S. Government. This material is also based upon work supported by the National Science Foundation under Grant No. 0321233. Any opinions, findings, and conclusions or recommendations expressed in this material are those of the author and do not necessarily reflect the views of the National Science Foundation. Other support has been provided by the Idaho State Board of Education, Boise State University Sabbatical Committee, Idaho Transportation Department, and Idaho Power. Again, views and conclusions contained herein are those of the author and should not be interpreted as necessarily representing the views of those who have provided the author support.

## 2 Conventions

- **Capitalization** In the text of this document, program names are often typed in upper case letters. This is to make the name stand out. However, when actually executing a program, use lower case letters for command line execution in a terminal. *Linux is case sensitive*.
- **File Naming** BSU programs expect file names with no more than 12 characters. If a file xxxx.seg is input to a program byyy, then the output data file will be byyyxxxx.seg. Naming scrolls right. So if data file byyyxxxx.seg is input to program bzzz, the output will be named bzzzbyyy.seg. This is predictable and a bit of a pain at times. It goes back to being able to write bash scripts and be able to predict the names of output for later stages in a sequence of processes. If you input a file with a long name, then you will get this message:  
**[ABORT] Input length violation [13]**
- **File Suffixes** Seismic data in BSEGY format are \*.seg files. Seismic data in SEG-Y format are \*.sgy files. Files in SEG-2 format are \*.DAT files. Bison files don't have a suffix (unless compressed as \*.gz) and are 8 characters long, typically upper case alpha numeric. Some BSU programs will output *GNU PLOT* \*.gp files.

### 3 Converting Between Data Formats

There are two types of seismic data formats.

- Data Exchange Formats
- Data Processing Formats

**3.0.0.1 Exchange Formats** These are formats meant to publicly defined so that users may exchange data regardless of what processing software they may ordinarily use. Examples would include SEG-Y, SEG-2, Bison, and SEG-D. The professional society, Seismic Exploration Geophysicists (SEG) has created all of these except the Bison format which was developed by makers of an engineering seismograph. For example, SEG-2 was developed for personal computers, Pullan (1990). There are other types of data exchange formats. For example, for Computer Aided Drafting, there is DXF (data exchange format). BSU software also can convert header information into DXF format which can then be used to make base maps.

Bison format is really just the format output by Bison seismographs. It can be integer or floating point. However, the floating point is extremely unusual, and I know of no other software than BSU which can process the Bison floats correctly. In particular,

```
Float 16 bit 2's complement mantissa, 4 bit exponent normalized float
generated from 32 bit internal fixed format standard instrument storage
format. Transmitted as follows: in groups of 5 words
```

```
LSbyte of first sample mantissa
MSbyte of first sample mantissa
LSbyte of second sample mantissa
MSbyte of second sample mantissa
LSbyte of third sample mantissa
MSbyte of third sample mantissa
LSbyte of fourth sample mantissa
MSbyte of fourth sample mantissa
```

Exponent word:

```
Least significant nibble of first byte for fourth word.
Most significant nibble of first byte for third word.
Least significant nibble of second byte for second word.
Most significant nibble of first byte for first word.
```

One must Check to see if 4 bit exponent is greater than 7  
(implies negative exponent in that case). Decode to negative  
4 bit number using assumption of 2's complement

If one has a choice, I recommend integer formats for recording and exchange. For exchange, I have used SEG-Y

**3.0.0.2 Processing Formats** There are a number of these. Seismic Unix (SU) has their format which is derived from SEG-Y. It is good for reflection seismic processing and is focused on vertical component geophones when it comes to headers. BSU uses what it calls BSEG-Y. It is very similar to SU, but headers include polarizations in 3-component directions for both sources and geophones. This is because the problems addressed by BSU include 3-components for both the source and geophone.

Both SU and BSEG-Y are similar to the exchange format, SEG-Y in that they use a 240 byte trace header followed by a data block. This way, the headers are locked to the data. Neither SU or BSEG-Y uses a reel header, since that goes back to the days of 9 inch tape.

BSU also supports dumping data to text and comma separated variable formats for those who just want to get the binary format out into something they can use in software like Matlab or Spread Sheets.

### 3.1 Conversion Utilities

The following programs do data format conversions:

- **BA2S 3.1.1** Converts an ASCII text file to BSEGY format.
- **BCNV 3.1.2** Converts between SEGY <-> BSEGY
- **BIS2SEG 3.1.3** Converts from Bison to BSEGY
- **BSWP 3.1.4** Byte swap. BSEGY <-> SUXDR
- **SEG2DUMP 3.1.5** Raw text dump of SEG-2 file
- **EGG2SEG 3.1.6** Converts between SEG-2 <-> BSEGY
- **GENB2S 3.1.7** Generates bash script, Bison to BSEGY
- **SEG2TXT 3.1.8** Converts BSEGY to ASCII text
- **SEG2CSV 3.1.9** Converts BSEGY to Comma Separated Variable

#### 3.1.1 BA2S

```
ba2s infile iorder ncol nrow dt
```

```

infile  = input file name
iorder  = 1 each row is a trace, columns are time axis (int)
         = 0 each col is a trace, rows are time axis (int)
ncol    = number of columns in matrix (int)
nrow    = number of rows in matrix (int)
dt      = sample interval in micro seconds (int)

```

#### 3.1.2 BCNV

```
bcnv infile endian compliance idirec idfc iunits hedfil
```

```

infile  = input file name
endian  0= Little Endian host (Linux PC)
        1= Big Endian host (IBM Mainframe)
compliance 1= SEGY Compliant (EBCDIC, IBM Float, BigEndian)
           0= (ASCII Reel Header, Float and endian of host)
idirec  0= BSEGY ==>SEG-Y (new *.sgy)
        1= SEG-Y ==>BSEGY (new *.seg)
idfc    1= floating point 4 byte output
        2= long integer 4 byte output
        3= short integer 2 byte output
         (uses reel header if SEG-Y input data)
iunits  1= meters
        2= feet
         (uses reel header if SEG-Y input data)

```

NOTE:

```

hedfil only input if idirec=0 BSEGY --> SEG-Y
(hedfil contains 3200 bytes, 40 records, 80char each
If hedfil='none', then blank lines after C used

```

### 3.1.3 BIS2SEG

No geometry setting of headers, just converts Bison to BSEGY.

```
bis2seg infile

    infile = input file name
```

### 3.1.4 BSWP

BSEGY and SU (before XDR) share many headers. However, if SU is compiled with XDR, the data and headers become byte swapped. SU with XDR is referred as SUXDR.

```
bswp infile idfc npts

    infile = input file name (4char minimum)

Data format
idfc  1=4 byte floating point
      2=4 byte integer
      3=2 byte integer
npts  =number of samples per trace
```

### 3.1.5 SEG2DUMP

Raw dump to a text file of SEG-2 data formatted data. Engineering seismographs like those available from [Geometrics](#) or that designed by the author [SeisRecorder](#).

The text file contains the data samples without applying any scaling due to fold or stack or instrument scaling factors. The input file \*.DAT is output as a \*.lst file name. File descriptor block and trace descriptor blocks for each trace are listed before the samples.

For example, if the data were integer recorded, the raw values as integers are displayed followed by the sample times.

This program is handy for debugging and evaluating recording parameters. If one wants a text file with all scale factors applied, consider converting the data to BSEGY format (program `egg2seg`) and apply either `SEG2TXT` or `SEG2CSV` programs which can produce text or comma separated value files.

For an input file 0000.DAT, output file 0000.lst sample follows:

```
ALL DATA VALUES ARE RAW SEG-2, No Scale Factors
```

```
Confirmed SEG-2 Data: 0X3A55
```

```
-----
FILE DESCRIPTOR BLOCK
-----
```

```
Size of trace pointer block = 32
Number of traces = 1
first string terminator character = 0
second string terminator character = 0
first line terminator character = 0XA
second line terminator character = 0
trace pointer (0) = 0XB8
number of characters=33
ACQUISITION_DATE 31/May/2019
Date (dayofyear.year: 151.2019
```

```

number of characters=30
ACQUISITION_TIME 13:45:52
TIME: 13 45 52

```

---

#### TRACE DESCRIPTOR BLOCKS

---

```

idfc = 2
npts = 500
CHANNEL_NUMBER 0
ALIAS_FILTER 100 6
HIGH_CUT_FILTER 100 6
FIXED_GAIN 19
DELAY 0.00
LINE_ID 0001
RAW_RECORD 0000.DAT
RECEIVER VERTICAL_GEOPHONE
SOURCE HAMMER
SOURCE_STATION_NUMBER 0000
RECEIVER_LOCATION 1.00 0.00 0.00
SOURCE_LOCATION 0.00 0.00 0.00
SAMPLE_INTERVAL 0.001000
DESCALING_FACTOR 1.1220185E-04
STACK 1

```

---

NOTE: All Values are Raw, No Scale Factors

[Indx]	Raw Value	(time)
[0]	-10521	(0.00000)
[1]	-10216	(0.00100)
[2]	-8080	(0.00200)
[3]	2373	(0.00300)
[4]	18318	(0.00400)
[5]	37544	(0.00500)
[6]	56084	(0.00600)
[7]	66536	(0.00700)
[8]	60356	(0.00800)
[9]	32890	(0.00900)
[10]	-11589	(0.01000)
[11]	-58586	(0.01100)
[12]	-88341	(0.01200)
[13]	-85594	(0.01300)
.	.	.
.	.	.
.	.	.

### 3.1.6 EGG2SEG

This is sample conversion of SEG-2 data to BSEGY It is just a simple conversion without any attempt to correct headers. To correct headers, see GENWAW 10.1.1. Or if survey data (\*.nez) are available, consider TOPCON2

**TOPCON2 Alternative** This alternative is to use a survey \*.nez file and topcon2

For example:

```
topcon2 a10001.nez 1061.dat 00A1 0.0 1 6 0186 0181 1061 0. 270 135 0 270
```

**BHED Fix Headers** This is an alternative to create a partial header file that can be edited

For Example:

```
bhed 1061.seg 1061.hed 1
```

Edit 1061.hed file for correct elevations, x, y, and z. Also, any ~@ characters like in PHONE= here need to be replaced with ascii characters.

Then, run

```
bhed 1061.seg 1061.hed 0
```

to upload the new headers producing a file bhed1061.seg

### 3.1.7 GENB2S

Generates bash script to convert BISON to BSEGY This is used typically to QC data, and minimal headers are created (geometry is zeroed out). It is an interactive program.

Example run, type in a terminal:

```
genb2s

enter 5_LETTER ALPHA PREFIX
LOGNO
enter 3digit FIRST FILE number
001
enter 3digit LAST FILE number
007
OUTPUT =====> gob2s
```

Make gob2s executable

```
chmod +x gob2s
```

Then run gob2s

Then run BMRG to perhaps look at the first trace from each file (vertical component down hole here)

```
bmrgr r00 001 007 1 1 1
```

### 3.1.8 SEG2TXT

This program converts a BSEGY seismic file to an ascii text file.

```
seg2txt infil tmin tmax fstrc lstrc timelist
```

```
infil =input file name
tmin  =minimum time
tmax  =maximum time
fstrc =first trace
lstrc =last trace
timelist = 0 do NOT add column of sample times
          = 1 Do add first column of sample times
```

One can convert only a portion of the BSEGY file, selecting a time and spatial window. If desired, an extra column of sample time values can be output as the first column.

### 3.1.9 SEG2CSV

Program SEG2CSV converts an entire BSEGY seismic file to a comma separated variable (csv) file with an additional column of sample times. Each column is a seismic trace.

```
seg2csv infile
```

```
infile = input file name (4char minimum)
```

## 4 Header Information

There are several codes that dump information about file contents. These are:

- **BDUMP 4.0.1** Shows partial headers of a BSEGY file
- **SEG2DUMP 3.1.5** Raw dump of SEG-2 acquisition files
- **BHELP 5.0.1** Lists all the programs in Basic Seismic Utilities
- **man pages 5.0.2** An information system in Linux or Unix

### 4.0.1 BDUMP

Most user interest is in a single shot gather, so the program focuses largely on receiver headers plus a single header for the shot. In that case, one may suppress shot header display after the first. Sometimes, as in reciprocal refraction shooting, the interest is in a geophone gather. In that case, one should display all shot headers. The command is issued from a terminal:

```
bdump infile iskip
```

```
infile:    =name of input file
iskip:     =0 suppress shot header display after 1st
           =1 display all shot headers
```

The output file is bdump.lst which may be viewed in any editor or cat to the screen. A sample of the output:

```

-----
|-----|
| PARTIAL SEG2 HEADER DUMP |
|-----|
|          k007.seg        |
|-----|
-----
Length = 2000 samples          | Shot Elevation = 998.4
Sample Interval = 0.00025 sec. | Shot Depth = 0.0
Delay Time = 0 msec.          | Up Hole Time = 0 msec
Low Cut Filter = 8 Hz.        | Shot X-COORD = 9927.00
High Cut Filter = 500 Hz.     | Shot Y-COORD = 9773.13
Line ID: BNK2                 | Shot Date (year.moday) = 1995.0628
Shot Orientation:             | Shot Time (hr:min) = 12:29
Azimuth= 0 Deg. Vertical=180 Deg. | Charge Size (grams)= 0
-----
TRACE|SHOT| STATION | OFFSET| RECEIVER          | VERT|1STBRK|K-GAIN|AZI|VER|
# |REC.|SHOT REC| | ELEV. X-COORD Y-COORD|FOLD|(SEC.)| (dB) | | |
-----|-----|-----|-----|-----|-----|-----|-----|-----|-----|
1 | 7| 024 001| 74.01| 1024.97 9981.25 9815.88|20|0.0580| 40 | 0| 0|
2 | 7| 024 002| 73.46| 1024.38 9980.50 9816.25|20|0.0552| 40 | 0| 0|
3 | 7| 024 003| 72.95| 1023.97 9979.76 9816.54|20|0.0540| 40 | 0| 0|
4 | 7| 024 004| 72.31| 1023.50 9978.88 9816.79|20|0.0531| 40 | 0| 0|
5 | 7| 024 005| 71.75| 1022.84 9978.09 9817.17|20|0.0557| 40 | 0| 0|
-----

```

Each trace row is a geophone associated with the shot.



## 4.0.2 SEG2DUMP

This is described above in section 3.1.5, since it is a complete text conversion of a SEG-2 exchange format program.

# 5 Software Documentation

## 5.0.1 BHELP

This code can be used to see a brief description of all the BSU codes streamed to a terminal screen. You may wish to pipe it through less or more programs. For example:

```
bhelp | more
or
bhelp | less
```

Here is a partial output of using “bhelp | less”.

```
babs.c    rectify seismic traces
bagc.c    automatic gain control of traces (scale in time and space)
ba2s.c    FORMAT CONVERSION: ASCII TEXT ---> BSEGY (no geometry setting)
bamp.F90  amplitude analysis by frequency (K-V Solid)Downhole Sph. Div.
bamx.F90  amplitude analysis by frequency (K-V Solid)SurfaceWaves Cyl.Div
bbal.c    balances two data sets to have same MAV (mean absolute value)
bcad.f    plot seismic traces as CAD (*.dxf; digital exchange file)
bcar.c    apply moving average (box car) filter as function of time
bcnv.c    FORMAT CONVERSION: BSEGY <---> SEG-Y (BSU=*.seg, SEG-Y=*.sgy)
bcrd.f    coordinate rotation and translation, BSEGY geometry headers
bdat.c    datuming program for refraction data (easier for picking)
bdcn.f    deconvolution (profile or trace mode), prediction or error out
bdif.f    differentiates w.r.t. time using Bilinear Transform method
bdum.f    generate dummy data set with user defined impulse position
bdump.f   generate a dump of selected BSEGY header values
bedt.f    edit BSEGY seismic file (traces, time, sample interval, etc.)
bequ.c    trace equalize data by L2 norm or Maximum Absolute Value
bext.c    extract traces from a merged data set based on header values
bfil.f    ARMA FILTER of seismic traces (low-, band-, or high-pass)
bfit.f    Solves for interval velocity from times in headers (VSP)
bftr.f    FILTER traces with other *.seg traces, or namelist from bdump
bfxt.f    F-X Transform of seismic traces
bgar.c    exponential GAIN recovery, by range specification
bgaz.c    exponential GAIN recovery, by depth specification
bhed.f    up/down load selected header information from/to a text file
bhelp.c   this listing of BSU package contents
bhod.F90  hodogram by PCA to determine down-hole tool orientation
bint.f    numerical integration of seismic traces (trapezoidal rule)
bis2seg.c FORMAT CONVERSION: BISON ---> BSEGY (no geometry setting)
bkil.f    either kill (delete) or zero seismic traces
bmed.f    median mix of seismic traces (spatial)
bmix.f    mean mix of seismic traces (spatial)
bmrq.f    merge traces from many files to a single file
bmrk.f    mark first break picks with a delta function on waveform
bmst.f    MASTER illustrates programing in BSU, FORTRAN
bnez.c    GEOMETRY: Create survey *.nez (Northing, Easting, Elevation) file
bnfd.c    MODELING: computes near and far field in elastic whole space
bnos.f    MODELING: generate band-limited random noise traces
```

## 5.0.2 man pages

From a terminal, type: man program name.

For example,  
man babs

```
babs(1)                Basic Seismic Utilities BSU                babs(1)
```

### NAME

babs - BSU program rectifies seismic traces (C-Language Version)

### SYNOPSIS

```
babs [ -h | infile ]
```

### DESCRIPTION

Basic Seismic Utilities (BSU) rectifies seismic traces by taking the absolute value. Functionally equivalent to the master program, cmst.c, this version is cleaned up and reflects the fact that there is only one command line argument necessary. C-Language Version.

### Options

-h Online help giving details on command line arguments

infile Input file name

### NOTE:

If invoked with no options, will prompt user for input parameters.

### EXAMPLE:

```
babs w001.seg
```

File w001.seg is processed by babs. Output traces are rectified.

### FILES

babsxxxx.seg

named according to convention (first 4char babs, the next 4char are the first 4char of the input file name, suffix .seg)

standard output

produces a progress bar

babsxxxx.lst

Echo check of input parameters in listing file.

### SEE ALSO

bhelp(1), cmst(1)

## 6 Plotting

BSU uses a number of ways to plot seismic data. Depending on how BSU is compiled, it can employ PLPLOT libraries, GNUPLOT libraries, OCTAVE, and old style line printer inspired text plots.

- **TRAPLT 6.0.1** Line printer inspired trace and spectrum (ASCII text)
- **BPLT 6.0.2** Plot seismic data with choice of output formats (PLPLOT or GNUPLOT depending on how compiled)
- **TPLT 6.0.3** Plot seismic trace (GNUPLOT)
- **QPLT 6.0.4** Plot seismic traces scaled by max amplitude (GNUPLOT)
- **CAPLOT 6.0.5** down-hole dispersion and amplitude decay (PLPLOT or GNUPLOT)
- **Octave TRAPLT 6.0.6** Octave version, trace and FFT spectrum
- **Octave YULE WALKER 6.0.7** Octave ALL POLE spectrum
- **Octave SEISAZI 6.0.8** Octave plot azimuth of down-hole horizontal components
- **Octave REFPLLOT 6.0.13** Octave first break analysis
- **Octave PROFPLOT 6.0.11** Octave plot of a shot profile
- **Octave HODOPLOT 6.0.9** Octave hodogram plot of two channels in same file
- **Octave HODO2PLOT 6.0.10** Octave hodogram plot of two channels in different files

### 6.0.1 TRAPLT

Inspired by old school line printer plots, the code produces a text file, traplt.lst, that can be viewed with only a terminal.

```
traplt infile tmin tmax trace# tzero ilin

infile: =name of input file
tmin:   =start time in seconds
tmax:   =end time in seconds
trace#: =trace number to list and plot
tzero:=zero time for phase reference; spectral plot
ilin:  spectrum plot 1=linear 0=dB
```

Here are some portions of the listing for an example case. The “j” column is the sample number, x(j), column is the sample amplitude in microvolts.

```
max= 0.1924409E+06 min=-0.1765663E+06
j      x(j) .....
51  0.4216680E+05 |          .***** |
52  0.4289184E+05 |          .***** |
53  0.3646188E+05 |          .***** |
54  0.2400264E+05 |          .**** |
55  0.7193160E+04 |          .* |
56 -0.1169604E+05 |          *. |
57 -0.2911608E+05 |          ****. |
58 -0.4088844E+05 |          *****. |
59 -0.4132728E+05 |          *****. |
60 -0.2852460E+05 |          ***. |
61 -0.3510720E+04 |          * |
```

```

62  0.2860092E+05 | .****
63  0.6088428E+05 | .*****
64  0.8744362E+05 | .*****
65  0.1032228E+06 | .*****
66  0.1058558E+06 | .*****
67  0.9560988E+05 | .*****
68  0.7389681E+05 | .*****
69  0.4399848E+05 | .*****
70  0.9291957E+04 | .**
71 -0.2699820E+05 | ***.
72 -0.6130404E+05 | *****.
73 -0.9099250E+05 | *****.
74 -0.1149951E+06 | *****.
75 -0.1331593E+06 | *****.
76 -0.1470877E+06 | *****.
77 -0.1580969E+06 | *****.
78 -0.1668928E+06 | *****.
79 -0.1731701E+06 | *****.
80 -0.1765663E+06 | *****.
81 -0.1765091E+06 | *****.
82 -0.1723306E+06 | *****.
83 -0.1633629E+06 | *****.
84 -0.1505412E+06 | *****.
85 -0.1346094E+06 | *****.
86 -0.1165979E+06 | *****.
87 -0.9709806E+05 | *****.
88 -0.7677788E+05 | *****.
89 -0.5510304E+05 | *****.
90 -0.3155832E+05 | ****.
91 -0.6506277E+04 | *
    
```

A portion of the spectrum listing:

```

maxa=0.636E+07 mina=0.231E+03 maxp=0.180E+03 minp=-.180E+03 tzero= 0.000
j  freq  amp  phz  ....linear scale.....-180.....0.....+180
1  .00000  0.3 -180.0 |* |* |. |
2   3.9   0.4  40.6 |* | |. * |
3   7.8   0.6 -93.3 |** | |* |. |
4  11.7   1.1 158.3 |*** | |. * |
5  15.6   2.0  70.8 |***** | |. * |
6  19.5   3.3  -6.6 |***** | |* |. |
7  23.4   5.2 -81.0 |***** | |* |. |
8  27.3   7.2 -155.5 |***** | |* |. |
9  31.2   9.0 129.5 |***** | |. * |
10 35.2  10.0  54.8 |***** | |. * |
11 39.1   9.8 -18.3 |***** | |* |. |
12 43.0   8.7 -88.4 |***** | |* |. |
13 46.9   7.3 -154.4 |***** | |* |. |
14 50.8   6.0 143.6 |***** | |. * |
15 54.7   5.1  83.5 |***** | |. * |
16 58.6   4.3  23.8 |***** | |* |. |
17 62.5   3.5 -33.7 |***** | |* |. |
18 66.4   3.0 -87.6 |***** | |* |. |
19 70.3   2.6 -139.2 |***** | |* |. |
20 74.2   2.3 169.6 |***** | |. * |
    
```

```

21  78.1   2.0 120.4 |*****          |           .      *      |
22  82.0   1.8  74.4 |****          |           .      *      |

```

## 6.0.2 BPLT

Depending on how conditionally compiled, the program will either use PLPLOT or GNUPLOT libraries. The command line arguments are:

```
bplt infile idev iorient itype ltr Ltr tmin tmax istyl amp percent xaxis yaxis
```

```

infile  =  input file name
idev    =  output device
          0=  xwin/wxt (Linux/MS Windows)
          1=  Post Script
          2=  xfig
          3=  jpeg
          4=  PDF
iorient =  orientation
          0=  landscape
          1=  portrait
itype   =  select non-time axis type
          0=  trace number
          1=  offset
          2=  geophone z-coord
          3=  geophone x-coord
          4=  geophone y-coord
          5=  shot z-coord
          6=  shot x-coord
          7=  shot y-coord
ltr     =  first trace to plot
Ltr     =  last trace to plot
tmin    =  minimum time to plot
tmax    =  maximum time to plot
istyl   =  style of plot
          0=  wiggle plot
          1=  black/white variable area
          2=  black/grey variable area
amp     =  amplitude for 1 trace deflection
percent =  percent overplot 100= 1 trace
xaxis   =  length of x-axis (non-time) in inches
yaxis   =  length of y-axis (time) in inches
          (if xaxis and yaxis absent, 6.0 by 4.0 inches

```

For example, `bplt c008.seg 4 0 1 1 100 0 0.6 1 2.E+4 200 7.0 3.0`

**6.0.2.1 Trace Equalization** Program BPLT is a true amplitude plot, the x-axis label indicates the amplitude of a single trace deflection. At times, there is a need to see detail in both low and large amplitude portions of a shot gather. This can be done by running program BEQU on the data, and then plotting that. For example,

`bequ c008.seg 0. 0.5` equalizes each trace amplitude using the L2 norm. Then run BPLT on file `bequc008.seg`. For example,

```
bplt bequc008.seg 4 0 1 1 100 0 0.6 1 2.0 200 7.0 3.0
```

**6.0.2.2 xplot bash script** As an example of how to set up a script to do the above sequence with the additional flexibility of scaling by one of the following normalizations:

- Peak Absolute Value of profile
- L2 Norm of profile

- Trace by trace L2 norm (this matches the above example in Figure 2)

```
#!/bin/bash
OR=0 #orientation 0=landscape 1=portrait
if test "$1" = "-h"
then
echo "USAGE: xplot filename tmax scaling"
echo 'Scaling Choices:'
echo '1= Peak Absolute Value of profile'
echo '2= L2 Norm of profile'
echo '3= Trace by trace L2 Norm'
else
if test "$1" = ''
then
echo 'Enter input file name'
read FILEN
else
FILEN=$1
fi

if test "$2" = ''
then
echo 'Enter tmax'
read TMAX
else
TMAX=$2
fi

if test "$3" = ''
then
echo 'Enter Scaling Choice'
echo '1= Peak Absolute Value of profile'
echo '2= L2 Norm of profile'
echo '3= Trace by trace L2 Norm'
read SCL
else
SCL=$3
fi

NAME='basename $FILEN .seg'
NAME4='echo $NAME | gawk -F " " '{print $1$2$3$4}'' '
```

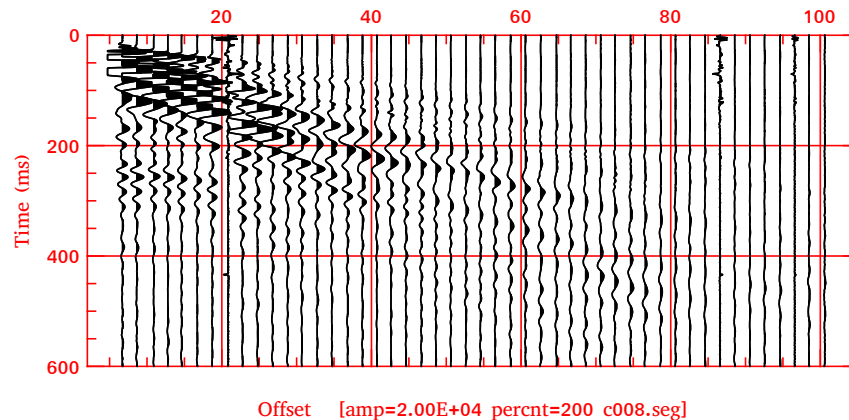


Figure 1: Example of a trace by offset in meters plot, written to file bplt.pdf

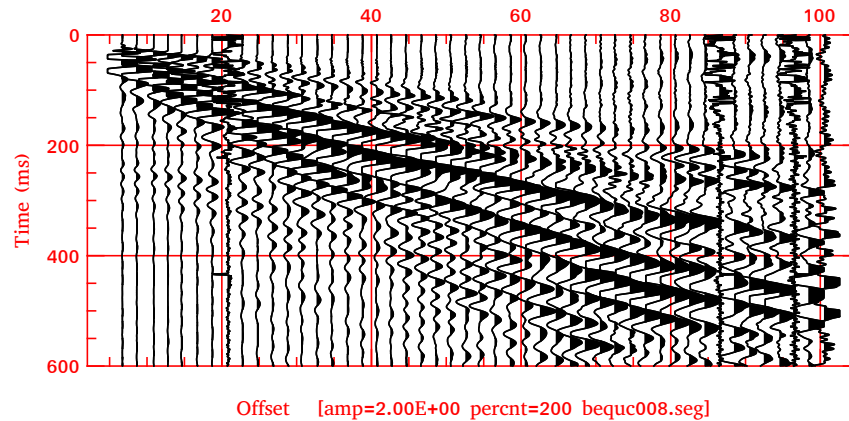


Figure 2: Trace equalized version of Figure 1

```

case $SCL in
1)
bscl $FILEN 1 5000 3
AMP='gawk '/Peak Absolute Value/ {print $4}' bscl$NAME4.lst'
rm -f bscl$NAME4.*
PFILEN=$FILEN
bplt $PFILEN 0 $OR 0 1 500 0 $TMAX 1 $AMP 200
;;
2)
bscl $FILEN 1 5000 1
AMP='gawk '/L2 Norm of Data Set=/ {print $6}' bscl$NAME4.lst'
rm -f bscl$NAME4.*
PFILEN=$FILEN
bplt $PFILEN 0 $OR 0 1 500 0 $TMAX 1 $AMP 200
;;
3)
bequ $FILEN 0 $TMAX
PFILEN=bequ$NAME4.seg
AMP=4
bplt $PFILEN 0 $OR 0 1 5000 0 $TMAX 1 $AMP 200
rm -f bequ$NAME4.*
;;
esac
rm -f bplt*.lst
fi

```

### 6.0.3 TPLT

This program plots a single trace to an X11 screen. The command is

```
tplt infile trace_number tmin tmax
```

```

infile = input file name (4char minimum)
trace_number = trace number to plot
tmin = minimum time in seconds
tmax = maximum time in seconds

```

It also outputs a GNUPLOT file, graph.gp which can be edited or not, and run as a bash script. For example, to create a PDF file of the plot, comment out (insert a # symbol at the beginning of the line) the “set terminal” command and replace it as follows:

```

#set terminal x11 persist
set terminal pdf

```

```
set output "graph.pdf"
```

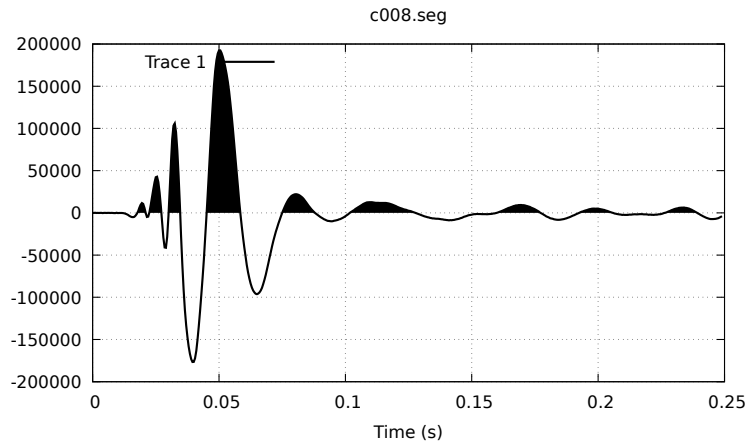


Figure 3: TPLT: Plot of the first trace in the file c008.seg. Units are microvolts if only instrument corrections have been applied. Of course that will change depending on the processing history.

#### 6.0.4 QPLT

A quick quality control plot in which each trace is scaled by its maximum value and the displayed by GNUPLOT to the X11 window. One can modify the display interactively. Pressing enter in the terminal will freeze the plot. Also output is a file, qgraph.gp, which can be edited for an alternative terminal. The program can be run with the following command line arguments:

```
qplt infile tmin tmax
```

```
infile = input file name (4char minimum)
tmin = minimum time in seconds
tmax = maximum time in seconds
```

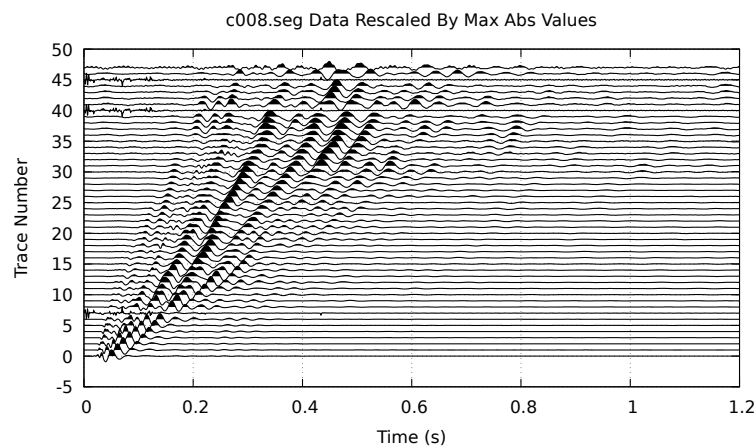


Figure 4: QPLT: Quality control plot showing traces, each scaled by the maximum value in the trace. Output includes X11 and qgraph.gp (GNUPLOT)



### 6.0.5 CAPLOT

Down-hole surveys can determine stiffness and damping of soils in shear. Two BSU programs, are used to measure S-wave velocity dispersion (BVAS) and amplitude decay with distance traveled (BAMP). These programs produce two files, bvas.his and bamp.his which can then be used in a joint inversion scheme to determine stiffness and damping. Program CAPLOT may be used to create an image file displaying the dispersion and decay measurements with 95% confidence bars.

```
caplot bvas_file bamp_file emin emax well year date idev
```

```
bvas_file =input file ( bvas.his)
```

```
bamp_file =input file ( bamp.his)
```

Title info from bvas and bamp runs:

```
emin      =minimum elevation (real)
emax      =maximum elevation (real)
well      =well name (char 4)
year      = year of survey (integer, 4 digits)
date      = 4 digit integer mmdd
idev      =device for plotting
0=X window display
1=Post Script *.ps file
2=Xfig *.fig file
3=JPEG *.jpeg file
4=PDF *.pdf file
```

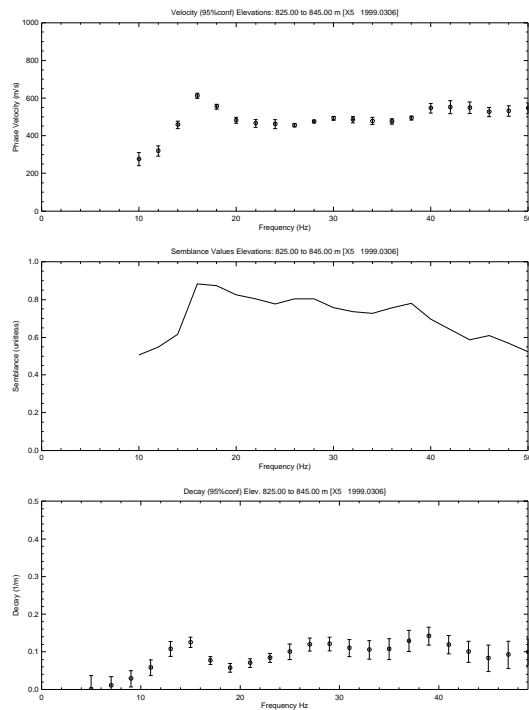


Figure 5: CAPLOT: Display S-wave dispersion and amplitude decay from program outputs of BVAS and BAMP programs.

### 6.0.6 OCTAVE TRAPLT

This is the octave version of TRAPLT. The following files are required to be in the directory where octave is started.

- **bsegin.m** Reads traces from BSEGY files
- **segyinfo.m** Reads header information from BSEGY files
- **traplt.m** Actually does the plotting

Start an octave session and then type  
traplt

You will be prompted for a file name, channel  
phase reference, and maximum frequency to  
display.

First shows a plot of the selected trace, mouse  
used to pick time zero for phase. Click OK then use mouse.

Uses FFT for amplitude spectrum

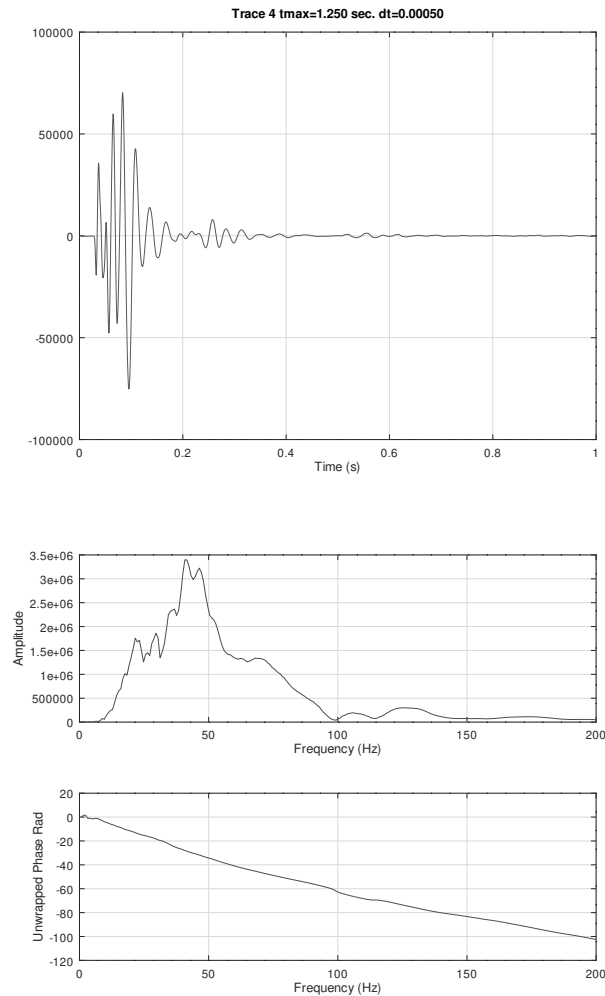


Figure 6: OCTAVE TRAPLT: Octave version of TRAPLT program. Octave is a mathematical interactive program like Matlab. Compare this plot to the all pole yule walker spectrum of figure 7

### 6.0.7 OCTAVE YULEWALKER

This program computes the ALL POLE spectrum for a signal. In addition to the octave plots, a GNUPLOT (plotspec.gp) file is output along with a data file yw.dat which is readable by the plotspec.gp file. The code is interactive and one uses the mouse to pick the order of the process on the autocorrelation.

The yulewalker.m program is run in octave.  
 You can run it on BSEGY, data or  
 you can run it on the autocorrelation of  
 BSEGY data.

In either case, the spectrum will be an  
 all pole spectrum. You will be prompted  
 to use your mouse to select the maximum  
 lag in the autocorrelation.

If you run BXCR followed by BSTK, you can  
 compute an average autocorrelation for the  
 entire shot gather, and run yulewalker.m on  
 that.

Other \*.m files required (bsegin.m and segyinfo.m)  
 must be in the same directory before yulewalker.m  
 is run in an octave session.

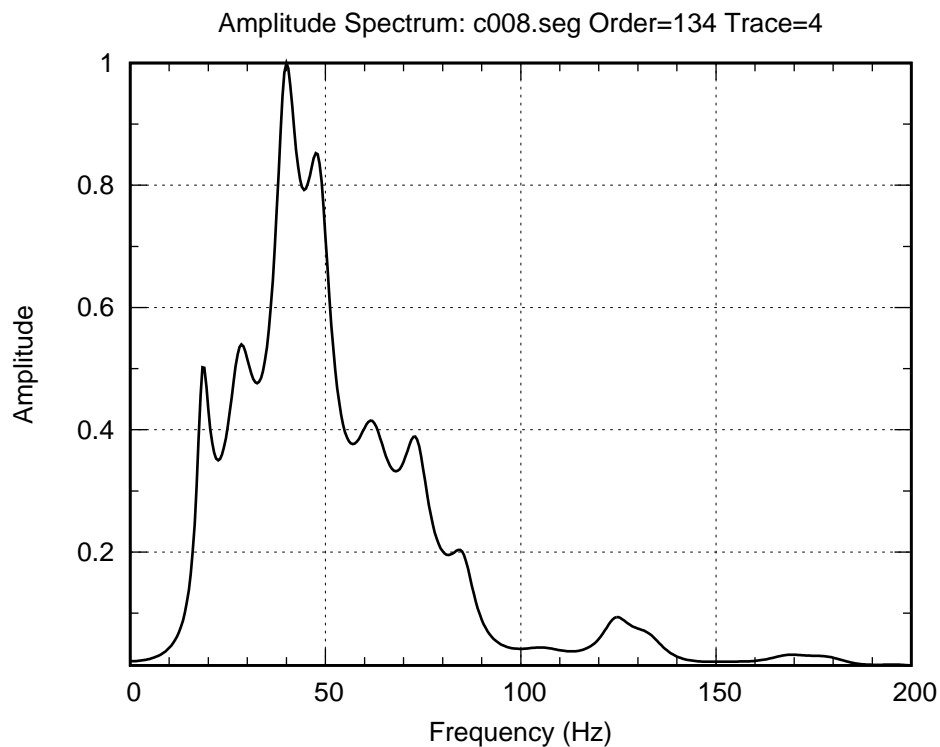


Figure 7: OCTAVE YULE WALKER: Octave program which computes the ALL POLE spectrum. Input can be either a seismic trace data or an autocorrelation of trace data (either must be in BSEGY format, \*.seg file). Compare to Figure 6 FFT plot. See BXCR 12.0.16 for how to create an autocorrelation as input.

### 6.0.8 OCTAVE SEISAZI

This program is run in an octave session. Start octave and type `seisazi`. The program will request an `*.seg` file name. It will then display a GUI showing the number of traces and the sample interval. Click on OK. The program will generate a plot of the horizontal component geophone. In a typical application, one extracts a single horizontal component from a collection of multi-component files using the BMRG program. The information is extracted from the headers.

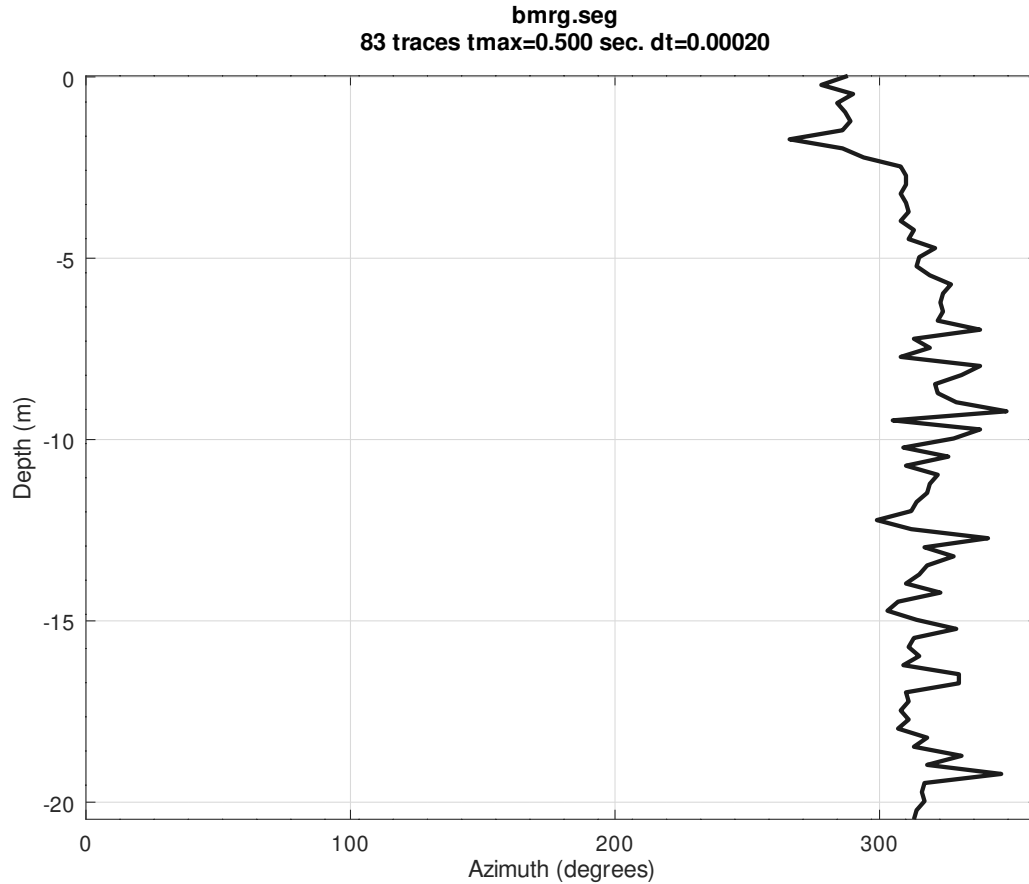


Figure 8: OCTAVE SEISAZI: Plots a horizontal component azimuth from the headers of an `*.seg` file. Here, the plot is of the T-component from a down-hole survey. Phone orientation was determined using program BHOD.

### 6.0.9 OCTAVE HODOPLOT

This program is run in an octave session. Start an octave session and type `hodoplot`. This program is for the case of two components in the same `*.seg` file.

Prompts follow:

1. enter file name. Example: `1001.seg` hardwired as a 6 channel shot record with 3 components down-hole and 3 components stationary reference phone at the surface.

Down-hole Phone

ch1=Vert

ch2=Radial

ch3=Transverse

Reference Phone

ch4=V

ch5=R

ch6=T

2. GUI, informative.
3. GUI, choose either the down-hole or surface reference phone.
4. GUI, choose component for X-axis, say T
5. GUI, choose component for Y-axis, say R
6. GUI, choose a scale factor, or default
7. GUI, choose a Tmax, say .05 seconds
8. GUI, click continue
9. GUI, choose the next Tmax, say 0.10 seconds
10. GUI, click continue

**KEY POINTS:**

\* There will always be a sign convention. Here, on the vertical phone, upward velocity produces a negative voltage. Do a tap test for your equipment.

\* The hodoplot.m program is hard wired to relate components to the GUI choices. If your phones use different channels for V, R, T, then you may need to modify the code. By the way, R and T are just arbitrary labels considering that down-hole phones will twist their orientation as they travel up or down the hole.

What you are doing is progressively working down in time, plotting the particle motion. With each step, just change the Tmax value, the Tmin value is automatically adjusted to the last step tmax value. Here the first arrival largest motion is in the direction of the R component. This means that it is mostly aligned with the source blow to the West. The T component is oriented mostly orthogonal to the source.

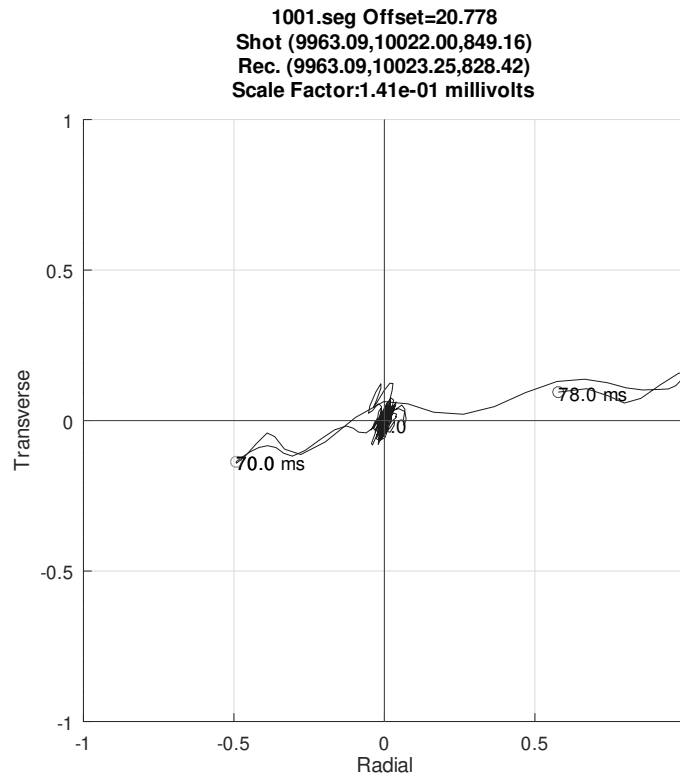


Figure 9: OCTAVE HODOPLOT: Plotting particle motion on the down-hole horizontal R- and T- components which are channels in the same \*.seg file. If components are in different files, use HODO2PLOT program instead (see 6.0.10).

### 6.0.10 OCTAVE HODO2PLOT

This is an octave program that reads two \*.seg files for plotting a hodogram. The files are checked to insure that the sample intervals and other parameters match.

Start an octave session and type  
hodo2plot

This hodo2plot.m file is different from hodoplot.m in that the components to be plotted are from two different \*.seg files, rather than a single \*.seg file. An example would be a Rayleigh wave problem where the vertical and radial components reside in different files.

1. prompt for which file is X-axis, which is Y-axis.

The code will check that the same number of traces and time sampling are used for both files.

2. GUI, choose a trace to plot.

3. GUI, choose scale factor or just default.

4. GUI, choose a Tmax value, say .05 seconds.

5. GUI, click continue

6. GUI, choose a new Tmax value, say .10 seconds

7. GUI, click continue

What you are doing is plotting the particle motion for increasing time steps.

KEY POINT:

\* There will always be a sign convention. Upward velocity produces a negative voltage on vertical phone. Do a tap test for your equipment.

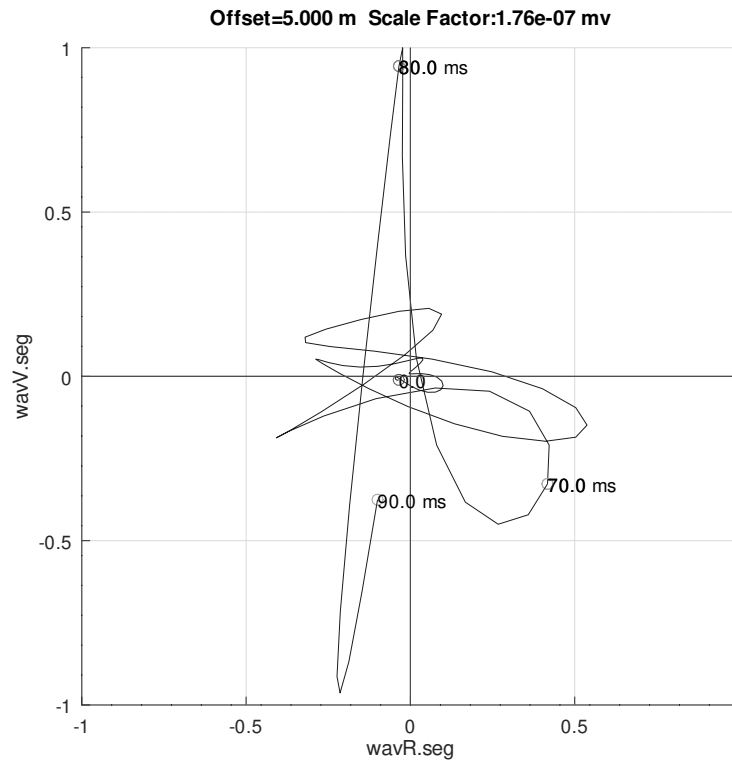


Figure 10: OCTAVE HODO2PLOT: Plotting particle motion on the Radial and Vertical components of a Rayleigh wave problem in which the channels reside in different \*.seg files. If components are in a single file, use HODOPLOT program instead (see 6.0.9).

### 6.0.11 OCTAVE PROFPLOT

This program just does a simple trace plot of a shot gather in octave. Depending on your octave installation, it is likely that you can zoom in for detail on the plot. This program also serves as an example of how to read a BSEGY \*.seg file. Requires segyinfo.m and bsegin.m in the same directory.

```
Start an octave session then type
profplot
```

```
The program will prompt for a file name. Type a full file name like:
c008.seg
for example.
```

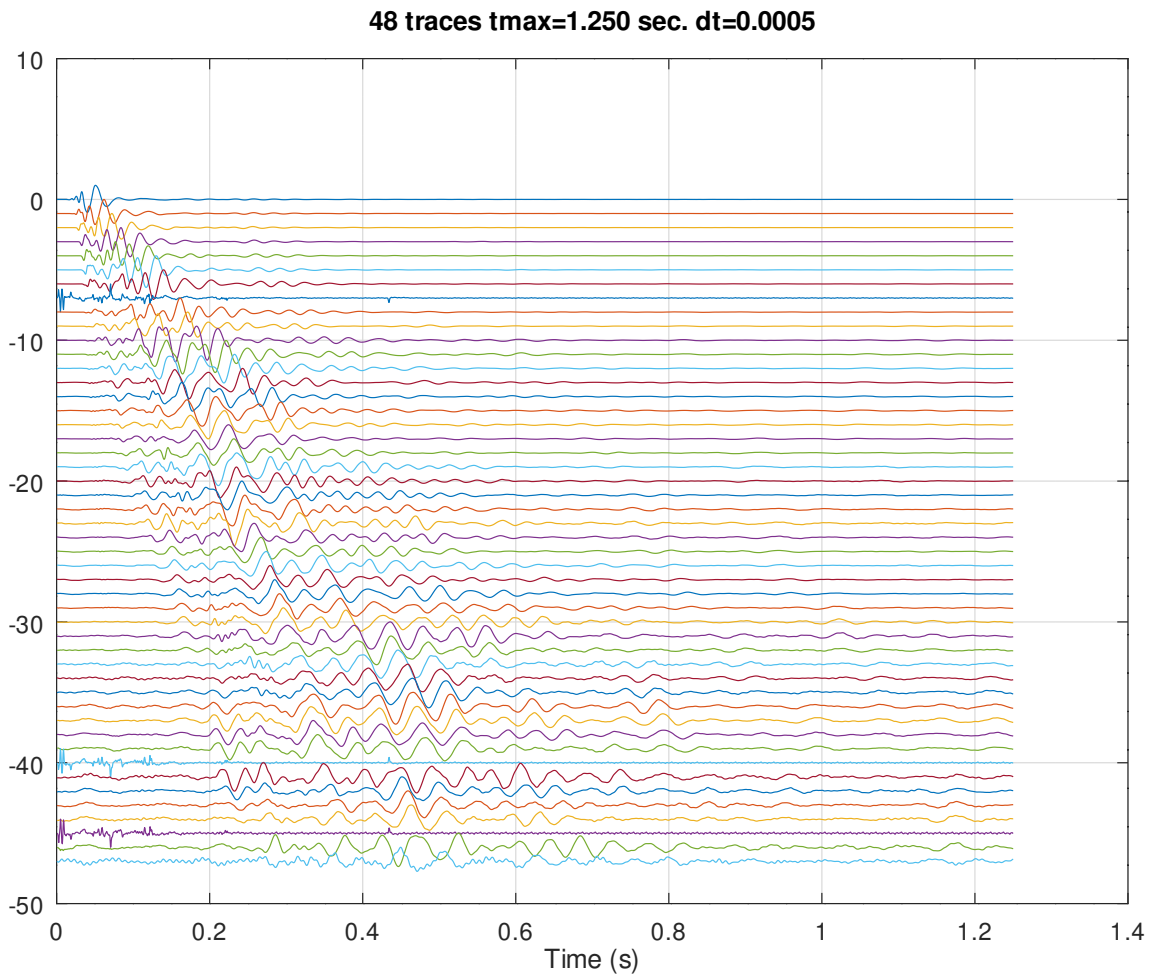


Figure 11: OCTAVE PROFPLOT: Plots a shot gather of traces in BSEGY formatted file, \*.seg. Traces are individually scaled by the maximum value. Compare to images Figure 2 and Figure 4.

### 6.0.12 OCTAVE SEGPIC

The program `segpic.m` is run in octave to plot each trace and permit picking with a mouse. The output is a file ending in `*.pic`. Program `BPIC` can be used to insert the pics into the `*.seg` file headers. See section 8.4 for an example of using `segpic.m` in conjunction with datuming program `BREF`. Also see `BPIC 8.4.3`.

Start an octave session, then type

```
segpic
```

1. prompt for file name, like `k007.seg` for example.
2. GUI shows number of traces and `tmax`.
3. GUI prompt for a clip factor and reduced `tmax` for good plotting resolution. Suggest clip factor of 3 and maybe `.1` for `tmax`, depending on when arrivals come in.
4. GUI page through each trace, using mouse to pic first arrival, typically a down motion with SEGY polarity conventions.

Pics are output to an ascii `*.pic` file. Trace (number, pick time):

```
1 0.05695853
2 0.05534562
3 0.04336406
4 0.05281106
5 0.05396313
for example.
```

Use program `BPIC` to insert pics to headers, for example:

```
bpic k007.seg 1 k007.pic 0.
```

The above command would be executed from an terminal, after octave session is ended.

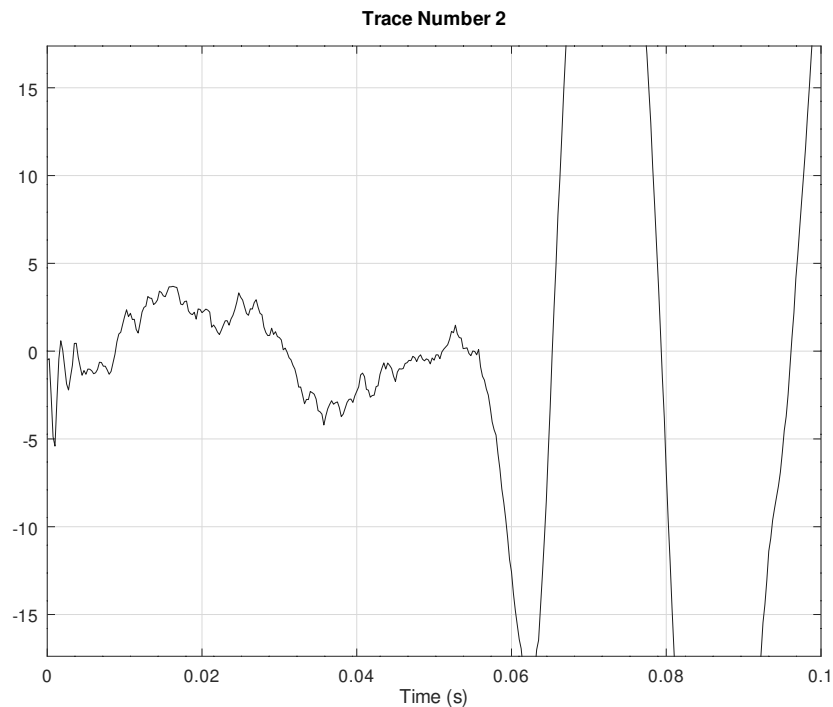


Figure 12: OCTAVE SEGPIC: Example of a trace for picking with mouse. First arrival refraction is at about 0.055 seconds.



### 6.0.13 OCTAVE REF PLOT

This program can be used to both plot and measure apparent velocities of refraction arrival time picks. It is up to the user to know how to determine which arrivals are refractions. The first arrival picks must have been done first and inserted into the \*.seg file headers (see programs segpic.m, BDAT, BPIC, PICRESTORE).

Start an octave session, then type  
refplot

One is prompted for the file name. For example:  
k008.seg

Choose either stations or offsets  
Pick a segment to get started,  
Click yes  
Then 3 mouse clicks, click a near offset  
then a far offset limit, a line will be  
fit (OLS), a third click will print the  
estimated velocity and 95% confidence  
values where you click on the plot.  
Chi<sup>2</sup> info GUI, then choose to do another  
segment or not.

When picking the near and far offset  
limits, only the horizontal, x-axis  
position of the mouse matters.

A Postscript output file, plot.ps, is created and  
can be viewed with ghostscript.

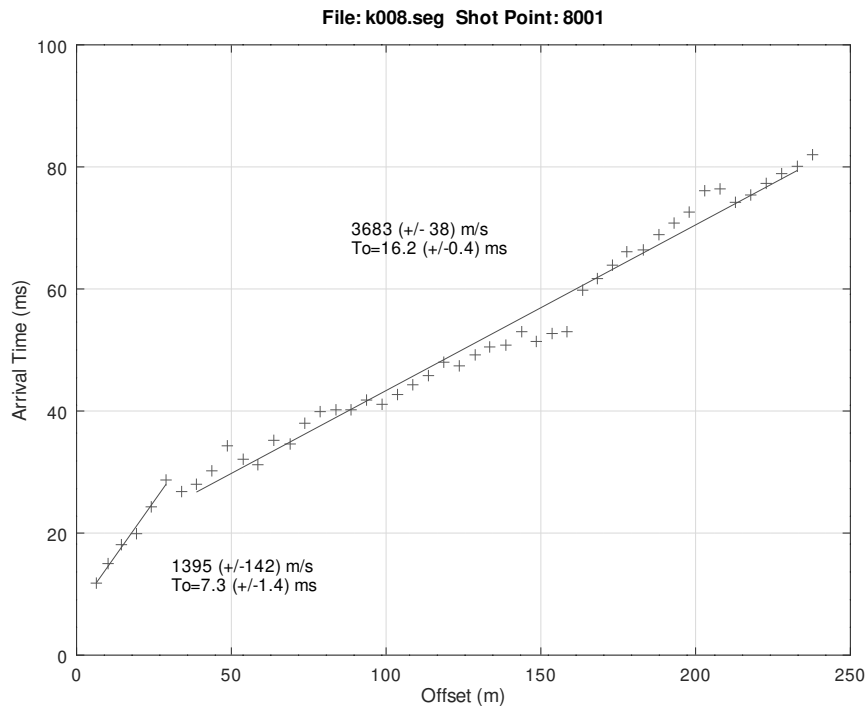


Figure 13: OCTAVE REF PLOT: Plots first break picks which have been added to headers with BPIC. Then use mouse to pick line segment (start,end), followed by a mouse click to plot refractor apparent velocity result. See section 8.4.6, estimating a cross-over distance for program BREF.

## 7 Surface Seismic

### 7.0.1 BRED

Correctional velocity can be applied to a data set to static shift data into a linear alignment (direct waves or refracted head waves). Alternatively, one can apply hyperbolic (NMO, reflection) correction to the data. Flattening the data on a refracted arrival can make picking first breaks easier in some cases (see section 8.4)

```
bred infile iopt vred t0 tshift
```

```
infile: =name of input file
iopt:   =statics option
        1=linear trend
        0=hyperbolic NMO trend
vred:   =reduction velocity
t0:     =zero offset time (iopt=0 ONLY)
tshift: =bulk static shift
```

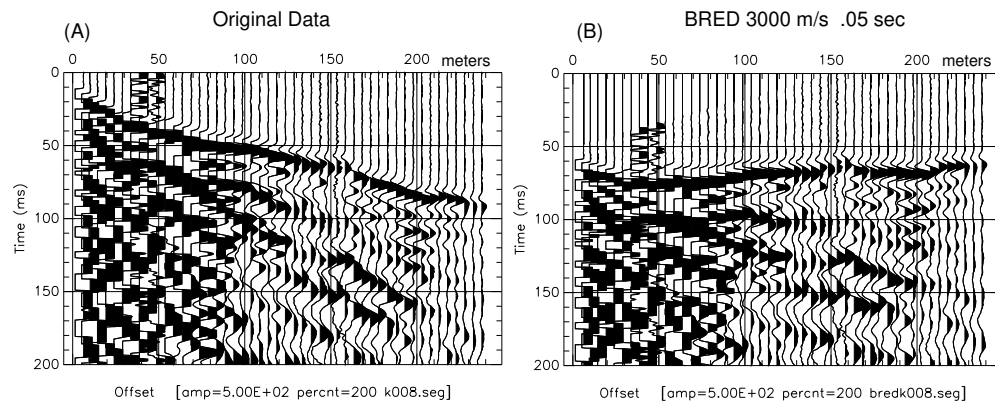


Figure 14: (A) Plot of a shot gather, (B) BRED: linear trend, 3000 m/s reduction velocity, .05 seconds offset. See section 8.4.0.1 for an example of picking data with BRED.

### 7.0.2 BVAX

Run BVAX for surface wave dispersion measurement. A number of image files are created, and the file `bvax.his` is available for use in the inversion program `invR1.m` (run in octave). To run `invR1.m` in octave, execute `build_disper_oct` script to build an extension to octave. Edit the `bvax.his` file to remove any measurements that are zero or bogus velocities. NOTE: BVAX determines *PHASE* velocities in the time domain.

```
bvax infile xmin xmax vmin vmax nvel . . .
      fmin fmax delf bwd iskp ivscn
```

```
infile =input file name
xmin   =minimum offset (float)
xmax   =maximum offset (float)
vmin   =minimum velocity
vmax   =maximum velocity
nvel   =number of velocity increments
fmin   =minimum frequency Hz
fmax   =maximum frequency Hz
delf   =frequency increment Hz
bwd    =filter bandwidth Hz
iskp   =skip filtering (if files already exist)
        1=YES 0=NO (-1=NO and delete when done)
```

```
ivscn  =output velocity scan data sets  
        1=YES  0=NO
```

```
EXAMPLE: bvax c008.seg 1.0 100. 100. 500. 200 10. 50. 1. 1. -1 0
```

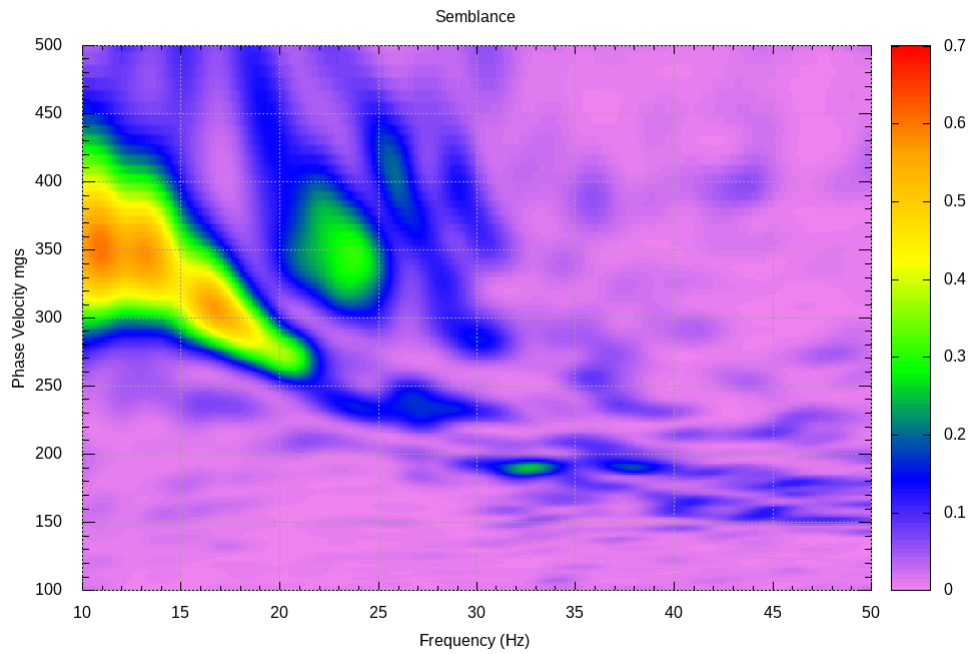


Figure 15: BVAX: Phase velocity semblance display file, clrplot.png. For details on semblance, see [Sheriff \(1991\)](#). Semblance provides a measure of the degree to which the data were aligned at a trial velocity.

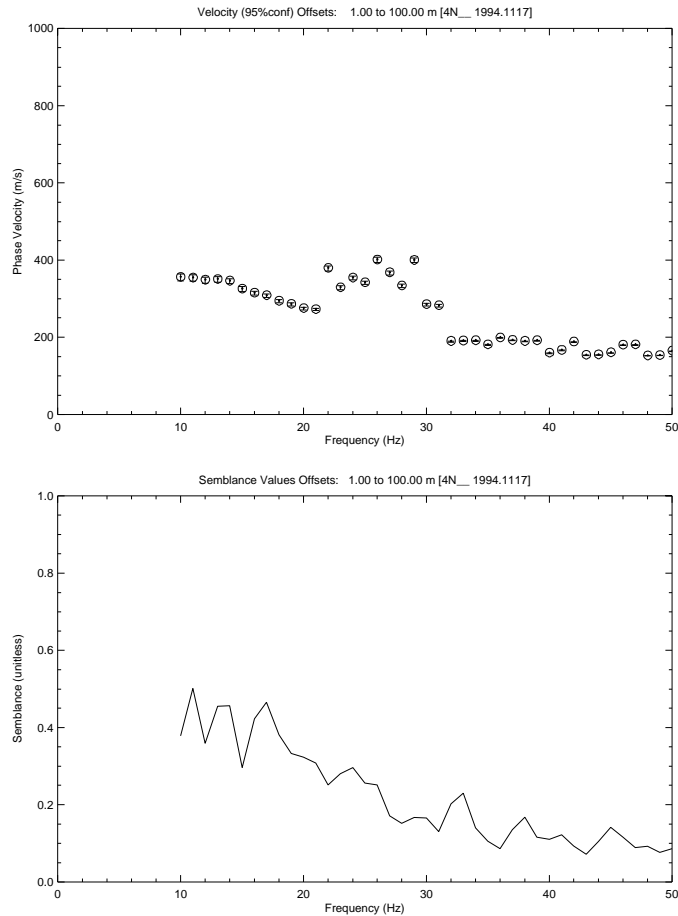


Figure 16: BVAX: Phase velocity semblance display file, bvax.ps

There are some useful data at the frequencies above and below the 20-30 Hz range. The file, bvax.his should be edited to remove questionable data with a lot of scatter (perhaps due to higher modes), or in some cases where zero velocity is returned due to a failure to find a phase velocity. That happens when the range of velocities scanned is too limited, or when there is no signal. Once edited, an inversion in Octave can be done with program invR1.m.

### 7.0.3 BAMX

Program BAMX computes amplitude decay with frequency. The code attempts to measure the viscoelastic alternative to an elastic earth. It is similar to the BAMP code which is used in down-hole measurements of viscoelasticity. BSU software does NOT have code to invert surface wave data under a viscoelastic representation, at this time. In this case, the decay is modest, but if large decay were present, one might wish to develop a viscoelastic surface wave inversion code.

```
bamx infile rmin rmax fmin fmax delf bw tmax
```

```
infile =input file name
rmin   =scan gate: (min_offset)
rmax   =scan gate: (max_offset)
fmin   =min band pass center frequency (Hz)
fmax   =max band pass center frequency (Hz)
delf   =frequency step (Hz)
bw     =bandwidth of filter (Hz)
tmax   =time gate: max_time
```

## 8 Inversion Codes

Forward problems take an earth representation and compute a corresponding geophysical expression of that representation. The inverse problem goes the other way and computes an earth representation from geophysical data. Basic Seismic Utilities (BSU) is focused on near surface problems. The typical representation of the earth is a soil profile with S- or P- wave velocities as a function of depth being the object of interest.

### 8.1 Surface Waves

The surface waves of interest in BSU software are Rayleigh waves. These are a mixture of SV- and P-wave motion that satisfy Hookes law  $F = k \cdot x$  and Newton's law of motion,  $F = m \cdot a$ . The particle motion is largely elliptical and can be measured on both vertical and in-line radial (horizontal) component geophones. BSU codes compute features from seismic data, specifically a dispersion curve. The soil profile representation is 1-D, varying only in depth. Inversion is done in Octave. The difference between **SASW** (section 8.1.2) and **saswv** (8.1.3) codes is in the type of file read. **SASW** read a BSEGY formatted file, **saswv** reads a text file of cross power spectrum. Program **invR1** reads a bvax.his file (see 8.1.1 and 7.0.2).

#### 8.1.1 OCTAVE invR1, Rayleigh Wave Inversion

This program uses the BVAX output file, bvax.his, to invert Rayleigh wave, fundamental mode, under an elastic representation. See section 7.0.2 for details on BVAX. A companion forward problem octave code is **FwdR1.m**, see section 9.2.1.

To run invR1.m in octave, first execute **build\_disper\_oct** script to build an extension to octave. Edit the bvax.his file to remove any measurements that are zero or bogus velocities. Program invR1.m is hard wired to read bvax.his, so that should be the name of any edited file that will be used by invR1.m. The octave files are located at the /usr/local/share/octave/site-m/ directory.

The code is an iterative inversion which runs for a user number of inversion steps. Default is 2, but recommend 5 as a useful number. Increasing the number of singular values employed will provide additional detail in the inversion result. However, if you use too many, noise in the data may inject details in the result that are not reliable. Or the code can become unstable if too many singular values are used.

1). Enter initial soil representation file:

File model.txt is used to set an initial model of control points

For example, with 3 control points:

```
3
200 300 500
.0 2.0 15.
```

Velocity	Depth
200 m/s	0 m
300 m/s	2.0 m
500 m/s	15.0 m

2). GUI Choose P-wave velocity option. Click on Vp/Vs ratio OR Vp=fixed (if fixed, GUI enter Vp m/s and Density kg/m3

3). GUI Choose density parameters, Poisson Ratio, grain density, porosity, degree water saturation.

4). GUI Informs user of Vp/Vs ratio and constant density to be used.

NOTE: Code will seek a S-wave velocity profile consistent with these results.

5). GUI Select number of singular values to use, layer thickness (constant in meters), locksw (switch to lock some conditions), and number of iterations to do. Typically, 5 or more are good, but for the first run, 2 is wise in case things go sideways.

locksw	meaning
0	free bottom control, velocity and depth
1	(default) lock bottom depth, free bottom velocity
2	lock bottom velocity, free bottom depth
3	lock both bottom velocity and depth

The bottom is the deepest control point, top of the bounding half-space.

Program FwdR1.m is a manual forward program that can be used to do the inversion manually, or to explore the fast and slow limits based on confidence limits.

The edited bvax.his file from the example in section 7.0.2 was run using a constant VpVs ratio for just 5 iterations. Density was held constant. Other settings from a GUI are Poisson Ratio=0.33, grain density 2.67 g/cc, porosity 0.33, degree water saturation 1.00. This results in VpVs=1.99 and density=2169 kg/m<sup>3</sup>. Only 3 singular values included, layer thickness, deltz=0.1 meters. The resulting velocity model

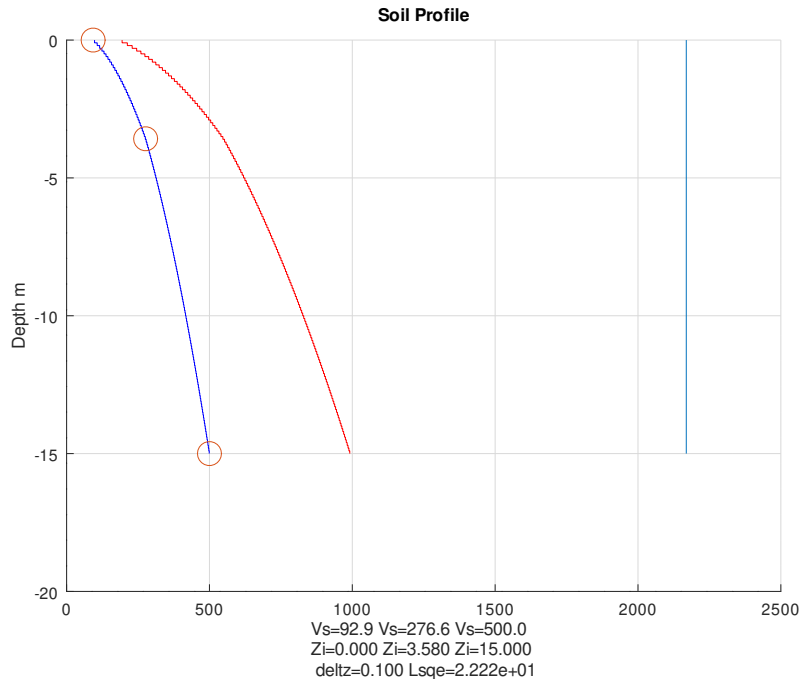


Figure 17: invR1: After 5 iterations, the resulting soil model is shown. The S-wave velocity with inverted control points is shown as the Blue curve (m/s). The Red curve is the P-wave velocity, and at the far right is the constant density (kg/m<sup>3</sup>)

Other output includes text files of the solution as well as a fast and slow 95% limit cases.

solution.txt (Three rows: number of control, S-velocities, Depths)

```
3.00000000e+00
9.28725027e+01 2.76581965e+02 5.00000000e+02
0.00000000e+03.00000000e+00
```

slow.txt

```
9.08153279e+01 2.75856068e+02 5.00000000e+02
0.00000000e+00 3.59321540e+00 1.50000000e+010 3.58000000e+00 1.50000000e+01
```

fast.txt

```
3.00000000e+00
9.49296774e+01 2.77307861e+02 5.00000000e+02
0.00000000e+00 3.56678460e+00 1.50000000e+01
```

**8.1.1.1 Solution Uncertainty** The above \*.txt files can be converted to alternative slow and fast plots of the soil profiles showing S-velocity with depth.

- One can uncomment the two plotvel() functions under the Plus and Minus Limits section of the invR1.m code. These are lines 620 and 630 of the current version of invR1.m Running the octave program, invR1.m, with the same parameters, but now with the fast and slow plotvel() calls will display the solution surrounded by the fast and slow solutions. In some cases, there will not be much difference, but using the zoom function on Figure 2 of the octave program output can be used to see the difference.

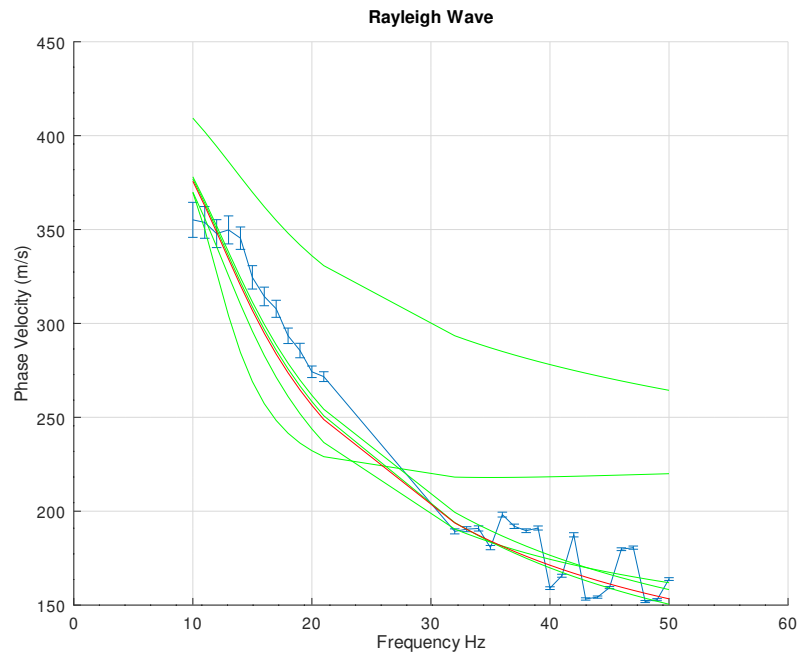


Figure 18: invR1: Progress of the inversion. The initial model dispersion is the fastest green curve. The green curve is the dispersion after 5 iterations. Data from bvax.his is in blue.

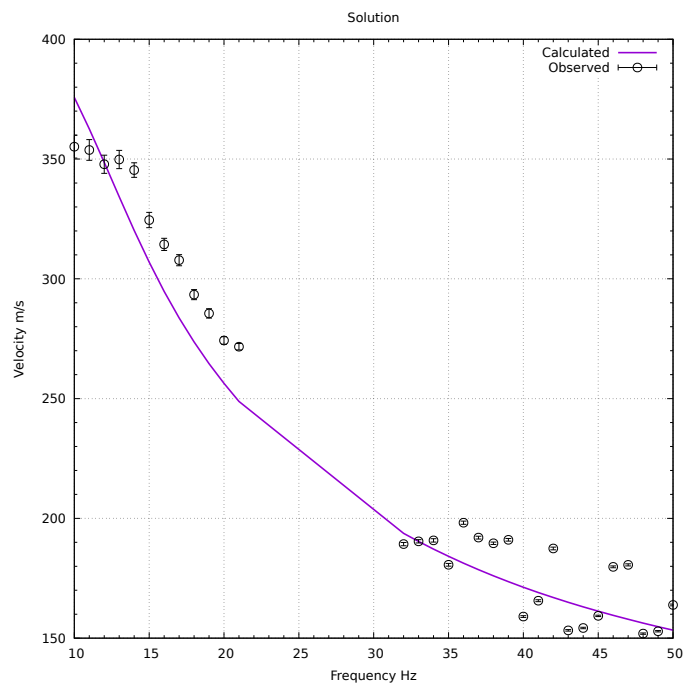


Figure 19: invR1: The code also generates a GNUPLOT file, dispcrv.gp, which shows the final solution when run with the gnuplot program.

- The other option is to use the octave program, FwdR1.m (see section 9.2.1 ), and when prompted for the model, enter either fast.txt or slow.txt (instead of model.txt) to compute these cases and their fit to the data. After the first plot, end the program with the GUI and it will generate additional plots showing both the fast or slow model and fit to the data.

### 8.1.2 OCTAVE SASW

In theory, only two traces are needed to compute a dispersion curve. Program SASW.m permits one to select two traces and compute Rayleigh-wave velocity dispersion. Depending on the trace spacing and spectral selection, the code recommends a maximum spacing between the two traces (to avoid aliasing).

Start an octave session and then type  
SASW

Note, capital letters are important since that agrees with the file SASW.m

This code takes two signals from a shot gather to compute a cross spectrum leading to a dispersion curve.

Prompts:

- 1). enter file name, example: c008.seg
- 2). GUI Pop up to select fmin, fmax vmin vmax
- 3), Info GUI pops up and shows both time and spatial sample intervals.

Recommended trace separation is indicated on last line. If high frequencies are chosen, then too large a separation between the two geophone stations can lead to aliasing.

- 4). GUI enter tmax, near trace number, far trace number.

For example:

tmax =1.0

trace R1 = 2

trace R2 = 3

(this would follow a recommendation that R2-R1 be no larger than 1)

When only two offsets are used, one should always look at the entire shot gather first and select traces likely to be dominated by the fundamental mode (typically close to the source).

The program produces two figures. One shows the cross power spectrum and coherence (Figure 20). The other figure shows the dispersion over a range of frequencies selected in the GUI prompt when the code is run (Figure 21).

The code determines a time shift,  $\Delta t$ , between geophones with a separation of  $\Delta x$  to compute a phase velocity at each frequency of interest. If  $\Phi$  is the unwrapped phase angle at a frequency of interest, then

$$\Delta t = \frac{\Phi}{2\pi} \cdot T, \quad (1)$$

where  $T$  is the period for a frequency  $f$  ( $T = 1/f$ ). The phase velocity at the frequency  $f$  is

$$C(f) = \frac{\Delta x}{\Delta t}. \quad (2)$$



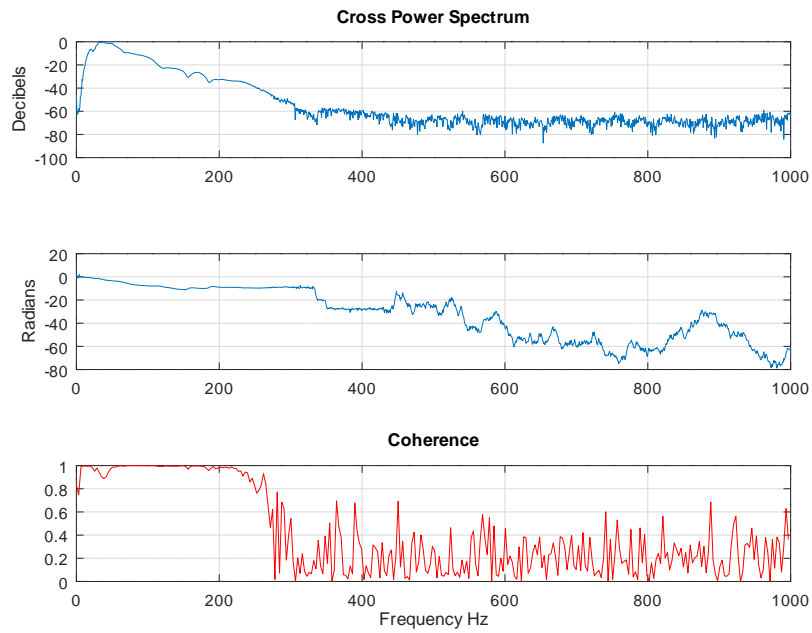


Figure 20: SASW: Cross spectrum amplitude and coherence reveal what range of frequencies provides useful dispersion information.

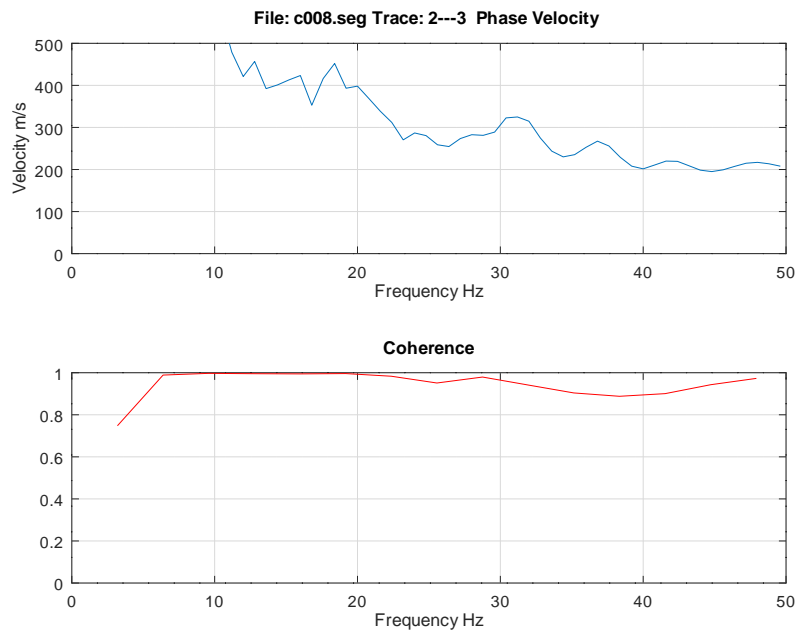


Figure 21: SASW: Dispersion computation over limited range of frequencies selected in the GUI.

### 8.1.3 OCTAVE saswv

This program was developed to read a text file with a measured cross power spectrum. It was used in a Benchmark Test sponsored by the Geo-Institute of the American Society of Civil Engineers (ASCE). The format of the text file is shown by the following first few lines of one instance:

```
dX = 32
R1= 45 R2 = 77 S = -7 T-Rex Shaker
  forward
f (Hz) |Gxy| (volts) Ph (Gxy) (deg) Coherence
5 0 -66.75 1
5.5 0.01 -74.25 1
6 0.01 -77.12 1
6.5 0.01 -79.08 1
7 0.01 -88.2 1
7.5 0.02 -103.03 1
8 0.02 -111.59 1
.
.
.
```

To run the program, start octave and type  
saswv

- 1). Enter the text file with the cross spectrum
- 2). Select a range of frequencies to plot

Details on the data set are found in [Michaels \(2014\)](#).

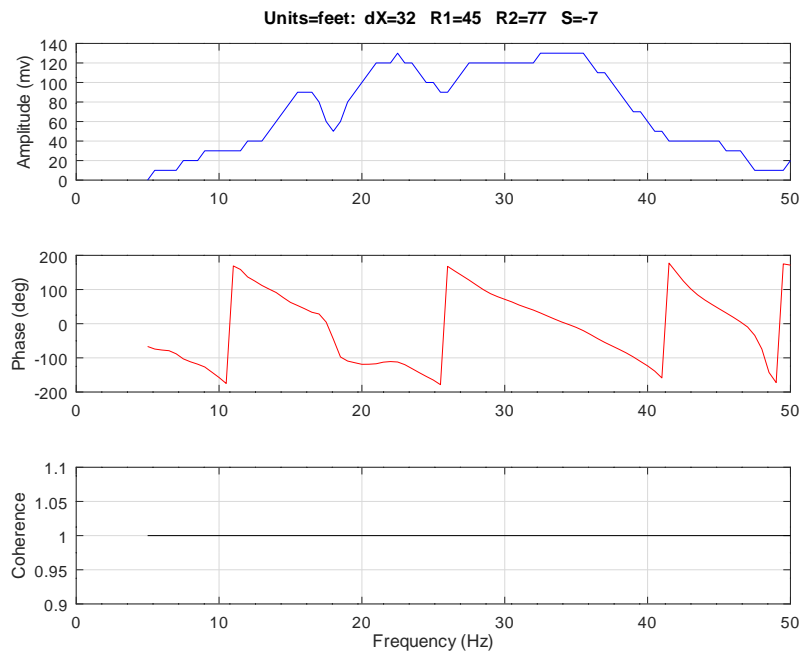


Figure 22: saswv: Cross power spectrum from data [Michaels \(2014\)](#).

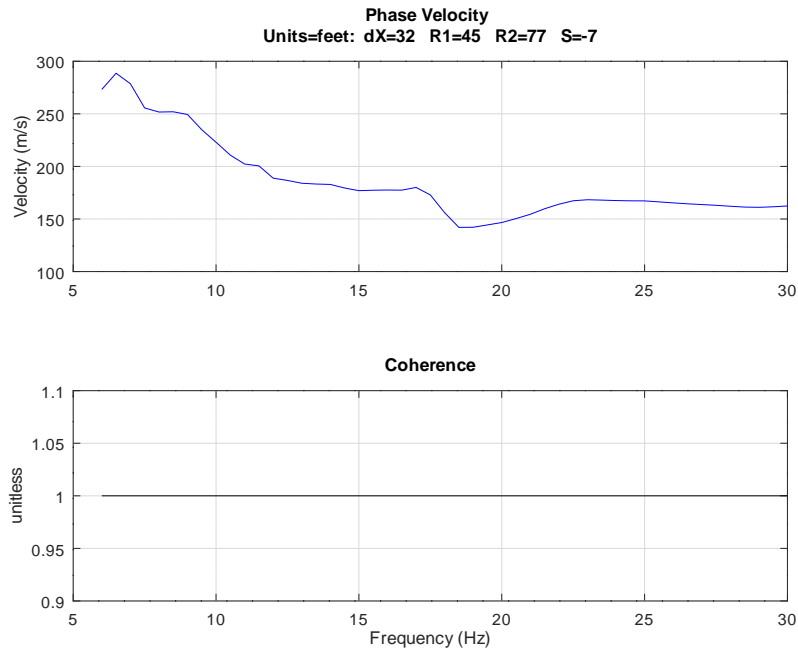


Figure 23: saswv: Dispersion computed from data [Michaels \(2014\)](#).

## 8.2 Down Hole Seismic

Down-hole processing codes include:

- **BFIT 8.2.1** Fit an interval velocity to vertical times
- **BVEL 8.2.2** applies correctional velocity by static shift
- **OCTAVE VFITW 8.2.3** fit interval velocities to vertical times of picked down-hole data.
- **OCTAVE VPLOTT 8.2.3.1** replot VFITW solutions, nice axes
- **BVSP 8.2.4** fits 3-layer model to picked down-hole data
- **BVAS 8.2.5** measures body wave dispersion SH-wave data
- **BAMP 8.2.6** measures body wave decay SH-wave data
- **OCTAVE CAINV3 8.2.7** inverts for stiffness and damping from **BVAS** and **BAMP** results.

### 8.2.1 BFIT

Vertical times correct for the source horizontal offset. If the vertical distance between the source and the geophone is  $Z$ , if the horizontal offset of the source from the bore hole is  $H$ , and if the straight line slant distance from source to geophone is  $S$ , then the cosine of the angle,  $\theta$ , between the vertical and the slant is  $\cos(\theta) = \frac{Z}{S}$ . The slant time is  $T_s = \frac{S}{V_i}$  where  $V_i$  is the interval velocity. Typically we don't measure  $S$ , but do measure  $H$ . So the angle,  $\theta = \arctan(\frac{H}{Z})$ . The vertical time is then:

$$T_v = T_s \cdot \cos(\theta) = \frac{Z}{V_i} = \frac{S}{V_i} \cdot \cos(\theta), \quad (3)$$

where  $T_s$  is the observed arrival time. Except in large horizontal offsets, the correction is modest. This program computes a straight line fit to vertical times,  $T_v$ . A similar program in OCTAVE is VFITW 8.2.3. The command line arguments are:

```
bfit infile emin emax labl

infile = input file name (4char minimum)

emin   =minimum elevation for interval
emax   =maximum elevation for interval
labl   =2 character ID label for interval
```

Example for the X5 borehole: `bfit twave.seg 820. 840. X5`

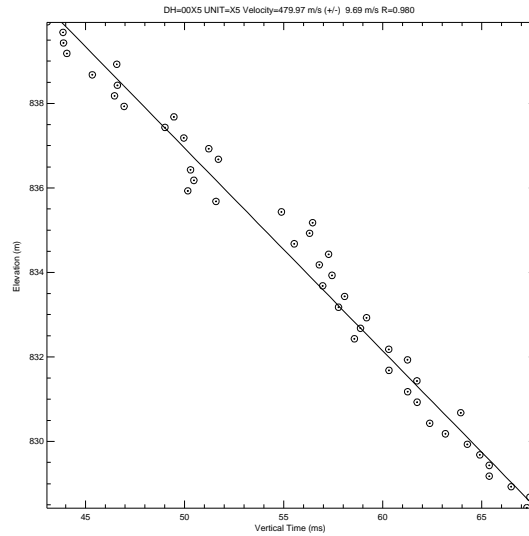


Figure 24: BFIT: Straight line fit yields interval velocity by least squares. Title has the value of the velocity, 479  $\pm$ 10 m/s.

### 8.2.2 BVEL

This program can apply a correctional velocity to a down-hole data set. See **BRED 7.0.1** for the same process on surface data. The command line arguments are:

```
bvel infile vel ifast
infile: =input file name
vel     =phase velocity to apply
ifast   0=slow option (fft phase rotation
        1=fast option (sample shift)
```

Example:  
bvel twave.seg 500. 1

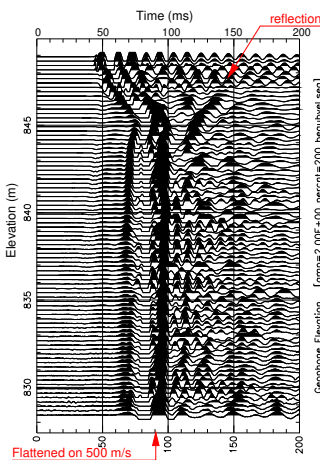


Figure 25: BVEL: Data flattened on 500 m/s (direct wave in bedrock). Overburden is slower (about 100 m/s). Reflection off top of bedrock shown.

### 8.2.3 OCTAVE VFITW

Given a BSEGY (\*.seg) down-hole data set with first break pics in the headers (see **SEGPIC 6.0.12**), this program computes vertical times from the observed slant times. The user uses the mouse to select a start and ending depth,

and the vertical velocity is computed. Then the mouse is used to place a label. After all the desired intervals are picked, the user may elect to replot the data using the **OCTAVE VPLOT 8.2.3.1** program.

Start an octave session

```
vfitw
enter file name
GUI: choose units      <feet | meters>
GUI: choose vertical axis <depth | elevation>
GUI: enter title
GUI: save to disk ?    <yes | no>   (recommend yes here)
GUI: do an interval?   <yes | no >
loop here, use mouse to click start and end depth
/|\      use mouse to place label with velocity (+/- m/s)
|        GUI: do another interval? <yes | no >
|--if yes if no---|
          \|\
          exit
```

### Output Files:

**filename.seg.vt** and **filename.seg.vt2** (required for **VPLOT**).

#### 8.2.3.1 OCTAVE VPLOT

Start octave session

```
vplot
inter file name
GUI: Axes Limits <xmin | xmax | ymin | ymax> (x=vertical time, y=depth/elevation)
save plot to postscript
```

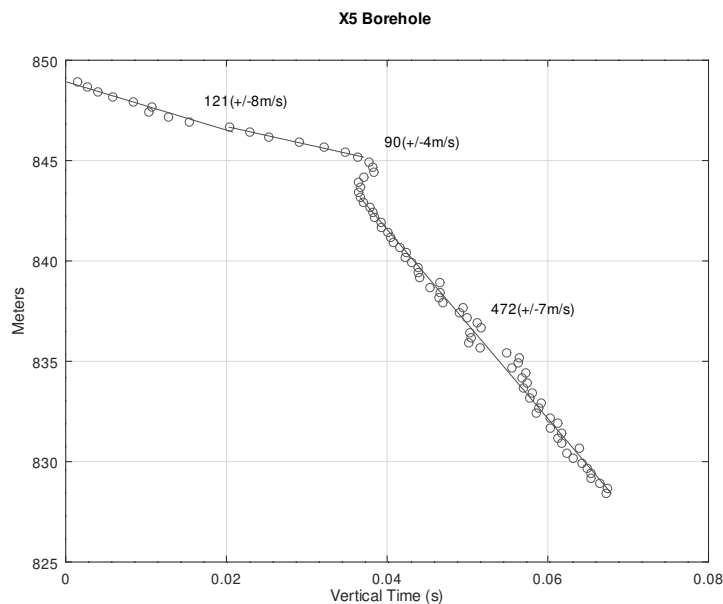


Figure 26: VFITW →VPLOT: Plot of vertical time vs. elevation, and interval velocities. Axes and placement of velocity labels by mouse.

### 8.2.4 BVSP

This code reads a down-hole survey in BSEGY seismic format. It fits a 3 layer velocity model to the arrival times. The output \*.lst file identifies the first arrivals as direct or refracted ray paths.

```

bvsp infile itmax zmin zmax
infile:  =input file name
itmax   =maximum number of iterations
zmin    =minimum depth to include
zmax    =maximum depth to include

```

EXAMPLE OUTPUT SAMPLE FROM \*.lst

Iteration	LSQE	V1	V2	V3	Z1	Z2
0	0.02201	107.5	206.7	305.8	4.8	3.8
1	0.01985	108.6	221.8	316.6	4.7	3.9
2	0.01791	109.7	238.5	327.0	4.7	4.0
3	0.01616	110.6	253.2	337.0	4.6	4.1
4	0.01458	111.3	269.3	346.4	4.5	4.2
5	0.01202	112.0	285.5	355.4	4.5	2.0
6	0.01087	112.6	314.0	364.2	4.5	2.0
7	0.00984	113.1	345.0	372.6	4.5	2.0
8	0.00891	113.5	379.0	380.5	4.5	2.0
9	0.00807	113.9	416.9	388.0	4.5	2.0
10	0.00732	114.3	459.7	395.1	4.5	2.0

The values Z1 and Z2 are layer thicknesses. So for iteration 10, the first layer is 4.5 meters thick. The second layer is 2 meters thick, placing the top of the half space at 6.5 meters depth.

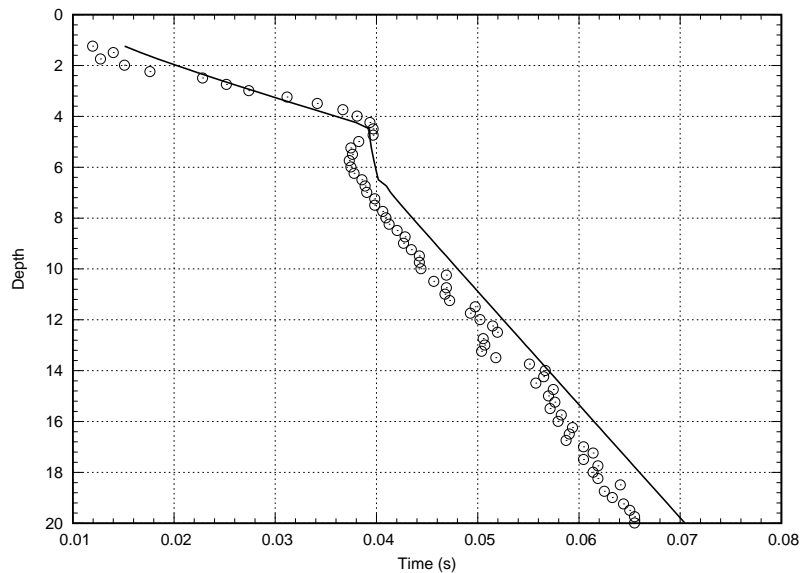


Figure 27: BVSP: Solution is a first layer, 4.5 meter thick  $V_s=114.3$  m/s, second layer 2.0 meter thick,  $V_s=459.7$  m/s, on top of a half-space with  $V_s=395.1$  m/s.

### 8.2.5 BVAS

Programs BVAS and BAMP 8.2.6 assume a viscoelastic medium. The results are inverted under a Kelvin-Voigt (KV) constitutive model with program **OCTAVE\_CAINV3** 8.2.7 Michaels (1998). Inversion under a Kelvin-Voigt-Maxwell-Biot (KVMB) model may be used to estimate hydraulic conductivity if porosity is available using **OCTAVE\_KD4kvmb** program (Michaels, 2006). SH wave enhanced down-hole data in BSEGY format (\*.seg files) are processed for body wave dispersion. The **delf**, frequency increment should be no smaller than the reciprocal of the record length. For a 0.5 second recording, 2 Hz is the finest resolution.

```
bvas infile emin emax vmin vmax nvel fmin fmax delf bwd iskp ivscn
```

```
infile =input file name
emin   =minimum receiver elevation (float)
emax   =maximum receiver elevation (float)
vmin   =minimum velocity
vmax   =maximum velocity
nvel   =number of velocity increments
fmin   =minimum frequency Hz
fmax   =maximum frequency Hz
delf   =frequency increment Hz
bwd    =filter bandwidth Hz
iskp   =skip filtering (if files already exist)
        1=YES 0=NO (-1=NO and delete when done)
ivscn  =output velocity scan data sets
        1=YES 0=NO
```

The output includes a file, **bvas.his** which can be processed by the inversion code, **cainv3** (section 8.2.7). The columns of the bvas.his file are also defined at the end of the \*.lst file which contains details of the run. For example:

Frequency	Phase Vel.	+/- m/s	Semblance	Tbar	Tvar
10.00	275.70	17.588079	0.5065	0.0342	0.0083
12.00	319.28	13.666108	0.5476	0.0144	0.0059
14.00	458.14	9.981112	0.6169	0.0053	0.0019
16.00	612.12	6.796203	0.8824	0.0016	0.0007
18.00	554.49	6.774652	0.8730	0.0022	0.0008
20.00	481.10	7.960438	0.8251	0.0028	0.0014
22.00	466.08	11.081243	0.8044	0.0026	0.0021
24.00	461.89	12.072690	0.7758	0.0018	0.0022
26.00	454.63	4.133945	0.8045	0.0012	0.0007
28.00	475.22	3.958418	0.8036	0.0011	0.0007

Other outputs include a Postscript plot, bvas.ps, a QC plot, bvasqc.ps and a number of semblance plots.

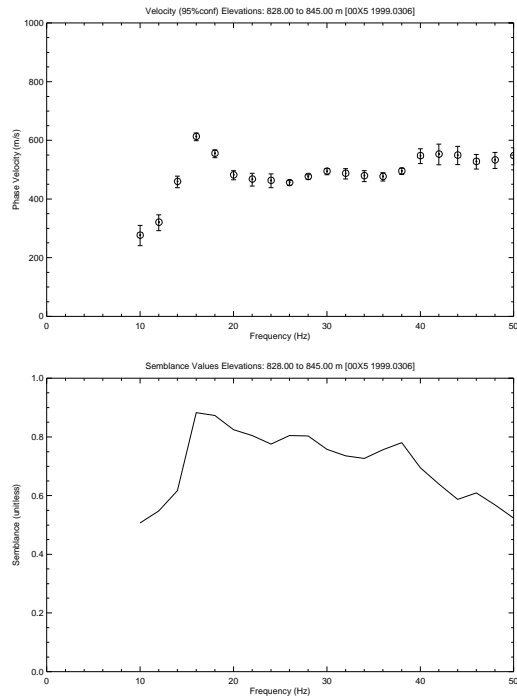


Figure 28: BVAS: SH body-wave dispersion and semblance results for down-hole data. These are the automated picks for maximum semblance as seen in Figure 29. Viscous, Kelvin-Voigt behavior is an increase in velocity with frequency (Michaels, 1998).

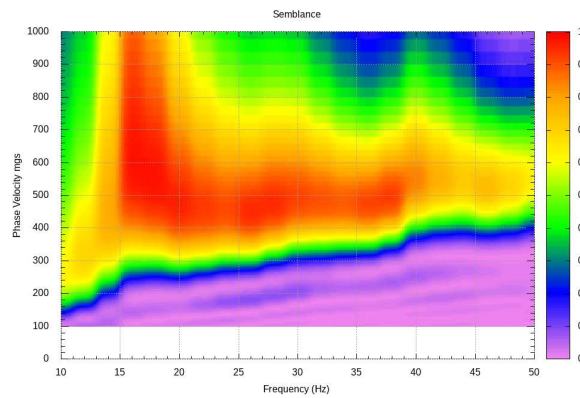


Figure 29: BVAS: SH body-wave semblance results for down-hole data.



### 8.2.6 BAMP

SH-wave enhanced down-hole data in BSEGY format (\*.seg files) are processed for amplitude decay with frequency. Beam spreading is corrected for assuming spherical divergence. The frequency increment, **delf**, should be no smaller than the reciprocal of the record length. For a 0.5 second recording, 2 Hz is the finest resolution.

```
bamp infile emin emax fmin fmax delf bw tmax
```

```
infile =input file name
emin   =scan gate: (min_elev)
emax   =scan gate: (max_elev)
fmin   =min band pass center frequency (Hz)
fmax   =max band pass center frequency (Hz)
delf   =frequency step (Hz)
bw     =bandwidth of filter (Hz)
tmax   =time gate: max_time
```

Output includes a file, **bamp.his** which can be used in procedure **cainv3** (see section 8.2.7). Also output are Postscript files, **bamp.ps** and **bampqc.ps**. The bamp.his file is 3 columns (frequency, decay 1/meters, standard deviation). For example:

5.00	0.0000	0.0186
7.00	0.0104	0.0121
9.00	0.0287	0.0109
11.00	0.0576	0.0105
13.00	0.1077	0.0099
15.00	0.1251	0.0070
17.00	0.0771	0.0055
19.00	0.0574	0.0059
21.00	0.0703	0.0059
23.00	0.0840	0.0062

### 8.2.7 OCTAVE CAINV3

Down-hole SH-wave data are inverted for stiffness and damping with this Octave program. Required are the \*.his file results from programs **BVAS** (section 8.2.5) and **BAMP** (section 8.2.6). Files **bvas.his** and **bamp.his** are required for each depth interval of interest. In order to compute uncertainty error bars, the depth interval should include as many subsurface stations as possible. Since **cainv3.m** is a joint inversion of body wave velocity dispersion and amplitude decay (corrected for beam divergence), the \*.his files do not need to include exactly the same subsurface stations, as when there is a need to remove poor data from one or both \*.his files. A companion program that computes the forward problem is **cafwd3.m** see section 9.1.1 .

The governing differential equation for this problem is a 3rd order PDE that is formulated as a 1-D plane wave problem (hence the need for the BAMP program to correct for beam divergence).

$$\frac{\partial^2 u}{\partial t^2} = C_1 \frac{\partial^2 u}{\partial x^2} + C_2 \frac{\partial^3 u}{\partial t \partial x^2} \quad (4)$$

where “u” is particle displacement, “t” is time, “x” is the coordinate in the direction of wave propagation,  $C_1$  is the stiffness ( $\frac{m^2}{s^2}$ ), and  $C_2$  is the damping ( $\frac{m^2}{s}$ ). Equation (4) reduces to the elastic wave equation when the damping value,  $C_2 = 0$ . In that case, the phase velocity is constant for all frequencies, and the wave does not experience any decay (for a 1-D plane wave). In the elastic case, the phase velocity will be  $\sqrt{C_1}$  .

In the more general case,  $C_2 \neq 0$ , and there will be both velocity dispersion and exponential, inelastic amplitude decay. A solution of equation (4) is

$$u(x,t) = \exp(-\alpha x) \cdot \cos(\beta x - \omega t),$$

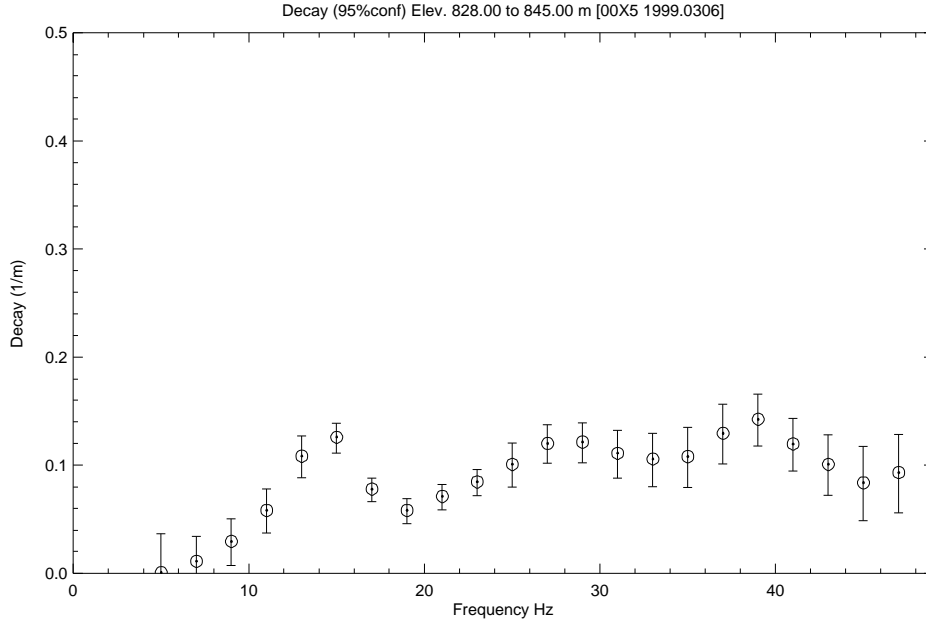


Figure 30: BAMP: SH body-wave amplitude decay for down-hole data same as seen in Figure 28 velocity dispersion. Corrected for beam spreading, a viscous, Kelvin-Voigt material, the decay should increase with frequency (Michaels, 1998).

where the wavenumber is complex and given by  $\beta + i\alpha$ . Michaels (1998) shows that the inelastic decay of a plane wave will be given by

$$\alpha = \frac{4\sqrt{D}\omega^2 C_2}{(2\omega C_2)^2 + D^2}$$

where  $\omega$  is angular frequency (rad/s) and the quantity,  $D$ , is given by

$$D = 2 \left( C_1 + \sqrt{C_1^2 + \omega^2 C_2^2} \right). \quad (5)$$

The phase velocity,  $c$ , varies with frequency according to the following relationship

$$c = \frac{2\omega^2 C_2}{D\alpha}. \quad (6)$$

The values for  $C_1$  and  $C_2$  can be expressed in terms of the following :

$$C_1 = \frac{(\beta^2 - \alpha^2) \omega^2}{(\beta^2 + \alpha^2)^2}, \quad (7)$$

and

$$C_2 = \frac{2\alpha\beta\omega}{(\beta^2 + \alpha^2)^2}. \quad (8)$$

Determination of  $C_1$  and  $C_2$  is by nonlinear joint inversion of the phase velocity,  $c$ , and inelastic decay,  $\alpha$ , over a range of frequencies. The inversion is currently performed in the Octave procedure, *cainv3.m*. Initial estimates of stiffness and damping are obtained at the frequency corresponding to the largest  $\alpha$  measured by *bamp*. First,  $C_1$  is found by evaluation of equation (7). In that computation,  $\beta = \frac{\omega}{c}$ . Then,  $C_2$  is estimated from equation (8).

RUNNING CAINV3:

Start an octave session, type `cainv3`

GUI, use mouse to pick min and max frequencies, click OK

and then use the mouse. Horizontal position is all that is read. Focus one of the panels.

You can exclude some frequencies, and that will create an fbx vector. If you include all frequencies, you may get an error statement (since it can't write out something that does not exist). Typically not a problem when you run caplot3.m later. Don't worry about it.

GUI, C1=stiffness, C2=damping initial estimate for the 3rd order wave equation.

GUI, Choose weighting

GUI, Choose balance between damping and velocity, .5 good idea

Plots, update as inversion progresses

GUI, continue LSQE plot

GUI, continue Chi squared plot

GUI, save results to disk, yes if you want to run caplot3.m

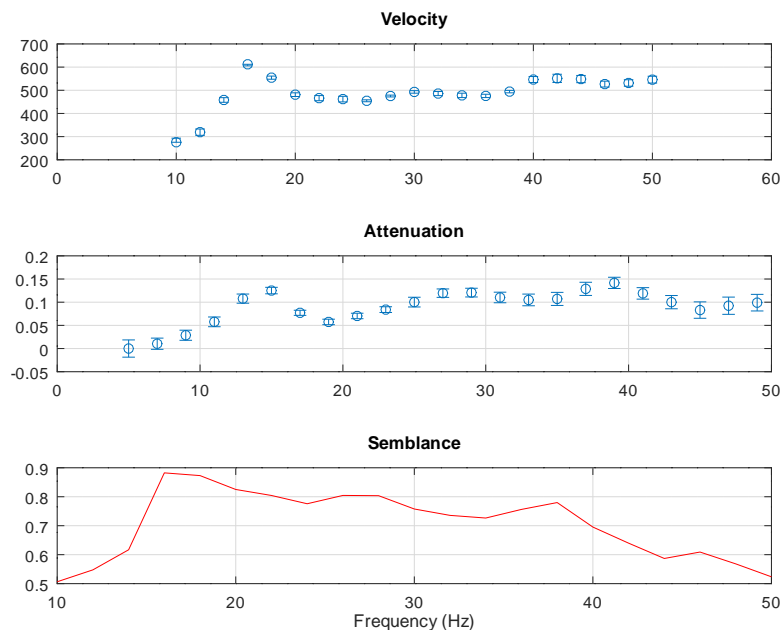


Figure 31: CAINV3: First display. Use mouse to pick frequency limits for analysis, low and then high.

After running `cainv3`, you may wish to make nice plots. For this, there is program `caplot3` (8.2.8).

### 8.2.8 OCTAVE CAPLOT3

Once an acceptable solution is determined for stiffness and damping by running `cainv3` (8.2.7), improved plotting of results are done with this program.

Running `caplot3`:

Start an octave session, type `caplot3`

GUI, show grid?

GUI, enter limits to plot

GUI, only plot data used, or plot all data (recall mouse selection of range of data to include above)

Plot, save or plot, shows fit and error bars with data observed

GUI, Go to Attenuation?

GUI, select axes limits

Plot, save or plot, shows fit and error bars with data observed  
 GUI, go to Chi<sup>2</sup> plot?  
 GUI, select axes limits  
 Plot shows change in Chi<sup>2</sup> by iteration

NOTE:

Plot Titles show C1, C2, Relaxation time

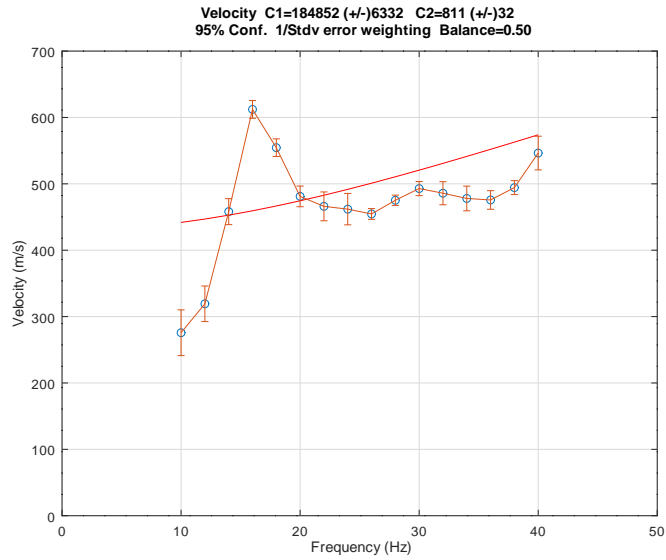


Figure 32: CAPLOT3: Plot of velocity dispersion, measure and calculated (solid line) only over frequency range used in cainv3 (8.2.7). Weighting by reciprocal of standard deviations.

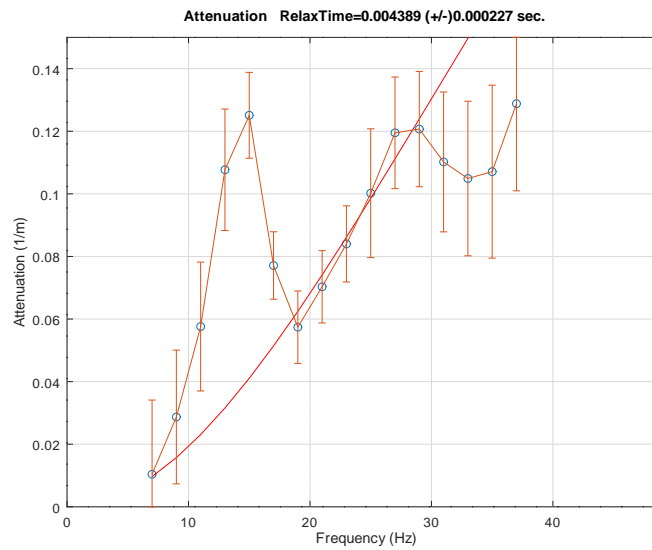


Figure 33: CAPLOT3: Plot of decay, measure and calculated (solid line) only over frequency range used in cainv3 (8.2.7). Weighting by reciprocal of standard deviations. Relaxation time about 4 msec. Relaxation time is  $T_r = \frac{C_2}{C_1}$ .

### 8.3 Relating Permeability to Damping KVMB

If porosity information is available, it can be combined with stiffness and damping results (viscoelastic, KV) under an alternative constitutive model (KVMB) to estimate permeability. This would not be absolute, but rather relative permeability (hydraulic conductivity, units of meters/second). The theory is found in [Michaels \(2006\)](#).

While the constitutive model is structured on highly simplified assumptions, it captures the behavior of granular soils saturated with water or other fluids when shaken by S-waves. Inertial damping resulting from shaking is predicted to peak at some hydraulic conductivity. Damping decreases on one side of the peak due to pore sizes being too small to permit significant relative motion between the frame and fluids. On the other side of the peak, damping decreases because the pores are so large that fluid moves easily with respect to the frame.

There are four octave programs provided with BSU that may be used with the KVMB soil model. Note, the intention is that this model is only valid in the context of **granular soils** under the assumption of **inertial damping** and **laminar flow**. The first three are forward problems, the 4th listed below is an inversion program.

- **OCTAVE\_kdKVMBscan.m** computes and plots KV damping ratio as a function of either hydraulic conductivity or uniform pore diameter (user option). User provides porosity and frequency of shaking. See [Figure 34](#).
- **OCTAVE\_fqKVMBscan.m** computes and plots KV damping ratio as a function of frequency for a user provided porosity and hydraulic conductivity. See [Figure 35](#).
- **OCTAVE\_kvKVMBscan.m** computes and plots KV (Kelvin-Voigt) damping ratio vs. KVMB (Kelvin-Voigt-Maxwell-Biot) damping ratio. See [Michaels \(2006\)](#).
- **OCTAVE\_KD4kvmb.m** [8.3.1](#) Inversion code that combines **frequency, porosity, stiffness, damping** to compute KV **damping ratios** and KD **hydraulic conductivity**. See the procedure [8.3.1.1](#).

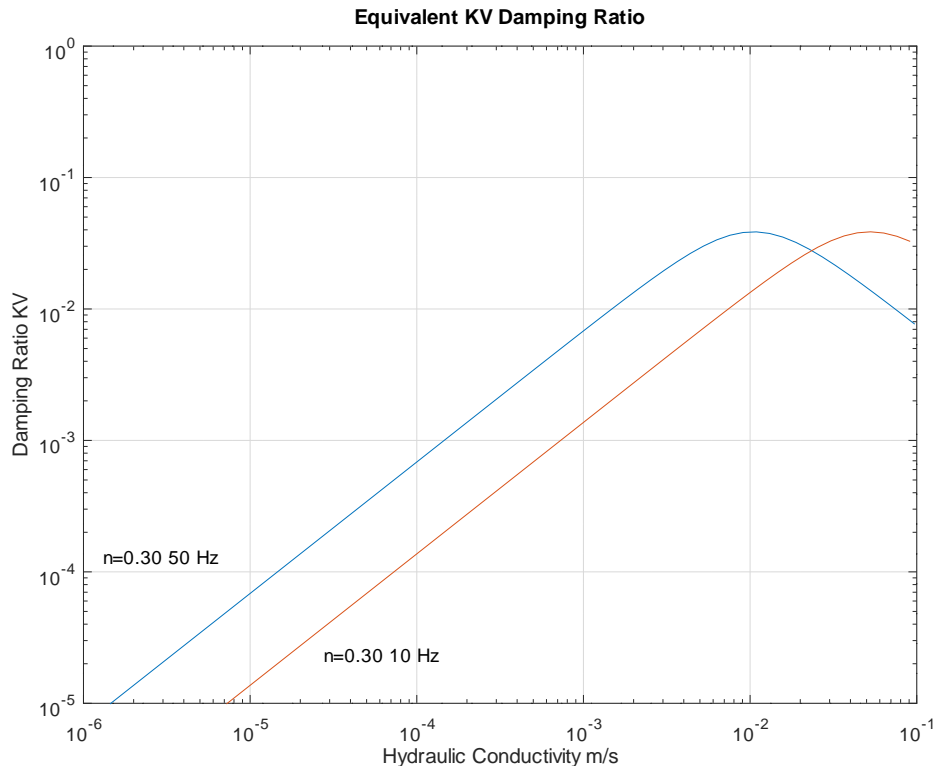


Figure 34: kdKVMBscan.m: Plots Kelvin-Voigt damping ratio vs. hydraulic conductivity for user provided porosity and frequency of shaking. Here, porosity is 30% and frequencies are 10 and 50 Hz. Left of the peak is coupled motion (small pores, fluid largely moves with frame). Right of the peak is uncoupled motion (large pores).

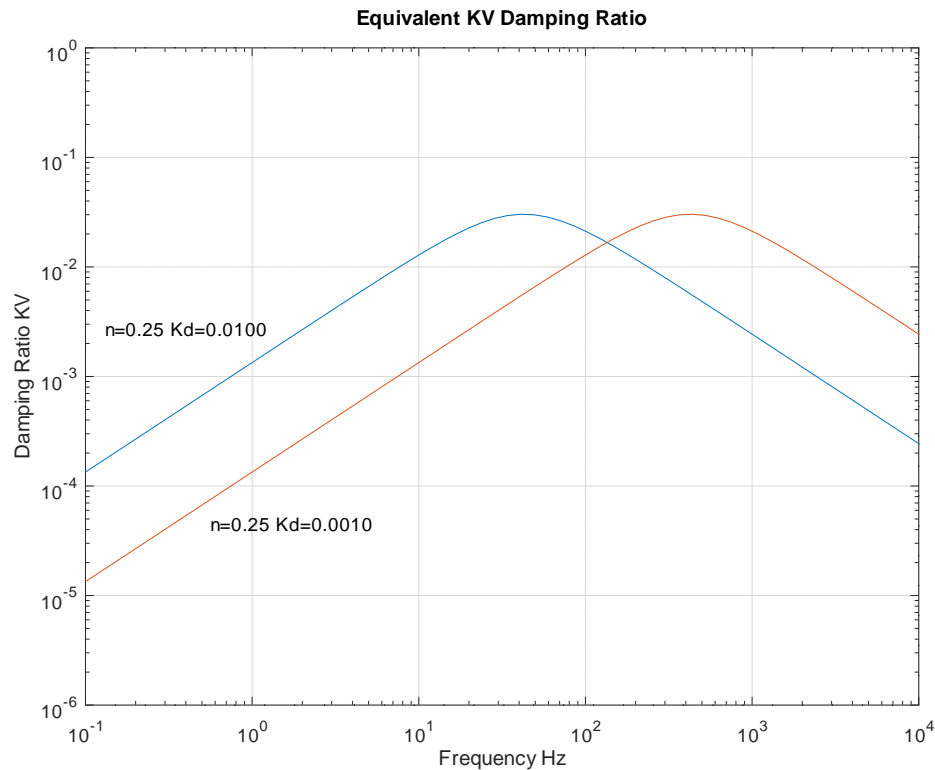


Figure 35: fqKVBscan.m: Plots Kelvin-Voigt damping ratio vs. frequency for user defined porosity and hydraulic conductivities. Here, porosity set at 0.25, two different cases of hydraulic conductivity  $K_d = .01$   $K_d = .001$   $m/s$ .

### 8.3.1 OCTAVE\_KD4kvmb

This Octave program combines the **C1, stiffness** and **C2, damping** wave equation coefficients (Equation 4) determined from body wave dispersion inversion (CAINV3, 8.2.7) with **porosity** and a desired frequency to return solutions for the Kelvin-Voigt (KV) **damping ratio**. There will be up to two possible solutions. The likely solution is the coupled solution, but the program returns the uncoupled solution as well. The coupled solution is more likely since most earth materials are not permeable enough to result in the uncoupled solution.

**8.3.1.1 Hydraulic Conductivity Procedure** The general procedure is as follows:

1. Drill a bore hole.
2. Do a down hole survey with a source that generates SH-waves.
3. Invert body wave dispersion and decay with CAINV3 8.2.7. This will provide **stiffness**, C1, and **damping**, C2 values under a Kelvin-Voigt model (Equation 4).
4. Select a relevant **frequency** and **porosity** and compute Kelvin-Voigt **damping ratios** using KD4kvmb 8.3.1. The coupled damping ratio is the most likely one. The program will also return the corresponding hydraulic conductivities,  $KD$  in  $m/s$ .

Note this solution should agree with figure like Figure 34 when that figure is computed for the same relevant frequency.

Why use the Kelvin-Voigt (KV) constitutive representation for a soil? The problem with the KV model is that frame and fluid masses are lumped together as one. The KVMB representation frees the two masses to move which leads to an estimate of permeability. Engineering practice has been to use the KV representation, as in resonant column analysis. The CAINV3 joint inversion of wave dispersion and decay to values of stiffness and

Enter Parameters

Frequency (Hz)  
12

n (porosity)  
.25

C1 stiffness (m<sup>2</sup>/s<sup>2</sup>)  
10000

C2 damping (m<sup>2</sup>/s)  
5

+/-stdevC1  
100

+/-stdevC2  
1

+/-stdev{porosity}  
.01

OK Cancel

Figure 36: Prompt for input in KD4kvmb.m run

damping follows that same KV practice. However, the KVMB representation can also be used to map up to two (coupled and uncoupled) cases of equivalent KV damping. It is also possible that there will be only one result of KV damping if one is at the peak of the curve. Consider drawing a horizontal line to intersect a curve like those in Figure 34.

An example of running KD4kvmb is shown to illustrate the final step of the procedure 8.3.1.1. The results of the run shown in Figure 36 are:

```
SOLUTION (+/- 95 Percent Confidence)
Freq=12(Hz) Resonator_L=1.33(m)
Damping Ratios: Peak=0.030293 Wave=0.018850 (+/-0.01450)
Coupled (b_case): DR=0.018850 KD=0.01224(+/-0.0117m/s)
UnCoupled (a_case): DR=0.018849 KD=0.09169(+/-0.0881m/s)
Porosity: 0.250 (+/-0.038)
Relaxation Time Tr=C2/C1=0.50 msec
```

The notation is as follows:

- **Peak** damping ratio (DR) is the theoretical maximum for the case at hand.
- **Wave** damping ratio is the result of the CAINV3 joint inversion of a down-hole SH-wave survey.
- **KD** is the estimate of hydraulic conductivity ( $m/s$ ). There are two of these unless the wave equals the peak damping ratio.
- Relaxation time is in milliseconds. It is analogous to the time it takes a sponge to recover its shape when squeezed under water and then released. The more permeable the sponge, the quicker the water can enter the sponges pores.

## 8.4 Refraction Shooting

There are a number of programs for refraction surveys.

- **BRED** (7.0.1) apply linear or hyperbolic correction in time to traces. This sometimes makes picking first breaks easier. It can also be used in flattening reflections for a totally different type of problem. See section 7.0.1 for more.
- **PICRESTORE** (8.4.1) Restore segpic.m picks on data reduced by BRED. 6.0.12
- **BPIC** (8.4.3) automatic first break picker, or inserts picks from a file (segpic.m 6.0.12 or suxpicker SU program).
- **BSHF** (12.1.1) static shift by header pics to QC picks.
- **BDAT** (8.4.5) datuming program for refraction data (adjusts data to the shot elevation using an overburden velocity)
- **BREF** (8.4.6) Direct wave (8.4.6.1), and Refraction (8.4.6.3) analysis setup.
- **OCTAVE DELAYTM** (8.4.6.3) delay time solution.
- **OCTAVE DELAYTMR** (8.4.6.4) reciprocal delay time solution (cross-river shooting).

**8.4.0.1 BRED Example Flow** Using BRED to flatten the refracted arrivals in time is not necessary, but can make it easier to pick first breaks. See **BVEL** 8.2.2 for the down-hole version. That is, it aids in identification of the refracted arrival (if the reduction velocity is approximately equal to the apparent horizontal velocity of the refracted arrival). It can also get the refracted arrival in a more confined time window, and this permits scaling the display for better resolution of the first arrivals. Here is an example flow. The situation is sandy soil over bedrock. The bedrock velocity is about 3000 m/s. See figure 14 for the original and reduced in time data plot.

```

bred k008.seg 1 3000. .05

octave
  segpic
    bredk008.seg (input file prompt answer)
    3 for clip, 0.15 for maximum time
    use mouse to pic first major down deflection
    (this produces output file bredk008.pic )
    exit

picrestore bredk008.dat bredk008.pic > out.pic

bpic k008.seg 1 out.pic 0.

mv bpick008.seg k008.seg

bshf k008.seg 0 1 .05

```

First, the data are reduced by a 3000 m/s velocity, bulk shift of .05 seconds to make picking easier. An octave session is started and the segpic program is run (requires segpic.m, segyinfo.m, bsegin.m files be in the directory). Exit octave, and note that file bredk008.pic contains the picks as two column text file (channel, pic time). Picks are corrected for the reduction velocity by running picrestore (see 8.4.1). The picks are then uploaded to file k008.seg by program BPIC 8.4.3. This produces file bpick008.seg which is then renamed k008.seg using the move command. A quality control (QC) check is done by using program BSHF 12.1.1, static shifting the data to align on .05 seconds using the header values uploaded during the BPIC step. Figure 37 shows the data aligned on .05 seconds (red arrow). Now file k008.seg is ready for further refraction analysis.



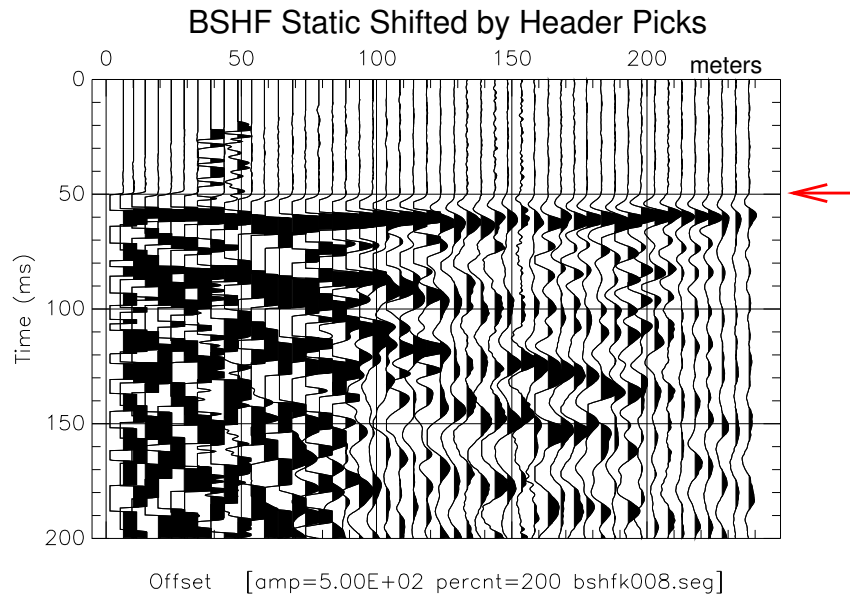


Figure 37: BSHF: After picks uploaded to headers with BPIC, data are static shifted to align on .05 seconds using header values. This is a quality control step. See example flow, section 8.4.0.1.

#### 8.4.1 PICRESTORE

Only required if data have been reduced by a velocity with BRED (7.0.1). Since the data have been reduced, we need to adjust the pic file to original recording time. File **bredk008.dat** has recorded the static shift that were applied in the BRED process. This \*.dat file is combined with the \*.pic file in **picrestore** to restore the picks themselves to the original time. Program BPIC is then run with this corrected pic file (out.pic in the example, created by redirection).

```
usage: picrestore bredxxxx.dat bredxxxx.pic >out.pic
FLOW
1)bred flattens data, saved: bredxxxx.dat
2)OCTAVE: segpic.m pics 1st breaks
segpic.m outputs bredxxxx.pic (tr,pic)
3)picrestore stdout stream:
adjusted pics, removing the flattening
4)bpic xxxx.seg 1 out.pic
(pics inserted into original file before bred)
```

#### 8.4.2 BMRK

Picking times on seismic data can be done a number of ways, and program **BPIC 8.4.3** is used to insert the pick times into the headers. This program will insert a spike (either positive or negative) at the pick time for quality control.

```
bmrk infile ipol fscl

infile =name input file to mark picks
ipol   =polarity of tick mark
       -1=negative
       1=positive
fscl   =scale factor to multiply tick mark size
```

Figure 38 shows an example of how it displays.

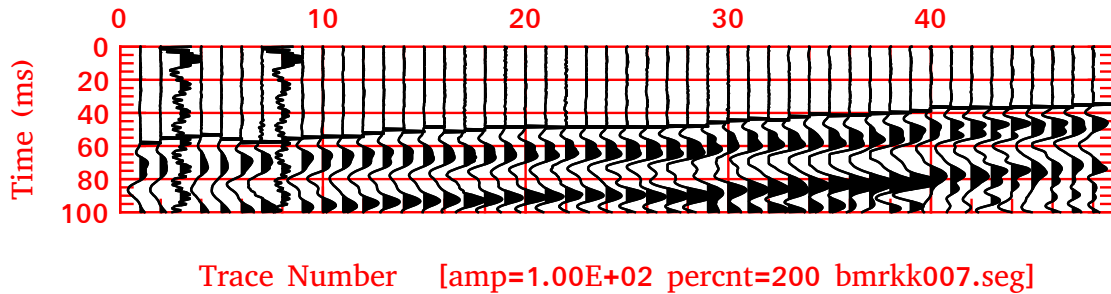


Figure 38: BMRK: Inserting a + spike to mark pick times.

### 8.4.3 BPIC

This program can either auto pick first breaks, or accept a text file with picks from another process like **sepic** 6.0.12. I recommend NOT using the autopicker (iswop1=0). It is usually better to manually pick data, particularly in noisy environments. The command line arguments to consider are below the dashed line, IF ISWOP1 = 1 or 2. See section 8.4.0.1 for an example flow where this program is used. The autopick command line arguments can be seen by typing **bpic -h** in a terminal, or using the man page for BPIC (**man bref**).

-----IF ISWOP1 = 1 or 2-----PIC FILE-----

```
bpic infile iswop1 picfil bulkst
```

```
infile
  iswop1:  option 0=autopic
           1=UPload pic file to header
           2=DOWNload headers to pic file
  picfil  =name of pic file
  bulkst  =bulk static shift to add to picks in sec
```

### 8.4.4 BSHF

This program can use static shifts to align data on either header pick values or values in a separate \*.pic file.

```
BSHF infile ipic ishf tshift picfil
```

```
infile:  =input file name
ipic:    1=static shift data by picks pick file
         0=static shift data by picks in headers
ishf:    0=add static shifts
         1=subtract static shifts
tshift:  =bulk static to add to picks
picfil:  =file name with picks (for ipic=0)
```

For an example, see the flow in section 8.4.0.1. See Figure 37 for a QC alignment check.

### 8.4.5 BDAT

Not all refraction data are shot with a source on the surface of the ground, and the ground is not always flat. BDAT can shift data by static shifts to datum the data to the shot elevation. This program is often used with buried explosive shots. The up-hole time can provide a value for the overburden velocity.

```
bdat infile vel iapply
```

```
infile = input file name
vel    = velocity of overburden
iapply = switch
        1= apply static to datum
        -1= remove static, restore to observed
```

NOTES: 1 Statics are computed from geometry in headers.  
The receiver is datumed to the shot elevation.

2 Intended use is to make recognition of first arrivals easier, and picking easier.

3 First break pick headers are adjusted with each datum change

BDAT uses a weathering zone approach, and computes the static shift using the header values:

- **sz** Shot elevation (top of hole if buried)
- **sd** Shot depth
- **rz** Geophone elevation

The static shift is computed as:

$$T_s = \frac{(sz - sd - rz)}{vw} \quad (9)$$

where  $vw$  is the weathering velocity.

### 8.4.6 BREF

BREF can build matrices for ground consistent inversion of either direct waves, or refracted head waves. The method is the delay time method for refractions, and the formulation is limited to a single refractor under overburden (Michaels, 1995). One typically decides on the offset where the transition from direct wave to refracted wave arrivals occurs (cross-over distance, Figure 13, OCTAVE refplot). The arguments are:

```

bref line# nshots rmin rmax irefdir irecip infil1 infil2 infil3

line#: =line number for refraction/direct analysis
nshots =number of shot records to use (<=10)
rmin   =minimum offset to include
rmax   =maximum offset to include
irefdir =0 refraction analysis
        =1 direct wave analysis
irecip  =0 normal shooting, shot fixed
        =1 reciprocal shooting, receiver fixed
infil1  =input file 1
infil2  =input file 2
infil3  =input file 3 etc....

```

Note that the argument, **irefdir** determines if the setup will be for direct wave or refracted wave. The following data examples are taken from Michaels (1999). Also, see the **bsu-user-guide3-1.pdf** for more.

**8.4.6.1 Direct Wave** In this example, we have 3 shot profiles, k004 is split spread. The minimum and maximum offset range for direct wave analysis is 0 to 30 meters. See Figure 13 to see how a cross-over distance is estimated. The command issued from a terminal is:

```
bref 008 3 0. 30. 1 0 k004.seg k008.seg k009.seg
```

The output files are:

- **G008** system matrix columns: (shot, receiver, offset)
- **D008** data vector, direct wave (arrival time,channel)
- **E008** elevation vector, (trace number, station number, elevation)

Analysis requires program **direct.m**. Start an Octave session and type **direct** to start the program. Answer the GUI questions. The result and a file **plot.ps** will be output (see Figure 39).

**8.4.6.2 Theory** The basic travel time equation for the direct wave between shot A and geophone 1 is

$$X_{a1} \cdot \frac{1}{V_1} = t_{a1} \quad (10)$$

where  $X_{a1}$  is the distance between the shot A and geophone 1. The overburden velocity is given by  $V_1$  and the observed first arrival time is  $t_{a1}$ .

We set up a matrix problem in the form

$$G \cdot m = d \quad (11)$$

which expands to

$$\begin{bmatrix} X_{a1} \\ X_{a2} \\ X_{b8} \\ X_{b9} \end{bmatrix} \cdot \begin{bmatrix} 1 \\ V_1 \end{bmatrix} = \begin{bmatrix} t_{a1} \\ t_{a2} \\ t_{b8} \\ t_{b9} \end{bmatrix} \quad (12)$$

The ordinary least squares (OLS) solution is given by [Menke \(1989\)](#)

$$m = [G^T G]^{-1} G^T \cdot d \quad (13)$$

It follows that the overburden velocity determination is  $V_1 = \frac{1}{m}$ .

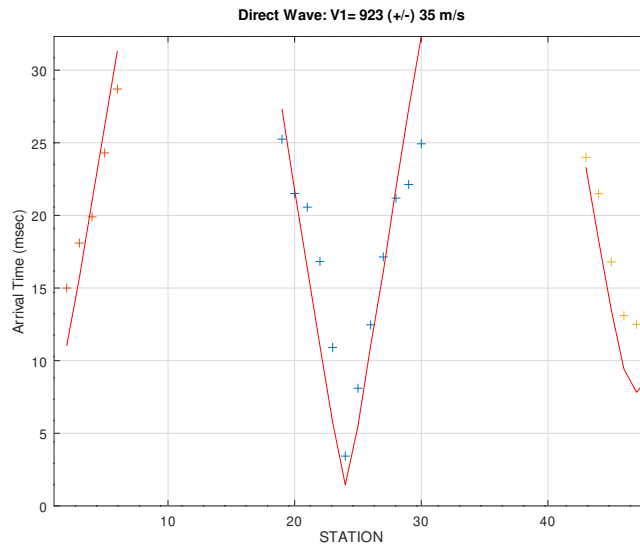


Figure 39: BREF: Output plot.ps for direct wave analysis. Title shows the least squares solution for the overburden velocity,  $923 \pm 35 \text{ m/s}$ . Range of offsets  $0 \rightarrow 30 \text{ m}$ .

**8.4.6.3 Normal Delay Time Refraction** In normal shooting, we work with shot gathers rather than geophone gathers in the reciprocal approach 8.4.6.4. Shot profiles k008.seg and k009.seg are used in this example. Offsets from 50 to 250 meters are used. The BREF command is

```
bref 0809 2 50. 250. 0 0 k008.seg k009.seg
```

The output files are:

- **G0809** system matrix columns: (shot, receiver, offset)
- **D0809** data vector, refracted wave (arrival time,channel)
- **E0809** elevation vector, (trace number, station number, elevation)

An observant reader will ask, why not include the split spread, k004.seg? That would be better, and one would normally do that. However, by taking only two reverse profiles, I can show you how to add constraint equations when needed. With only the two lines, the system matrix, G0809, will be singular. The problem is a lack of reverse profiles in the near offset ranges. Receivers in the offsets 50 meters and beyond all receive signals from both sources at the end of the lines, so they are OK. To get a solution, we need to add a couple of extra lines at the bottom of the G0809 matrix.

**Constraint** Make the shot and nearest geophone for that shot have the same delay time.

The first two columns of the G matrix correspond to the shots, k008 and k009 (columns 1 and 2 respectively). The first shot, k008, has a near geophone in column 13. We create a new row at the bottom of the matrix by placing a 1 in column 1 (for the shot) and a -1 in column 13 (for the nearest geophone with a refraction). In the last column, we put a zero instead of a distance since this is a constraint equation,  $T_{shot} - T_{geophone} = 0$ . To get the zero, we add a row to the bottom of the data vector, D0809. We put two zeros in this last row (one for the time column, one for the station number). Again, this is a constraint equation, not a data equation.

We do this again, adding one more row to the G matrix. This time, a 1 in column 2 (for shot k009). The negative one (-1) goes in column 39 corresponding to the nearest geophone with a refraction for shot k009. We then edit D0809 data vector with a pair of zeros as above. Note: *We leave E0809 alone, no need to change it.*

The rest of the matrix above the constraint equations are simply delay time equations.

$$T_{shot} + T_{geophone} + \frac{X_{sg}}{V_2} = T_{obs} \quad , \quad (14)$$

where  $T_{shot}$  is the shot delay time,  $T_{geophone}$  is the geophone delay time,  $X_{sg}$  is the distance on the surface of the ground between shot and geophone, and  $T_{obs}$  is the observed arrival time from the first break pick. The refractor velocity is given by  $V_2$ . Details on this approach are given in [Michaels \(1995\)](#). The solution is found by a weighted least squares (weighting minimizes the roughness of the solution, ie. makes it smoother at the slight cost of a poorer fit).

```
RUNNING delaytm.m
Start octave, type
delaytm
GUI, change G001 to G0809, etc
GUI, number of shots = 2
GUI, smoothness weight 0.1
GUI, shows refractor velocity =4122 m/s and shot delay times of 10.3 and 12.4 msec, OK
Plot showing delay times for geophones
GUI, overburden velocity 923
Plot shows ground elevation and refractor indicating a variable soil thickness.
Alternative solution, default to 10 meters, Plot shows how an alternative extreme
solution of constant soil thickness with overburden
velocity varying.
GUI, 2 constraint equations, OK
Plot shows fit of solution to observed times.
Preference is for variable soil thickness solution based on geologic context.
```

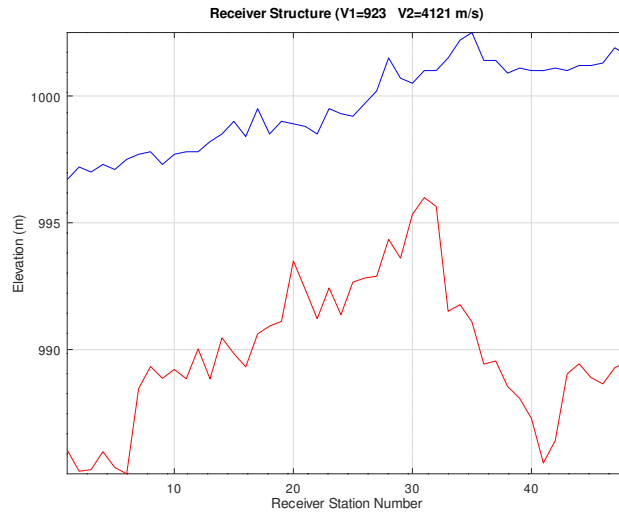


Figure 40: OCTAVE DELAYTM: Structure solution for shots k008 and k009. Ground surface in blue, top of bedrock in red. Soil velocity 923 m/s between blue and red. Bedrock velocity 4121 m/s.

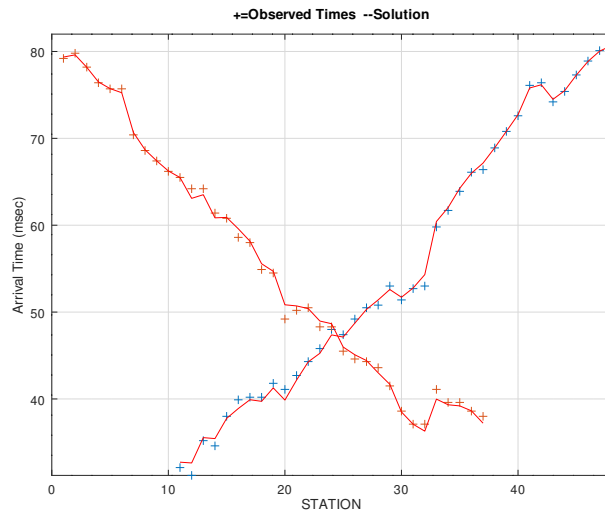


Figure 41: OCTAVE DELAYTM: Computed solution and observed times for k008 and k009.

**8.4.6.4 Reciprocal Delaytime Refraction** The difference between conventional refraction shooting 8.4.6.3 and reciprocal refraction shooting is that the former employs shot gathers and in this case, geophone gathers. The analysis is essentially the same. Reciprocal shooting is applied when crossing a river. Placing geophone in the river would subject the geophones to a noisy environment, particularly with a strong current bouncing the phones around. When an existing bridge needs replacement, one can deploy an airgun source from the bridge at stations of about 5 meters and record into geophone arrays on the river banks. An example of this approach is shown in Figure 42.

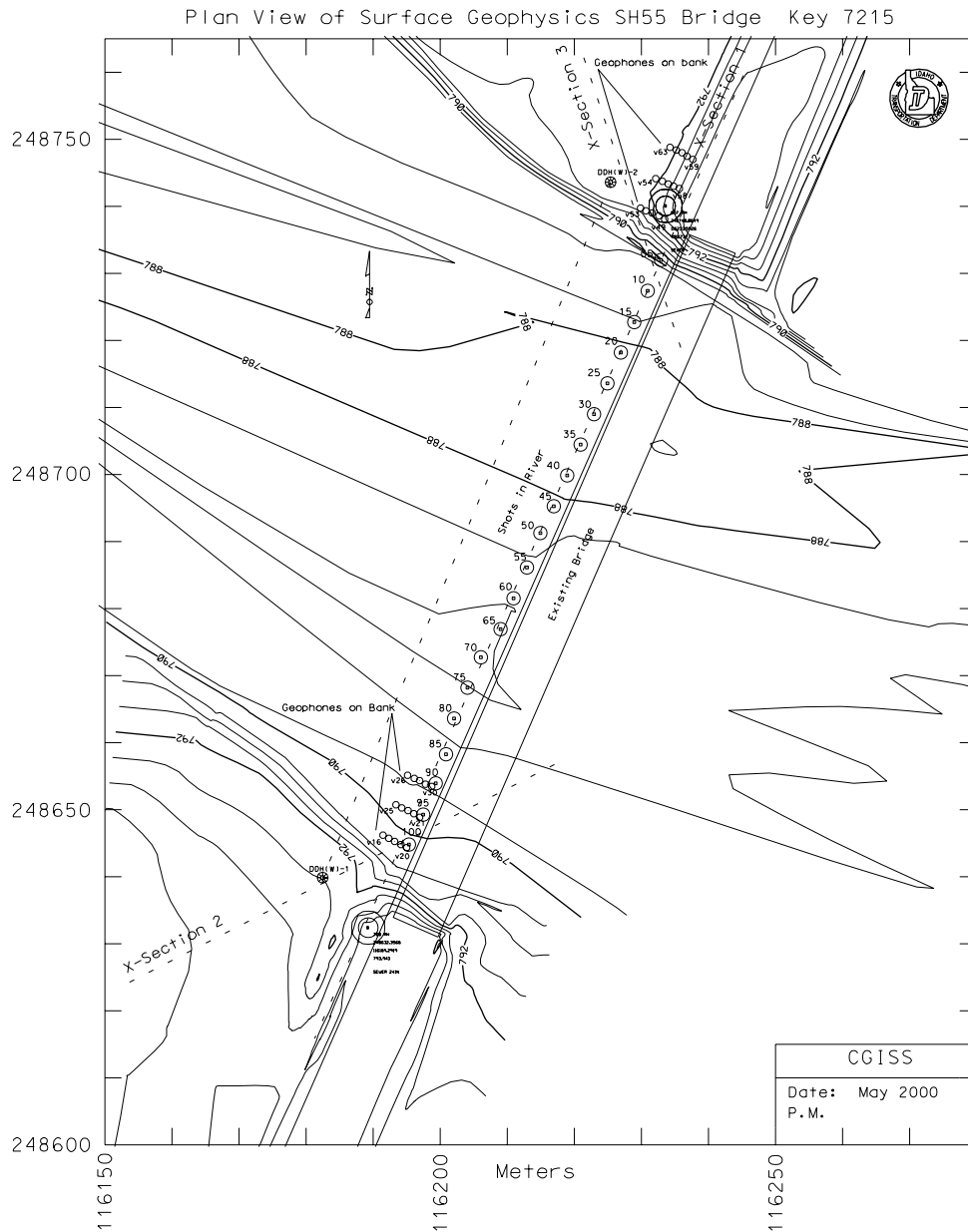


Figure 42: OCTAVE DELAYTMR: Reciprocal shooting across a river. Airgun source deployed at stations across bridge (Michaels, 2001a).

The geophone arrays can be summed to cancel traffic noise, beam forming to receive signals from the air guns. Recorded data are sorted into 3 northern geophone gathers, and 3 southern geophone gathers using the BEXT program (N000.seg, N001.seg, N002.seg, S001.seg, S002.seg, S003.seg). BREF is run with the following command:



```

  bref 001 6 10. 110. 0 1 N000.seg N001.seg N002.seg S000.seg S001.seg S002.seg

```

The BREF program produces output files G001, D001, and E001 as in the normal shooting example (8.4.6.3). Some editing is required. The BREF code detected one constraint equation on this run. Further, a need for 2 more constraints was found so that all delaytimes were made equal at stations 90, 95, 100. These constraints were strongly weighted (factor of 50), and the G001 matrix was edited as shown below (relevant tail of the matrix):

```

  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 1 0 0 0 0 0 0 0 0 45.363
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 1 0 0 0 0 0 0 0 40.364
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 1 0 0 0 0 0 0 35.382
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 0 1 0 0 0 0 0 30.329
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 0 0 1 0 0 0 0 20.336
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 1 0 0 0 15.075
  0 0 0 0 0 1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 1 0 0 10.564
  0 0 0 50 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 -50 0 0 0.000
  0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 -50 50 0 0.000
  0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 -50 0 50 0.000

```

The first 6 columns are for the 6 geophone gathers. The D001 matrix tail is shown here:

```

  0.03300    55
  0.03120    60
  0.02840    65
  0.02660    70
  0.02390    80
  0.01880    85
  0.01720    90
  0.00000     0
  0.00000     0
  0.00000     0

```

One extra file is required when shooting this way. One needs a water depth file which gives the depth to the river bottom below the surface shots. This is a two column ASCII text file (depth, shot#). One starts an octave session and then runs delaytmR.m:

```

delaytmR
GUI enter names of the files G001, D001, E001 and wds.data
GUI number of geophones on shore, 6
GUI smoothness weight, 0.1 a good value
GUI displays refractor velocity (2216 m/s) and geophone delay
times below geophones (7.8, 9.9, 11.0, 2.5, 8.7, 8.6 ms), OK
PLOT shows delay times under the trans-river sources
GUI enter an overburden velocity in m/s, for example 1500 m/s
PLOT showing structure if overburden velocity is as assumed
(ground or water surface, bottom of river, refractor structure)
GUI enter alternative limiting case of constant depth refractor
PLOT of alternative, constant depth refractor, V1 varies
GUI enter number of constraint equations, 3 OK
PLOT of the observed data and fit to the equations.

```

The resulting structural plot is shown in Figure 43 and the arrival time fit in Figure 44. Delay times,  $T_{dt}$ , are related to the distance from the surface to the refractor,  $H$  by the critical angle,  $\theta_c$ :

$$T_{dt} = \frac{H \cos(\theta_c)}{V_1} \quad (15)$$

In normal shooting, one solves for  $H$  using the critical angle  $\theta_c = \sin^{-1}(V_1/V_2)$ , the refractor velocity,  $V_2$ , being the last unknown in the G matrix setup, and a result of obtaining a solution. This is straight forward for the delay times on the river banks, under the geophones. Under the shots, the water layer creates some additional complexity to the problem (see the code, delaytmR.m). In short, if the refractor velocity,  $V_2$ , is greater than water velocity, then it is simple ray optics. Velocity  $V_1$  is the soil layer between the bottom of the river and the top of the refractor. To keep it simple in this example, I made  $V_1 = 1500$ , water velocity, and that would mean no ray bending from water to overburden soil. Recall that the ray parameter,  $p$ , is a constant,  $p = \sin(\theta_j)/V_j$ , and to get a critical refraction at bedrock, we need to know  $V_2$  relative to water velocity.

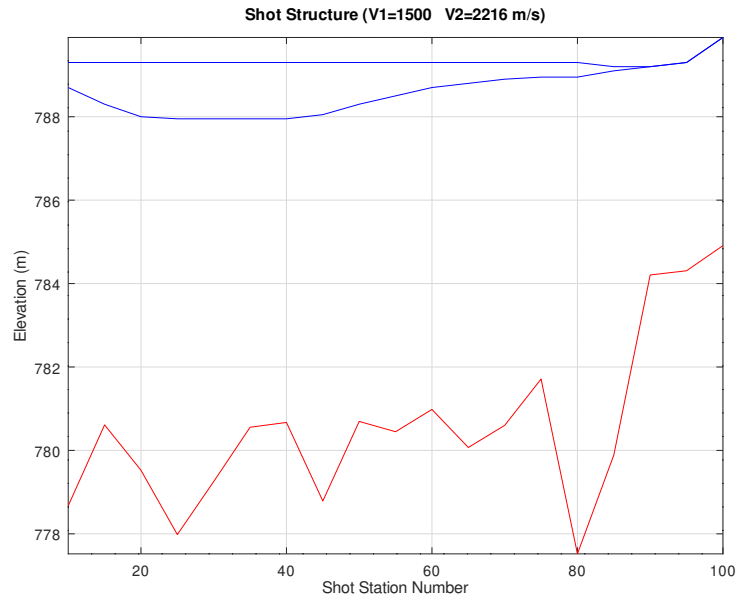


Figure 43: OCTAVE DELAYTMR: Structure assuming an overburden velocity of 1500 m/s. River water surface and bottom of river bottom in blue. Refractor structure in red.

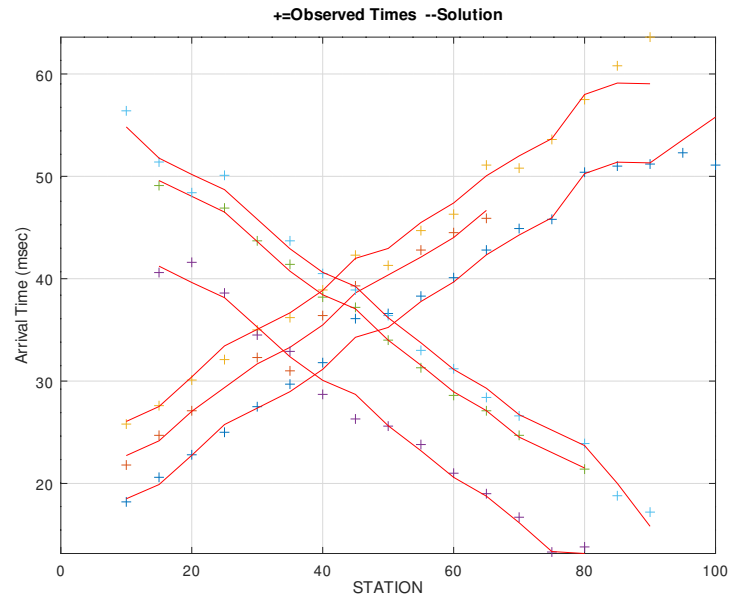


Figure 44: OCTAVE DELAYTMR: Observed arrival times and fit assuming an overburden velocity of 1500 m/s.

## 9 Forward Problem Codes

In the forward problem, a soil profile representation is used to compute a corresponding geophysical expression. The resulting geophysical data can then be compared to actual data. Another use is in planning geophysical surveys.

### 9.1 Down-Hole Seismic

Given an earth representation, compute the geophysical expression (ie. synthetic geophysical data).

#### 9.1.1 OCTAVE `cafwd3`

This code uses the same formulation as in `cainv3` (8.2.7). It can be run with data to compare to, or as a simple stand alone.

```
Start an octave session then type
cafwd3
```

##### FORWARD PROBLEM

```
Computes S-wave velocity dispersion and damping by manual input for
C1=stiffness m^2/s^2 and C2=damping m^2/s
```

```
GUI, with or without data? (requires bvas.his and bamp.his if with data)
```

```
GUI, use mouse to pick range of frequencies, OK
```

```
Plot, use mouse to limit,
```

```
GUI, C1 and C2 estimate, OK
```

```
computes fit, then
```

```
GUI, make changes to stiffness and damping for next forward calculation
```

```
iterate
```

The code also computes “Q” quality factor as a function of frequency.

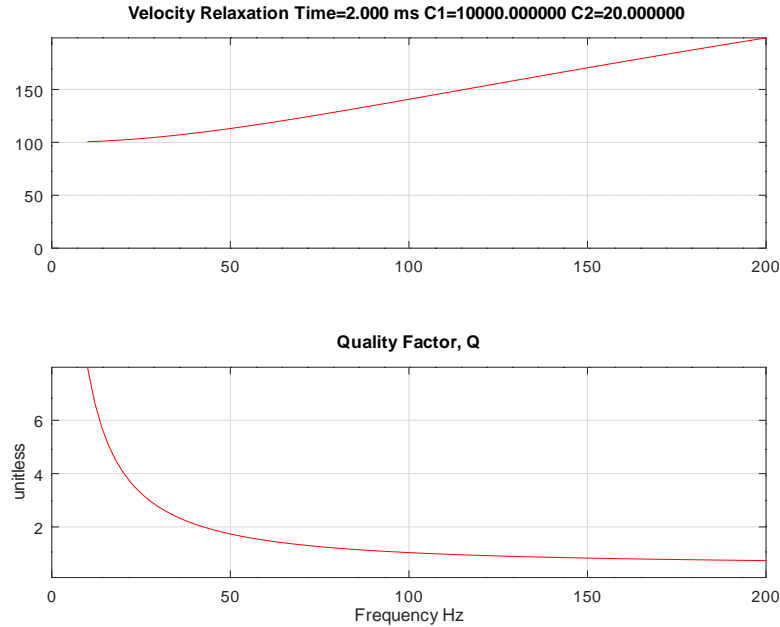


Figure 45: CAFWD3: Example without data, program's second plot showing quality factor,  $Q$ , The program's first figure plot expresses damping in terms of decay ( $1/m$  units) as in Figures 30, 31, and 33.

## 9.2 Surface Waves

BSU codes include C-language, Fortran, and Octave procedures.

### 9.2.1 OCTAVE FwdR1

This code uses the same formulation as `invR1.m` program (see section 8.1.1). The function `rwv.f` is built from `DISPER` (9.2.6). It computes a single forward problem described in the same way as in the inverse problem.

This program computes a demonstration Rayleigh wave dispersion curve. It uses `rwv.f` fortran function which must be compiled and linked to expand octave capability. The required files include:

```
rwv.f
wrapper.cpp
build_disper_oct (make executable, chmod +x if not)
```

Either run the build script first or start an octave session and be prompted to build the extension inside the octave run.

A file, like `model.txt`, contains the layered earth velocity model. Control points are interpolated by layers linearly interpolated by elastic parameters, not velocity.

Control Point file, like `model.txt`, is as follows:

1. first line is number of layers
  2. second line is S-wave velocities
  3. third line is depth values where those velocities apply.
- ```
|Example: for nlay=3 vi=shear velocity, zi=depth layer top |
| nlay |
| v1 v2 v3 |
| z1 z2 z3 |
```

GUI prompts are available for relating P-wave velocity to S-wave velocity. The layer thickness, once entered, will be held constant during the run of all other model changes.

Also plotted are actual observed data (see line 275). Running this program over and over again, one can manually invert the observed data in

bvax.his (see program BVAX), by trying to fit the observed data with a computed curve.

The disper() function returns a vector pv with fundamental and any higher modes. Here, it is demonstrated how to select and plot fundamental mode.

### 9.2.2 LAMB

Program LAMB computes a solution to Lamb's problem. This solution includes surface and body waves that radiate from a vertical impact on a half-space medium. The code is specific to a single medium property where the P- to S-wave velocity is fixed,  $V_p/V_s = \sqrt{3}$  (Lamb, 1904). For additional theory, see Mooney (1974). The command line arguments are:

```

lamb xmin dx np tmax fsamin vs den itype sfrq sdamp gfrq gdamp pol stab
xmin = minimum geophone offset (m)
dx   = spacing of geophones (m)
np   = number of geophones
tmax1 = maximum time for seismogram (s)
fsamin= sample interval (seconds)
vs   = shear wave velocity (m/s)
den   = mass density (kg/m3)
itype = type of traces output
      1= ground displacement, step function source
      2= ground particle velocity, step function source
         (or ground displacement, impulse source)
      3= ground displacement, source wavelet=damped resonator
      4= ground particle velocity, source wavelet=damped resonator
      5= geophone displacement with source wavelet
      6= geophone particle velocity with source wavelet (geophone voltage)
      7= source wavelet displacement (at source)
      8= source wavelet velocity (at source)
      9= source wavelet geophone displacement (at source)
     10= source wavelet geophone velocity (at source)
sfrq  = source wavelet high-cut frequency (hz)
sdamp  = source damping (fraction of critical, example .7)
gfrq  = geophone resonant frequency (hz)
gdamp  = geophone damping (fraction of critical, example .7)
pol    = polarity switch
      -1= SEG Sign Convention (up motion = negative = trough)
       0= TEST MODE, for display of normalized solution, NO 1/R etc.
      +1= REVERSE SEG Sign Convention (up motion = positive = peak)
stab   = stability factor, for derivative, moves pole off unit circle
         (not generally needed except for itype=8 (try stab=.16)
         since most other outputs have enough low-pass filtering)
         See function deriv for more

```

The code computes both vertical and horizontal motion (files **lambv.seg** and **lambh.seg**). The **itype** parameter selects the type of output signal.

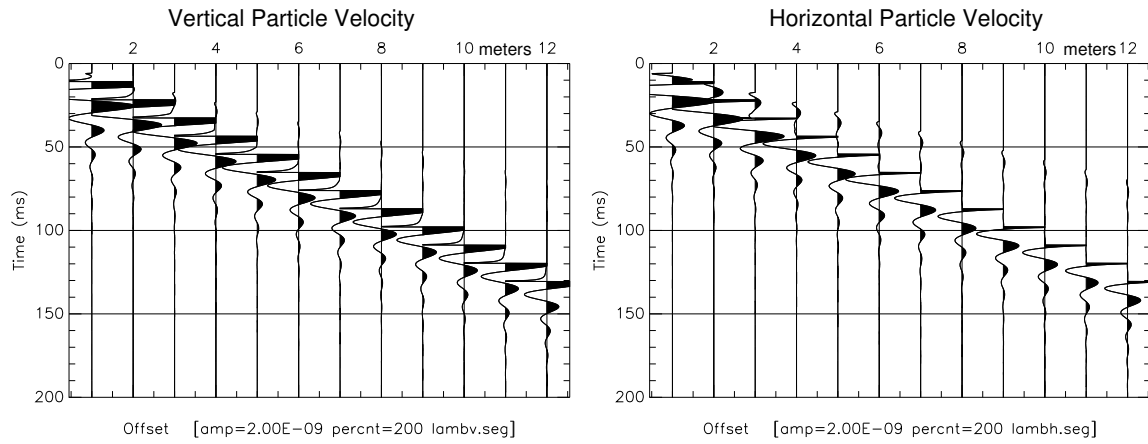


Figure 46: LAMB:Ground particle velocity solution for Lamb's problem, *itype* = 4.

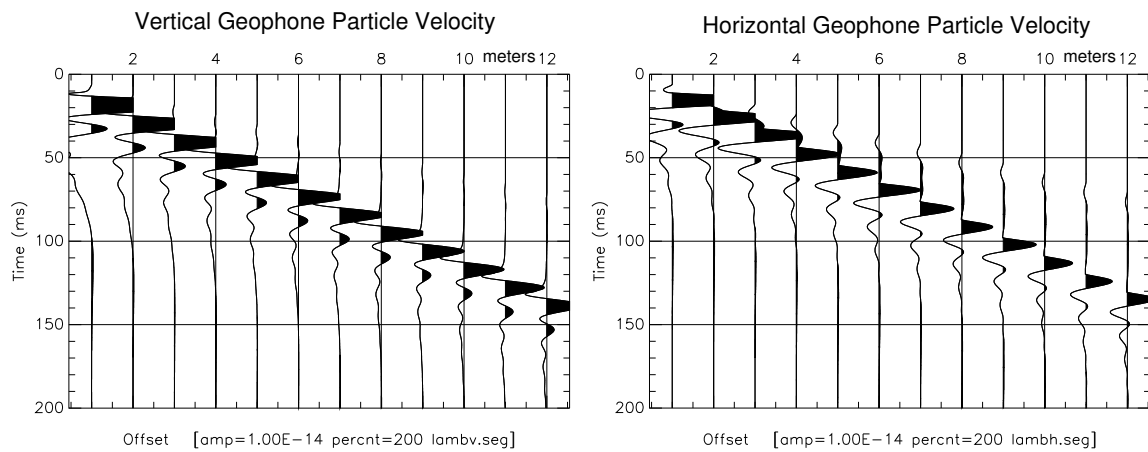


Figure 47: LAMB: Geophone (10 Hz, 0.7 damping) response, *itype* = 6.

The LAMB command corresponding to Figure 46 (*itype* = 4).

```
lamb 1 1. 12 .5 .0001 100. 1700. 4 70 .2 10. .7 -1 .2
```

The above command computes the ground particle velocity for offsets 1 to 12 meters, .5 second record, sample interval of 0.1 msec,  $V_s = 100m/s$ , density  $1700 kg/m^3$ .

LAMB was re-run to compute the particle velocity of the geophone element (which corresponds to geophone voltage) by changing *itype* to 6. This is shown in Figure 47.

```
lamb 1 1. 12 .5 .0001 100. 1700. 6 70 .2 10. .7 -1 .2
```

### 9.2.3 Near Field BNFD

The BNFD program computes the near field seismic radiation for a homogeneous whole space (Aki & Richards, 1980) (Eq 4.23, Vol. I). While not likely to correspond to any field recording, it aids in understanding the transition from near to far field without the complexity of any boundaries being present. Command line parameters follow:

```
bnfd infile xforce vp vs den alpha fc icomp ifield

infile = input file name (sets geometry, tmax, ntraces)
xforce = point force direction
        1= in positive x-axis direction
        2= in positive y-axis direction
        3= in positive (down) z-axis direction

vp      = p-wave velocity
vs      = s-wave velocity
den     = mass density

alpha   = exponential decay factor (pos) for wavelet
fc      = center frequency of wavelet Hz
        (for example, try 50 for alpha and fc)

icomp   = component of motion to output
        1= radial
        2= transverse
        3= vertical
ifield  = fields to include
        (SPN) binary coded
        0= wavelet only
        1= near field only
        2= far p-wave only
        4= far s-wave only
        3= near and far p-wave only
        5= near and far s-wave only
        6= far p- and s-wave only
        7= ALL: far S, far P, Near Field
```

Abbreviation SPN: S-wave, P-wave, and Near-field. Thus, (SPN)=(111)=7=all far S, far P, and N. For far-field P-wave only, (SPN)=(010)=2.

In the following example, all motions (*ifield*=7) are computed on the vertical and radial components. The headers are copied from an actual data set, *c008.seg*, to set number of samples, sample interval, and geometry. The commands are:

```
bnfd c008.seg 3 800. 200. 1800. 50 50 1 7
```

```
bnfd c008.seg 3 800. 200. 1800. 50 50 3 7
```

The first is for the radial (*icomp*=1) motion, the second is for the vertical (*icomp*=3) motion. Figure 48 shows motion in the vertical and radial directions. The template file, *c008.seg*, provides header data generating offset for each trace. While there are different definitions of near field in the literature, the bandwidth of the propagating wavelet (which sets the wavelet duration) and the difference between P- and S-wave velocity play the major role. The plots have been trace equalized using program BEQU to compensate for the dynamic variation in amplitude with offset.

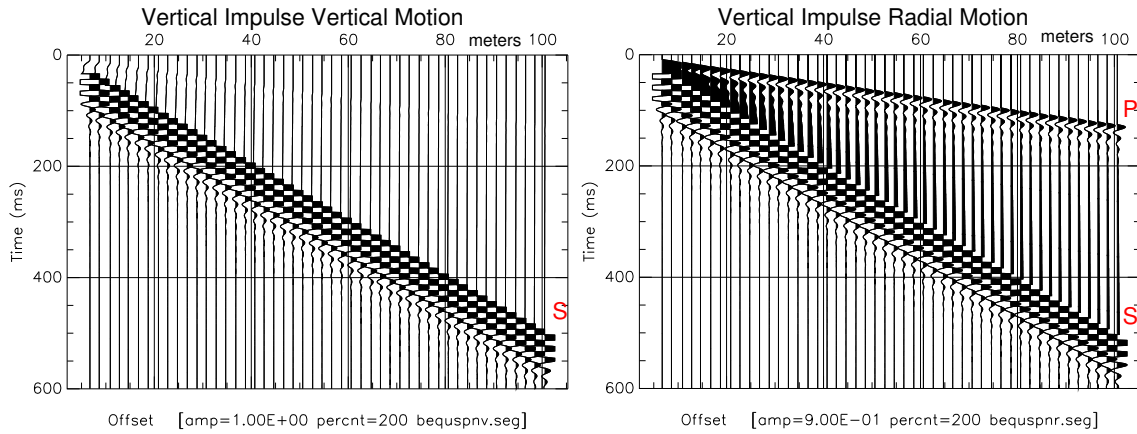


Figure 48: BNFD: Computing all fields (S-wave, P-wave, Near-field) The geometry is taken from a template file, c008.seg, and spans offsets from 7 to 100 meters. As offset increases, the far field P- and S-wave motion waxes as the near field wanes.

### 9.2.4 HALFSP

The Rayleigh wave solution for a half-space elastic medium is computed by HALFSP. The man page is read by typing `man halfsp` in a terminal. The output includes the velocity of the Rayleigh wave (`stdout`) and a file, **halfsp.tmp** which gives the motion-stress vectors for the frequency and range of depths required. An example log of this interactive program follows. One types from a terminal, `halfsp`.

```

ENTER RHO,VP,VS
1600,800,200
ENTER FREQ,NZ,Z0,ZEND
15,50,0,30
PHASE VEL= 190.2245
    
```

The units are  $kg/m^3$  for RHO (density),  $m/s$  for P- and S-wave velocities (VP and VS),  $Hz$  for FREQ (frequency), and meters for the top and bottom of the interval computed (Z0 and ZEND). NZ is the number of depth points computed. The top of the `halfsp.tmp` file:

```

HALFSP.F OUTPUT:
RHO=0.1600E+04
MU=0.6400E+08
LAME=0.8960E+09
FREQ=0.1500E+02
P-WAVE VELOCITY=0.8000E+03
S-WAVE VELOCITY=0.2000E+03
RAYLEIGH WAVE PHASE VEL= 190.2245
R1=Horiz. Displacement R2=Vertical Displacement
R3=Horiz. Stress R4=Vertical Stress
    
```

| DEPTH | R1             | R2             | R3            | R4             |
|-------|----------------|----------------|---------------|----------------|
| 0.0   | 0.2241030E+00  | -0.3974480E+00 | 0.3780026E+01 | -0.0000000E+00 |
| 0.6   | 0.1236450E+00  | -0.4410695E+00 | 0.4977292E+07 | 0.2806469E+07  |
| 1.2   | 0.5226342E-01  | -0.4611813E+00 | 0.8269704E+07 | 0.4662912E+07  |
| 1.8   | 0.2325356E-02  | -0.4647821E+00 | 0.1033810E+08 | 0.5829188E+07  |
| 2.4   | -0.3185774E-01 | -0.4570187E+00 | 0.1152439E+08 | 0.6498086E+07  |
| 3.0   | -0.5452060E-01 | -0.4416587E+00 | 0.1208169E+08 | 0.6812320E+07  |
| 3.6   | -0.6881093E-01 | -0.4214445E+00 | 0.1219681E+08 | 0.6877232E+07  |
| 4.2   | -0.7706446E-01 | -0.3983572E+00 | 0.1200719E+08 | 0.6770310E+07  |
| 4.8   | -0.8101030E-01 | -0.3738140E+00 | 0.1161344E+08 | 0.6548294E+07  |
| 5.4   | -0.8192512E-01 | -0.3488157E+00 | 0.1108883E+08 | 0.6252494E+07  |
| 6.0   | -0.8074815E-01 | -0.3240562E+00 | 0.1048634E+08 | 0.5912774E+07  |



To make a quick plot of the motion vectors, you can do something like this. Copy the `halfsp.tmp` file to a file like `data.dat`:

```
cp halfsp.tmp data.dat
```

Delete the first lines down to the first depth. So the top of the file becomes just columns of data:

```
0.0  0.2241030E+00  -0.3974480E+00  0.3780026E+01  -0.0000000E+00
0.6  0.1236450E+00  -0.4410695E+00  0.4977292E+07  0.2806469E+07
1.2  0.5226342E-01  -0.4611813E+00  0.8269704E+07  0.4662912E+07
1.8  0.2325356E-02  -0.4647821E+00  0.1033810E+08  0.5829188E+07
2.4  -0.3185774E-01  -0.4570187E+00  0.1152439E+08  0.6498086E+07
3.0  -0.5452060E-01  -0.4416587E+00  0.1208169E+08  0.6812320E+07
```

Then write a short **Gnuplot** script to plot the second and third columns as a function of the negative of the depth. Call it `plot.gp`:

```
set grid
set ylabel 'Depth (meters)'
p 'data.dat' u ($2):(-1)*($1) w l t 'horiz',\
'data.dat' u 3:(-1)*($1) w l t 'vert'
set terminal pdf
set output 'plot.pdf'
replot
```

Run the Gnuplot program from a terminal command line to produce Figure 49:  
`gnuplot -p plot.gp`

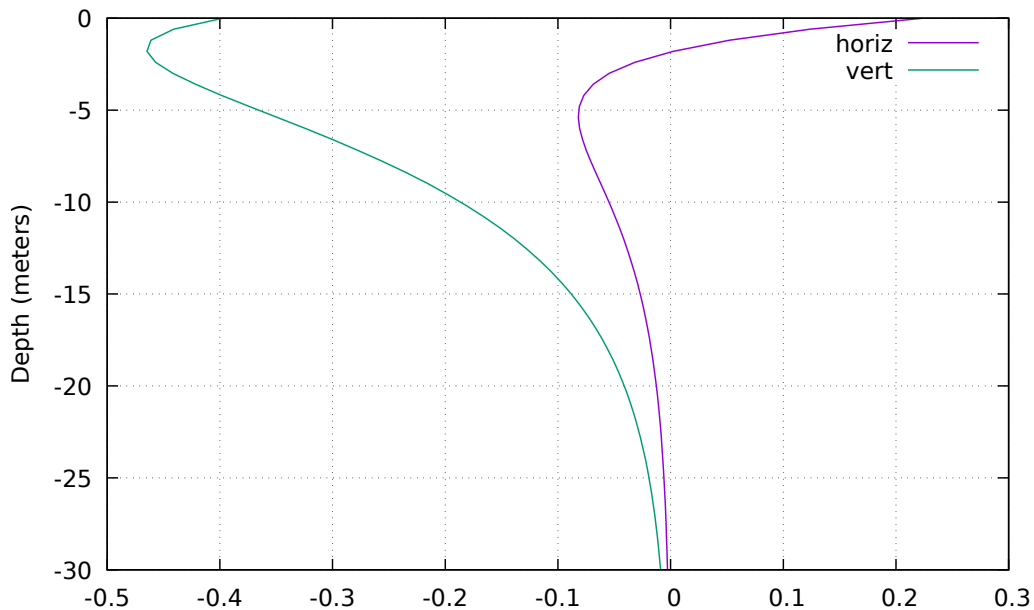


Figure 49: Gnuplot image created by the `plot.gp` script. The `-p` command line option of the `gnuplot` command makes the X11 plot persistent. Press the `q` key while mouse focus is in the figure to end the display. Then view the `plot.pdf` file with your favorite PDF viewer.

### 9.2.5 GENDIS

GENDIS is an interactive program used to prepare an input file for the **DISPER** program (9.2.6). **DISPER** is written in Fortran, and employs a namelist file for input. Use the man page for documentation (man gendis). There is no **-h** command line option since it is interactive. The following is an example log of a GENDIS run.

```

    enter: name of output file ( < 40 char)
disper.d
    enter: sample interval in seconds
.001
    enter: tmax for trace
1
    enter: minimum frequency
1
    enter: maximum frequency
100
    enter: maximum mode #
9
    enter: deltz step size
1.
    enter: number of control
3
    layer( 1) enter: beta,alpha,rho,depth(top)
100., 800., 1600., 0
    layer( 2) enter: beta,alpha,rho,depth(top)
100., 800., 1600., 1.0
    layer( 3) enter: beta,alpha,rho,depth(top)
300., 1500., 1700., 1.01
twice npts= 2048
twice tmax= 2.0480
output====>disper.d

```

The output file, disper.d, in this case is:

```

&disper
  nlay= 3,
rho= 0.1600E+04, 0.1600E+04, 0.1700E+04,
mu= 0.1600E+08, 0.1600E+08, 0.1530E+09,
lame= 0.9920E+09, 0.9920E+09, 0.3519E+10,
zi= 0.0000E+00, 0.1000E+01, 0.1010E+01,
deltz= 1.0000,
modemx=9,
nfreq=202, flo= 0.1000000E+01, delf= 0.48828122E+00, jsmax=300, ksw=0,
  pvlcty=0.0, pfreq=0.0, zend=100.0,
  ofile='disper.tmp',
  octav1='phase.m', octav2='mat2.m',
  curve='earth.crv', /

```

**9.2.5.1 SHOWMDL** This program provides an easier human view of a disper.d file. Type **showmdl disper.d** to display the file named disper.d:

```

show.tmp
  &disper
    nlay= 3,
rho= 0.1600E+04, 0.1600E+04, 0.1700E+04,
mu= 0.1600E+08, 0.1600E+08, 0.1530E+09,
lame= 0.9920E+09, 0.9920E+09, 0.3519E+10,
zi= 0.0000E+00, 0.1000E+01, 0.1010E+01,
deltz= 1.0000,
modemx=9,
nfreq=202, flo= 0.1000000E+01, delf= 0.48828122E+00, jsmax=300, ksw=0
  pvlcty=0.0, pfreq=0.0, zend=100.0,
  ofile='disper.tmp',
  octav1='phase.m', octav2='mat2.m',
  curve='earth.crv', /
    1      0.000      100.00      800.00      1600.0
    2      1.000      100.00      800.00      1600.0
    3      1.010      300.00      1500.00      1700.0

```

### 9.2.6 DISPER

After **gendis** ( 9.2.5 ) is run, a namelist file can be run to compute dispersion. The output includes a text file, **disper.tmp**, a data file capturing dispersion, **earth.crv**, and some Octave files, **model.m** and **phase.m**.

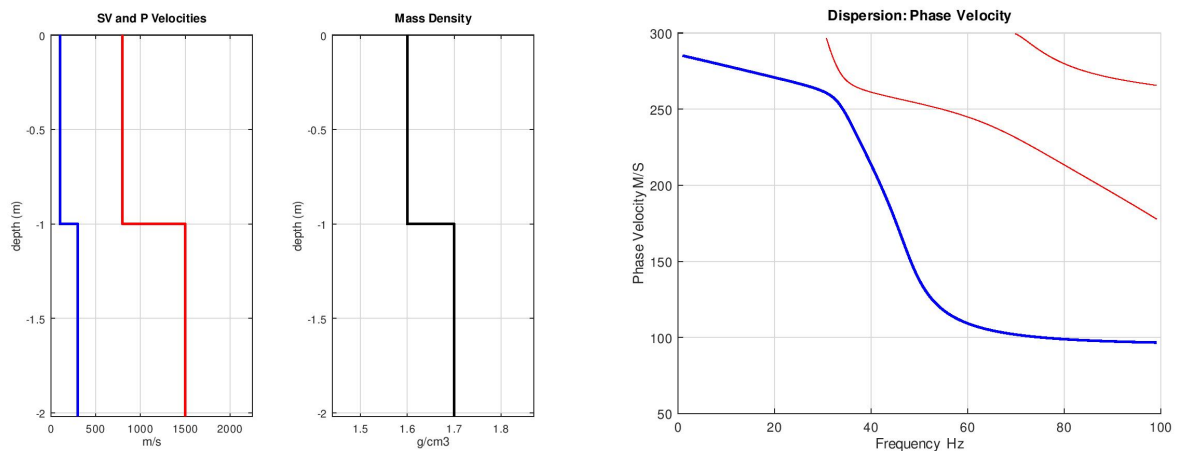


Figure 50: DISPER: The model and phase velocity plots. The model.m plot shows P-wave (red), S-wave (blue) velocity, and density (black). This is a layer over a half-space model. On the right is the phase.m generated plot showing the fundamental mode (blue) and two higher modes (red). The model (soil profile) was generated in **gendis** (9.2.5)

The computational function is the same as **rwv.f** used in the octave programs **FwdR1.m** 9.2.1 and **invR1** 8.1.1.

**9.2.6.1 Motion-Stress from disper.d** The file, **disper.d** can be edited and **disper** run in a different way, computing the motion-stress vector for a given frequency and mode. For example, from the file **disper.tmp**, note the phase velocity for a particular mode. Pick a frequency of interest, say 32.2265605 Hz. We scroll down **disper.tmp**:

```

31.2499981 | 259.6905012 291.4923132
31.7382793 | 258.6609733 286.9434926
32.2265605 | 257.4416940 282.8358906
32.7148417 | 255.9845241 279.2012189
33.2031230 | 254.2462253 276.0619743

```

We edit the **pvlcty=** line of **disper.d** to use the phase velocity. For the fundamental mode, the replacement is:

```

< pvlcty=0.0, pfreq=0.0, zend=100.0,
---
> pvlcty=257.4416940, pfreq=32.2265605, zend=10.0,

```

We also change **zend**, choosing a more relevant depth of interest which depends on the frequency. Low frequencies extend deeper than high frequencies. We also change the computational depth interval, **deltz**. The replacement is:

```

< deltz= 1.0000,
---
> deltz= .0010,

```

We change the computational increment from 1 meter to .001 meters. Review the relevant lines of the **disper.d** file shown above in section 9.2.5 for example. Running **disper** with the edited **disper.d** file will replace **disper.tmp** with a new version listing the motion-stress vector computed at the new **deltz** interval. Figure 51 is the result.

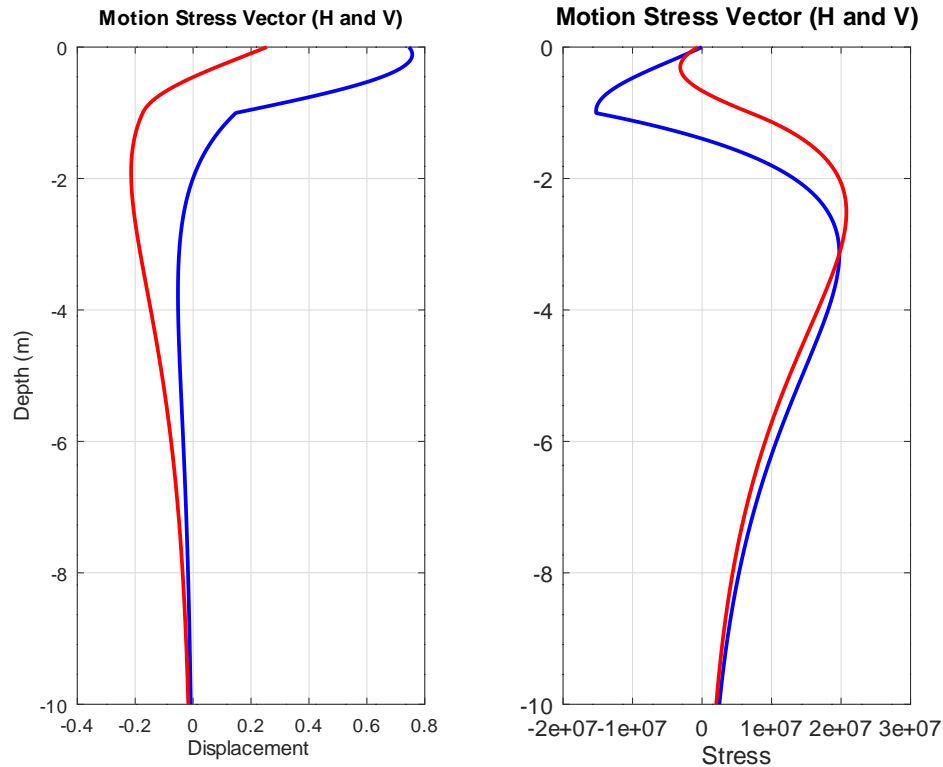


Figure 51: DISPER: Re-running **disper** to compute the motion stress vectors. See section 9.2.6.1 for how to edit **disper.d**. The file, **mat2.m** created this plot. Blue is horizontal, Red is vertical.

### 9.2.7 GENWAV

The output file **earth.crv** generated by DISPER ( 9.2.6) can be used as input to the program **WAVES** which will compute a synthetic Rayleigh wave only seismic data profile. This program, **genwav**, is used to interactively build a namelist file for program **WAVES** (see 9.2.8).

**IMPORTANT: The *tmax* and *sample interval* must agree when running *gendis* and *genwav* so that frequency sampling matches.** The following is a log of an example run of this interactive program.

```
genwav
  enter name of namelist file (40 char)
  Example: waves.d
waves.d
  enter name of dispersion curve file
  (this is file from disper.f)
  Example: earth.crv
earth.crv
  enter near offset:  xnear
1
  enter group interval:  delx
1
  enter number of receivers:  nrec
24
  enter minimum group velocity expected
100
  RECOMMENDED minimum tmax=   0.4800
  enter: maximum trace time, tmax
1.0
  enter: sample interval (seconds), fsamin
.001
  enter frequencies: fmin, fmax
1., 100.
```

```

enter maximum mode to include
9
enter ksw switch 0=c plot, 1=k plot
0
enter type of plot format, mapmat
0=octave (Matlab) 1=scilab
0
enter Output option 0=Vertical 1=Radial
0
enter source depth
0
enter (3) diagonal elements, moment tensor
0,0,1
Padded Radix 2 tmax= 2.0480
Number of points in signal= 2048
-----
.....Frequency interval= 0.48828122
NOTE: Frequency Interval MUST MATCH DISPER OUTPUT
WAVES will output signal length = 1.0/delf
IF MISMATCHED: CHANGE sample rate in WAVES
or RERUN DISPER
-----
Number of frequencies= 204
output in =====>waves.d

```

**9.2.7.1 Frequency Increment** Note that following the RECOMMENDED minimum  $t_{max}$ , the next two questions must agree with the **gendis** run, specifically enter **tmax** (the maximum recording time) and **fsamin** (the sample interval in seconds). The “Frequency interval” must be the same for both **disper** (which generates **earth.crv**) and the intended **waves** program run. If unsure, open the **earth.crv** file in your favorite editor, and compute the difference between consecutive frequencies (column 1). This frequency interval must match the one near the end of the **genwav** run (written to the terminal, see above for example). In this example:

```
0.1464843660000000102883178E+01 -0.9765624400000000315813509E+00 = 0.48828
```

### 9.2.7.2 Explanation of **genwav** parameters

- **xnear** near offset in meters
- **delx** spacing between channels in meters
- **nrec** number of receivers (ie. traces in shot gather)
- **minimum group velocity expected** used in estimating needed record length.
- **tmax** length of traces in seconds (need not match the recommended, depending on **gendis** run).
- **fsamin** sample interval in time, seconds
- **fmin, fmax** minimum and maximum frequencies. Recommend that these be wider than what you think you need, then filter back for your final result using a filter program (like **BFIL**). NOTE: Wavelet used is minimum phase, set by **fmin** and **fmax**.
- **maximum mode to include** Must be  $modemax \leq 9$
- **ksw** dispersion plot, sets wavenumber (1=k) or velocity (0=c)
- **mapmat** Format for plots. Recommend Octave (Matlab)
- **output option, irvsel** signals will be vertical or inline radial.
- **source depth** depth of source in meters
- **moment tensor diagonal** (radial,transverse,vertical). 0,0,1 is a vertical impulse. You can edit the **waves.d** file if you want a double couple instead.

The waves.d listing for this example:

```
&waves
ksw= 0, stepz=20,
modes=1,2,3,4,5,6,7,8,9,
fmin= 1.0000, fmax= 100.0000,
fsamin= 0.00100,
curve='earth.crv',
mapmat=0,
matlb1='matc.m', scilb1='matc.sci',
matlb2='matu.m', scilb2='matu.sci',
irvsel=0,
ofile='waves.tmp', /
&source
tm= 0.0, 0.0, 0.0,
    0.0, 0.0, 0.0,
    0.0, 0.0, 1.0, /
sz= 0.00, sy=0.00, sx=0.00, /
&recvr
nrec=24,
rz=24*0.0,
ry=24*0.0,
rx= 1.000, 2.000, 3.000, 4.000, 5.000,
    6.000, 7.000, 8.000, 9.000, 10.000,
    11.000, 12.000, 13.000, 14.000, 15.000,
    16.000, 17.000, 18.000, 19.000, 20.000,
    21.000, 22.000, 23.000, 24.000,
/
```

This file can be edited in case the **genwav** options don't cover what you want. If you want only the fundamental mode, for example, change the modes line:

```
modes=1,0,0,0,0,0,0,0,0,
```

The **irvsel** parameter is an easy way to change between **vertical** or **horizontal radial** signals on the receivers. For guidance on the moment tensor, **tmll**, see [Aki & Richards \(1980\)](#)

### 9.2.8 WAVES

WAVES computes elastic Rayleigh waves. Start with **GENDIS 9.2.5**, then run **DISPER 9.2.6**. Run **GENWAV 9.2.7** to define a simulation geometry and parameters. IMPORTANT: Make sure the frequency increments are consistent between **disper** and **waves** (see [9.2.7.1](#)). The **waves.d** file will then be input to **WAVES**. Outputs from WAVES include:

- **matu.m** Octave program to plot group velocity dispersion
- **matc.m** Octave program to plot phase velocity dispersion (redundant with phase.m output from **disper**).
- **m0.m** Octave program to plot wavelet and spectrum.
- **wavV.seg** or **wavR.seg** seismic shot gathers in BSEGY format.
- **waves.tmp** listing file for waves run.
- **waves.his** scaled lagrangian maximum for all runs made in the current directory. Smaller the better since integrating stiff equations.

The following are some plots generated from these outputs.

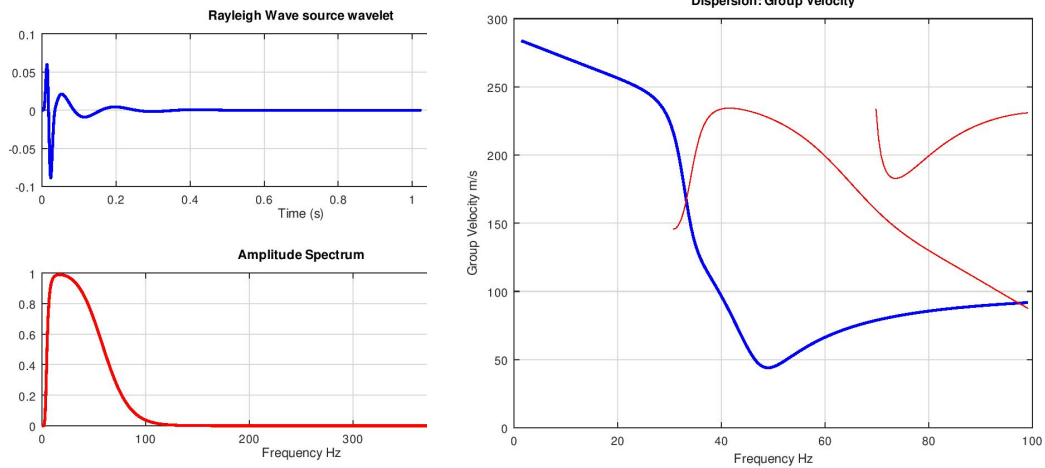


Figure 52: WAVES: Wavelet on left, group velocity dispersion on right. No significance to curve colors except that in the dispersion plot, the fundamental is Blue and higher modes are in Red. Soil representation is layer over half-space as shown in 9.2.5.1 and Figure 50 above.

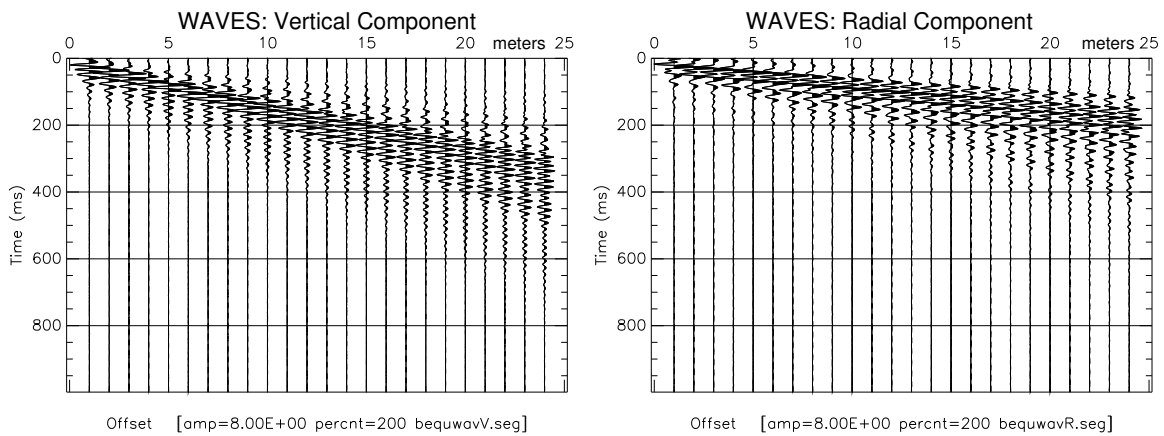


Figure 53: WAVES: Synthetic seismograms for Vertical (wavV.seg) and horizontal (wavR.seg) motion.

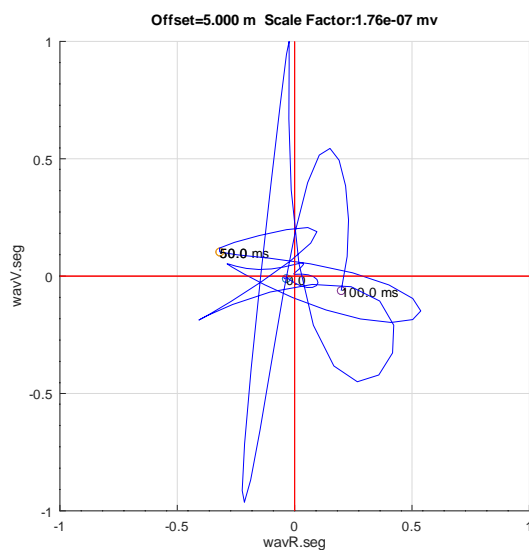


Figure 54: Hodogram for offset 5 meters. Requires bbegin.m, segyinfo.m, and hodo2plot.m in directory with wavV.seg and wavR.seg files (see 6.0.10).

While Rayleigh waves are often described to have elliptical retrograde motion, that is not always the case. Depending on the source depth, receiver depth, and the geologic profile, the motion can be either prograde or retrograde. In Figure 54 we see that the motion is complex, starting out with an ellipse with a sub-horizontal major axis, evolving to a vertical major axis of elliptical motion. An alternative case that illustrates retrograde elliptical motion is computed for the same waves.d. The revised disper.d uses a homogeneous half-space model (two points to describe). The result is shown in Figure 55:

```
show.tmp
&disper
  nlay= 2,
rho= 0.1700E+04, 0.1700E+04,
mu= 0.6800E+08, 0.6800E+08,
lame= 0.9520E+09, 0.9520E+09,
zi= 0.0000E+00, 0.1000E+01,
deltz= 1.0000,
modemx=1,
nfreq=202, flo= 0.1000000E+01, delf= 0.48828122E+00, jsmax=300, ksw=0
pvlcty=0.0, pfreq=0.0, zend=100.0,
ofile='disper.tmp',
octav1='phase.m', octav2='mat2.m',
curve='earth.crv', /
  1      0.000      200.00      800.00      1700.0
  2      1.000      200.00      800.00      1700.0
```

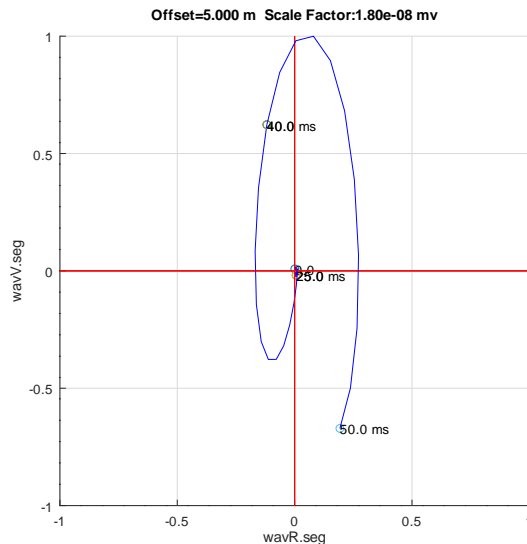


Figure 55: Hodogram at offset 5 meters for alternative half-space soil model, see show.tmp above. Sign conventions need to be taken into account when determining type of motion.

### 9.2.9 BDUM

This program reads a \*.seg file, adopts the headers and replaces the data with an impulse at a user specified time. It can be used to present an impulse response or to benchmark software and do testing. In the following example, the filter program is illustrated. See Figure 56.

```
bdum infile time_impulse
```

```
EXAMPLE: filter an impulse at 0.1 seconds, headers from c008.seg
bdum c008.seg .1
bfil bdumc008.seg 1 6 40 40 1
```



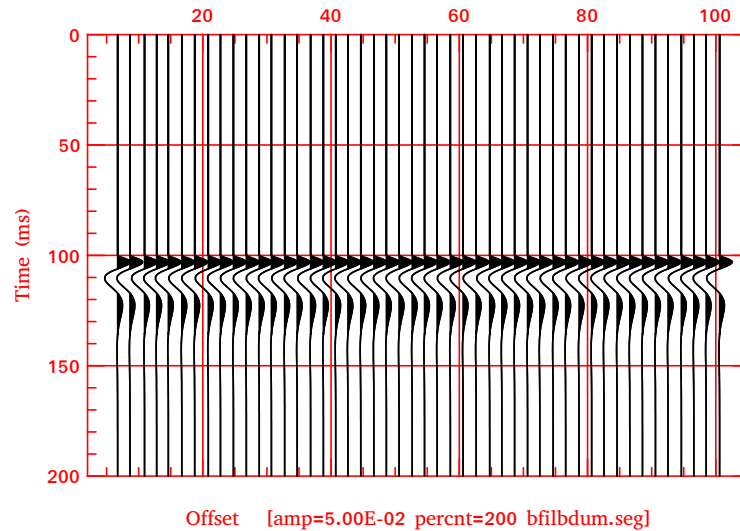


Figure 56: BDUM: Impuse replaced original data and filtered by BFIL program (band-pass 6 pole 40 Hz center, 40 Hz bandwidth, minimum phase).

### 9.2.10 OCTAVE rayleigh.m

Demonstration program on how to use the dispersion computation function `rwv.f` in an octave program. Required octave functions are:

- `rwv.f` computes dispersion, requires compilation and linking into the octave engine.
- `wrapper.cpp` wrapper code
- `build_disper_oct` script to compile `rwv.f` and build `disper.oct`
- `rayleigh.m` the demonstration code.

The build script has one active line plus some informational echo commands. The active line is:

```
mkcoctfile rwv.f wrapper.cpp -o disper.oct
```

In order for this to work, one must have the development package installed for Octave. In Debian 10 Linux, the packages needed installation are:

- `liboctave-dev`
- `octave-common`

Of course, these are not the only packages, Octave has a lot of packages that are of use. But the above packages will install the `mkcoctfile`. The `wrapper.cpp` and all BSU `*.m` files will be installed in `/usr/local/share/octave/site-m/` if you build BSU from the source tar archive. The `locate` command can also be helpful in other situations.

The `disper()` function returns a vector `pv` with fundamental and any higher modes. Here, it is demonstrated how to select and plot both fundamental and first higher mode. The higher mode becomes possible and recognized when the returned `pv(2)` value becomes  $> 0$ . This demonstration code searches backward to find the first non-zero case of the second component being non-zero.

Another similar code, `moho.m` is included that illustrates the same points as above. It shows how to display dispersion as a function of period rather than frequency.

## 10 Surveying, Setting Geometry, and Mapping

Setting geometry is the act of creating headers that include the locations of seismic sources and geophones. BSU includes programs for setting geometry as well as making maps and computer aided drafting (CAD) files from headers once they are set.

- **GENWAW 10.1.1** Labor intensive interactive conversion of SEG-2 (\*.DAT files) and setting geometry for each source and receiver.
- **GENREF 10.1.2** interactive, generates bash scripts for setting geometry on CDP reflection shooting. Bison data ONLY.
- **TOPCON 10.1.3** reads a survey \*.nez file and **Bison** seismograph file, creates \*.xyz file.
- **BHED 10.1.4** down-load or up-load header data from or to an \*.seg file.
- **TOPCON2 10.1.5** converts **SEG-2** (\*.DAT) to \*.seg format while setting geometry from command line arguments. **GENVSP 10.1.8** can be used first to set up bash scripts that use a \*.nez file and calls to TOPCON2.
- **GENSETG 10.1.6** interactive program creates files for setting geometry where phones fixed, shots move (reciprocal shooting)
- **SETGEOM 10.1.7** Run after GENSETG, takes the shot.txt, phones.txt, and \*.nez files created by GENSETG and applies them. A \*.nez file is Northing, Easting, Elevation text file.
- **GENVSP 10.1.8** interactive program for setting geometry in down-hole surveys.
- **GENBHOD 10.1.9** SH-wave source interactive program generates bash scripts to determine down-hole tool orientation by principle component analysis (PCA) of shot records. Program **BHOD 10.1.11** does the actual PCA.
- **GENBHODV 10.1.10** Vertical impact source interactive program generates bash scripts to determine down-hole tool orientation by PCA. Experimental, uses Rayleigh wave on horizontal component. Program **BHOD 10.1.11** does the actual PCA.
- **BHOD 10.1.11** performs PCA on down-hole data.
- **BNEZ 10.1.12** generates a \*.nez file from rules. Typically run twice, once for shots, once for geophones.
- **TOP2NEZ 10.1.13** converts a raw Topcon Total Station survey file to NEZ format.
- **TOP2DXF 10.1.14** reads a \*.nez file and converts it to a \*.dxf (CAD) file.
- **TOPBCRD 10.1.15** applies rotation and translations to coordinates in an \*.nez file. Program **BCRD 10.1.16** does this on \*.seg files.
- **BCRD 10.1.16** rotates and translates header geometry coordinates in an \*.seg file.
- **BCAD 10.1.17** creates a CAD \*.dxf file from \*.seg file headers.

## 10.1 Setting Geometry

These are interactive codes for setting geometry. They are run from a terminal with a question and answer format.

### 10.1.1 GENWAW

Basic Seismic Utilities (BSU) interactive program for setting geometry. Code optimal for a walk-a-way type of data collection. Code is for SEG-2 format (\*.DAT) files. This code prompts the user for shot and geophone locations. It should be run in the directory where the SEG-2 \*.DAT files are located. The code scans the directory contents and builds a list of the files needing to have headers corrected for geometry. One use that makes geometry setting less of a burden is to set geometry for temporary local coordinates (ex. line along x-axis), then employ BSU program BCRD to rotate and translate coordinates to a final system. This program is run if the SEG-2 headers were not correctly set during acquisition (a common occurrence).

#### EXAMPLE:

Here, the application is a walk-a-way with a fixed source at 0,0,0 and a single moving vertical component geophone, starting at an offset of 10 meters walking in 1 meter steps toward the source. **This is labor intensive, but very flexible.**

```

genwaw
Enter Number of Channels 1
GEOPHONE ORIENTATIONS
Geophone Az 90=East Ver 180=Down
Channel=0 Enter Geophone Orientation Az Ver
0 180
|-----|
|   Copyright (C) 2017 P. Michaels   |
|         All rights reserved         |
See GNU General Public License
waw: TIME: 15:38:11 DATE: 29/May/2020
Nsources= 10
SOURCE LOCATION-----
0000.DAT: Enter Source X Y Z
0 0 0
Trace:00 Enter Receiver X Y Z
10. 0 0
SOURCE LOCATION-----
0001.DAT: Enter Source X Y Z
0 0 0
Trace:00 Enter Receiver X Y Z
9 0 0
SOURCE LOCATION-----
0002.DAT: Enter Source X Y Z
0 0 0
Trace:00 Enter Receiver X Y Z
8. 0 0
.
. etc....
.
SOURCE LOCATION-----
0009.DAT: Enter Source X Y Z
0 0 0
Trace:00 Enter Receiver X Y Z
1.0 0 0

```

NOTE: If there were more than one trace in each \*.DAT file, there would be additional “Trace:” questions to answer.

Output includes the creation of a child directory, LST, in which the list files for each \*.DAT file are stored. These are the result of the interactive program calling EGG2SEG 3.1.6. In this example, there will be files 0000.seg through 0009.seg, each with a single trace, now with headers as entered in GENWAW. To merge these into a single file, use BMRG (see 11.0.1). The command would be

bmrgr 000 0 9 1 1 1

and that produces a file bmrgr.seg. The headers can be checked by running BDUMP 4.0.1:

```
-----
Length = 500 samples          | Shot Elevation = 0.0
Sample Interval = 0.00100 sec. | Shot Depth = 0.0
Delay Time = 0 msec.         | Up Hole Time = 0 msec
Low Cut Filter = 0 Hz.       | Shot X-COORD = 0.00
High Cut Filter = 100 Hz.    | Shot Y-COORD = 0.00
Line ID: 000                 | Shot Date (year.moday) = 2019.0423
Shot Orientation:            | Shot Time (hr:min) = 15:20
Azimuth= 0 Deg. Vertical= 0 Deg. | Charge Size (grams)= 0
-----
```

| TRACE # | SHOT REC | STATION SHOT REC | OFFSET | ELEV. | X-COORD | Y-COORD | VERT | 1STBRK | K-GAIN | AZI | VER |
|---------|----------|------------------|--------|-------|---------|---------|------|--------|--------|-----|-----|
| 1       | 0        | 0000 0001        | 10.00  | 0.00  | 10.00   | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 2       | 1        | 0000 0001        | 9.00   | 0.00  | 9.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 3       | 2        | 0000 0001        | 8.00   | 0.00  | 8.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 4       | 3        | 0000 0001        | 7.00   | 0.00  | 7.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 5       | 4        | 0000 0001        | 6.00   | 0.00  | 6.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 6       | 5        | 0000 0001        | 5.00   | 0.00  | 5.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 7       | 6        | 0000 0001        | 4.00   | 0.00  | 4.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 8       | 7        | 0000 0001        | 3.00   | 0.00  | 3.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 9       | 8        | 0000 0001        | 2.00   | 0.00  | 2.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |
| 10      | 9        | 0000 0001        | 1.00   | 0.00  | 1.00    | 0.00    | 1    | 0.0000 | 19     | 0   | 180 |

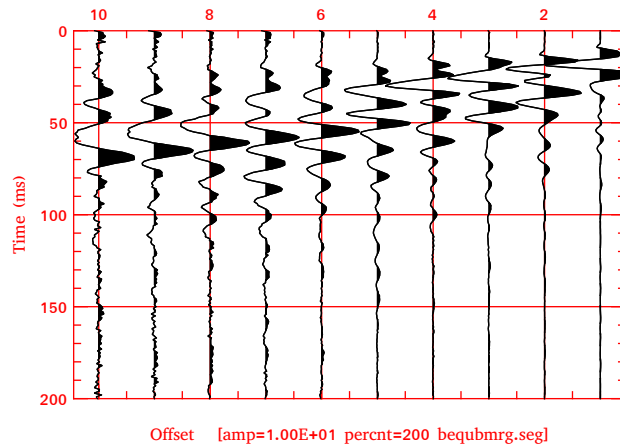


Figure 57: GENWAW: Example data from a small hammer source, trace equalized with BEQU 12.0.9.

### 10.1.2 GENREF

NOTE: This program is only for BISON data. For SEG-2 formatted data, consider GENSETG 10.1.6 and SET-GEOM 10.1.7. This is an interactive program that generates bash scripts for setting geometry when doing conventional “Roll-a-long” shooting, but no survey data are available. The program generates an \*.nez file and scripts:

- **geom** run this first, it runs program TOPCON to generate \*.xyz files.
- **geom2** run this second, it calls script **go1** which reads the \*.xyz files
- **go1** called by geom2

TIP: Make sure to chmod +x the scripts to make them executable. The following is an example log of a small run:

```

genref
|-----|
| Copyright (C) 2017 P. Michaels |
|   All rights reserved         |
see GNU General Public License
CDP Roll-a-long Pattern Generator
Bison Format Data
-----SOURCES-----
  Enter 6 char. name for nez file (ex. STP001)
ABC001
  Enter 4 char. LINEID
0001
  Enter Z-Datum: Elevation
100
  Enter number of shots
5
  Enter Shot Record Names 8char: First
FAC10001
  Enter Shot Record Names 8char: Last
FAC10005
  Enter First Shot Station Number
1
  Enter First Source: x, y, z
0,0,100
  Enter Last Source: x, y, z
5,0,105
  Enter number of receivers in a shot gather
24
  Enter TOTAL NUMBER of stations on line
48
  Enter First Geophone Station: x, y, z
0,0,100
  Enter Last Geophone Station: x, y, z
48,0,109
  Enter first shot NEAR GEOPHONE station
0,0,100
  Enter first shot FAR GEOPHONE station
24,0,105

```

The ABC001.nez (index, Northing, Easting, Elev) file looks like this:

```

0001      0.0000      0.0000      200.0000  SP001
0002      0.0000      1.2500      201.2500  SP002
0003      0.0000      2.5000      202.5000  SP003
0004      0.0000      3.7500      203.7500  SP004
0005      0.0000      5.0000      205.0000  SP005
0001      0.0000      0.0000      200.0000  VP001
0002      0.0000      1.0213      200.1915  VP002
0003      0.0000      2.0426      200.3830  VP003
0004      0.0000      3.0638      200.5745  VP004
0005      0.0000      4.0851      200.7660  VP005
0006      0.0000      5.1064      200.9574  VP006
.
. etc...
.
0041      0.0000      40.8511      207.6596  VP041
0042      0.0000      41.8723      207.8511  VP042
0043      0.0000      42.8936      208.0426  VP043
0044      0.0000      43.9149      208.2340  VP044
0045      0.0000      44.9362      208.4255  VP045
0046      0.0000      45.9574      208.6170  VP046
0047      0.0000      46.9787      208.8085  VP047
0048      0.0000      48.0000      209.0000  VP048

```

The labels “SP” are shot locations, the labels “VP” are voltage points (geophone) locations.

The **geom** file calls **topcon 10.1.3** for the **Bison** files FAC10001 etc., and looks like this:

```
topcon ABC001.nez FAC10001 0001 0.0 1 24 000 023 1 0. 0 0 0 0
topcon ABC001.nez FAC10002 0001 0.0 2 24 001 024 2 0. 0 0 0 0
topcon ABC001.nez FAC10003 0001 0.0 3 24 002 025 3 0. 0 0 0 0
topcon ABC001.nez FAC10004 0001 0.0 4 24 003 026 4 0. 0 0 0 0
topcon ABC001.nez FAC10005 0001 0.0 5 24 004 027 5 0. 0 0 0 0
```

The **geom2** script looks like this:

```
go1 001
go1 002
go1 003
go1 004
go1 005
```

The **go1** script looks like this:

```
bis2seg FAC10$1
bhed FAC10$1.seg FAC10$1.xyz 0
mv bhedFAC1.seg F$1.seg
rm FAC10$1.seg
```

### 10.1.3 TOPCON

For **BISON** data. The program combines \*.nez survey file (Northing, Easting, Elevation) and Bison files from a Bison seismograph to produce \*.xyz header files. The program **BHED 10.1.4** then reads the \*.xyz files and uploads them into the \*.seg files. The command line arguments are:

```
topcon topf bisf lid shdp is nch vp1 vpn ir esh isa isv ira ita
```

```
topf      =file name of topcon .nez file
bisf      =file name of Bison file with data
lid       =line ID
shdp      =shot depth
is        =shot location number
nch       =number of channels (<48)
vp1       =geophone location number channel 1
vpn       =geophone location number channel nchan1
ir        =shot record number
esh       =elevation adjustment to be added
isa       =source polarization azimuth (deg.)
isv       =source polarization vertical (deg.)
ira       =reference polarization R-axis (deg.)
ita       =reference polarization T-axis (deg.)
```

### 10.1.4 BHED

**BHED** either uploads or downloads header data into/from \*.seg files. The command line arguments are:

```
bhed infile header_file iupdn
```

```
infile      =input file name
header_file =file with selected header info
iupdn       =1 download headers to header_file
            =0 upload headers to BSEGY data set
```

Aside from initial upload of headers, one can also use this program to edit existing headers. Just download to a header file from an existing \*.seg headers, open the header file and edit. **HINT: Watch out for zeros. In particular, note that a binary zero is used to terminate character strings. Depending on how initial headers were set, it is possible that a header string might have a binary zero, often shown as a @ symbol in an editor like VI.**

As a sample, the top of a header file looks like this:

```

&BHED
LOWCUT=8      ,
HIGHCT=500   ,
LINE="4N__",
YEAR=1994    ,
DAY=1117     ,
HOUR=11      ,
MINUTE=46    ,
PHONE="VERT",
SDEPTH= 0.400000006 ,
UPHOLE= 0.00000000  ,
CHARGE=0     ,
SREC=8       ,
/
 1 0.0000    003 9668.130 10131.190 818.700    001 9670.780 10125.040 818.840  0 000 180 000 000
 2 0.0000    003 9668.130 10131.190 818.700    002 9671.670 10123.330 818.860  0 000 180 000 000
 3 0.0000    003 9668.130 10131.190 818.700    003 9671.120 10120.710 818.840 20 000 180 000 000
 4 0.0000    003 9668.130 10131.190 818.700    004 9673.280 10119.480 818.830 20 000 180 000 000
 5 0.0000    003 9668.130 10131.190 818.700    005 9674.080 10117.840 818.760 20 000 180 000 000
 6 0.0000    003 9668.130 10131.190 818.700    006 9674.990 10115.940 818.670 40 000 180 000 000
 7 0.0000    003 9668.130 10131.190 818.700    007 9675.950 10114.170 818.710 40 000 180 000 000
 8 0.0000    003 9668.130 10131.190 818.700    008 9676.910 10112.280 818.790 40 000 180 000 000
 9 0.0000    003 9668.130 10131.190 818.700    009 9677.660 10110.530 818.720 40 000 180 000 000
10 0.0000    003 9668.130 10131.190 818.700    010 9678.490 10108.690 818.720 40 000 180 000 000
11 0.0000    003 9668.130 10131.190 818.700    011 9679.280 10106.840 818.720 40 000 180 000 000
etc .....
42 0.0000    003 9668.130 10131.190 818.700    042 9705.330 10050.830 819.400 60 000 180 000 000
43 0.0000    003 9668.130 10131.190 818.700    043 9706.300 10049.040 819.400 60 000 180 000 000
44 0.0000    003 9668.130 10131.190 818.700    044 9707.120 10047.260 819.390 60 000 180 000 000
45 0.0000    003 9668.130 10131.190 818.700    045 9708.000 10045.380 819.420 60 000 180 000 000
46 0.0000    003 9668.130 10131.190 818.700    046 9708.910 10043.660 819.410 60 000 180 000 000
47 0.0000    003 9668.130 10131.190 818.700    047 9709.710 10041.910 819.380 60 000 180 000 000
48 0.0000    003 9668.130 10131.190 818.700    048 9710.460 10039.950 819.480 60 000 180 000 000

```

These are read as namelist files by BHED. The above was created by the command:

```
bhed c008.seg header.txt 1
```

and the header file was created with the name header.txt.

### 10.1.5 TOPCON2

For **SEG-2** data. The program combines \*.nez survey file data with the SEG-2 seismic file data to produce a BSEGY format file, \*.seg. One difference between this and the Bison data **TOPCON 10.1.3** procedure is that there is no need to run **BHED** with an intermediate \*.xyz file. This goes directly to \*.seg. The command line arguments are:

```
topcon2 topf seg2f lid shdp is nch vpl vpn ir esh isa isv ira ita
```

```

topf   = topcon file name
seg2f  = seg-2 file name
lid    = line ID
shdp   = shot depth
is     = shot location number
nch    = number of channels (nch<66)
vpl    = geophone station channel 1
vpn    = geophone station channel n
ir     = shot record number
esh    = elevation adjustment to be added
isa    = source polarization azimuth (deg.)
isv    = source polarization vertical (deg.)
ira    = reference phone polarization R-axis (deg.)
ita    = reference phone polarization T-axis (deg.)

```

An example of issuing the command for a down-hole surge:

```
topcon2 stp001.nez 1051.DAT 00X5 0.0 1 6 0156 0151 1051 0. 270 135 0 270
```

This combines survey file stp001.nez with SEG-2 data 1051.DAT. In the case of down-hole data, the only orientation of horizontal components is known for the reference phone. To determine the orientation of a down-hole phone see **GENBHOD 10.1.9** and **BHOD 10.1.11**. A partial listing of the resulting header dump by program **BDUMP 4.0.1** follows:

```

|-----|
| PARTIAL SEG2 HEADER DUMP |
|                               |
|           1051.seg           |

```

```

-----
Length = 2000 samples          | Shot Elevation =      849.2
Sample Interval = 0.00025 sec. | Shot Depth =         0.0
Delay Time = 0 msec.          | Up Hole Time =       0 msec
Low Cut Filter = 0 Hz.        | Shot X-COORD =    9963.09
High Cut Filter = 1000 Hz.    | Shot Y-COORD =    10022.70
Line ID: 00X5                 | Shot Date (year.moday) = 1999.1102
Shot Orientation:             | Shot Time (hr:min)  = 14:25
Azimuth=270 Deg. Vertical=135 Deg. | Charge Size (grams)= 0
-----
TRACE|SHOT| STATION | OFFSET|          RECEIVER          |VERT|1STBRK|K-GAIN|AZI|VER|
#	REC.	SHOT REC		ELEV. X-COORD  Y-COORD	FOLD	(SEC.)	(dB)		
1 | 0|0001 0156| 0.73| 848.96  9963.09  10022.00| 3|0.0000| 0 |270| 90|
2 | 0|0001 0155| 0.73| 848.96  9963.09  10022.00| 3|0.0000| 0 | 0| 90|
3 | 0|0001 0154| 0.73| 848.96  9963.09  10022.00| 3|0.0000| 24 | 0| 0|
4 | 0|0001 0153| 14.49| 834.68  9963.09  10023.25| 3|0.0000| 24 | 0| 90|
5 | 0|0001 0152| 14.49| 834.68  9963.09  10023.25| 3|0.0000| 24 | 0| 90|
6 | 0|0001 0151| 14.49| 834.68  9963.09  10023.25| 3|0.0000| 24 | 0| 0|

```

### 10.1.6 GENSETG

This program sets up files for a second program **SETGEOM 10.1.7** which does the actual setting of geometry for SEG2 data. A primary application is reciprocal refraction shooting where blocks of geophones are irregularly placed on banks of a river. Given the flexible nature of this pair of programs, it can be useful for other applications as well. This is an interactive program, and produces two text files, one for shots, one for geophones. An example log of a run is shown below:

```

gensetg

SHOTS: -----
Enter first shot file NAME number
1001
Enter last shot file NAME number
1004
Enter first SP label NUMBER
01
Enter increment for SP label NUMBER
01
PHONES: -----
Enter number of BLOCKS to define channels
2
BLOCK Number----- 1
Channels (1) through (?)
Enter last channel for this block
12
Enter first label VP NUMBER for this block
01
Enter label VP increment for this block
01
BLOCK Number----- 2
Channels (13) through (?)
Enter last channel for this block
24
Enter first label VP NUMBER for this block

```



50

Enter label VP increment for this block

1

The two files output are **shots.txt** and **phones.txt**. The shots.txt file contains the following:

```
1001.seg SP001
1002.seg SP002
1003.seg SP003
1004.seg SP004
```

The phones.txt file contains the following:

```
01 VP001
02 VP002
03 VP003
04 VP004
05 VP005
06 VP006
07 VP007
08 VP008
09 VP009
10 VP010
11 VP011
12 VP012
13 VP050
14 VP051
15 VP052
16 VP053
17 VP054
18 VP055
19 VP056
20 VP057
21 VP058
22 VP059
23 VP060
24 VP061
```

These SP and VP labels would correspond to those in an \*.nez file produced by a surveying instrument. This example might correspond to channels 1–12 being on one bank of a river, then a jumper cable might cross the river and connect to channels 13–24 with geophones on the other bank of the river. The 4 shots might then be taken with an airgun deployed from the bridge. In reality, there would likely be more shots than 4, but this illustrates the concept.

### 10.1.7 SETGEOM

After running **GENSETG 10.1.6**, one needs to also have a survey \*.nez file before proceeding. Continuing the example started in **10.1.6**, this might look like this:

```
1 0.000000 0.000000 100.000000 SP001
2 0.000000 2.000000 101.000000 SP002
3 0.000000 4.000000 102.000000 SP003
4 0.000000 6.000000 103.000000 SP004
5 2.000000 2.000000 100.000000 VP001
6 3.000000 2.000000 100.000000 VP002
7 4.000000 2.000000 100.000000 VP003
8 5.000000 2.000000 100.000000 VP004
9 6.000000 2.000000 100.000000 VP005
10 7.000000 2.000000 100.000000 VP006
11 8.000000 2.000000 100.000000 VP007
12 9.000000 2.000000 100.000000 VP008
13 10.000000 2.000000 100.000000 VP009
14 11.000000 2.000000 100.000000 VP010
15 12.000000 2.000000 100.000000 VP011
16 13.000000 2.000000 100.000000 VP012
17 2.000000 8.000000 125.000000 VP050
18 3.000000 8.000000 125.000000 VP051
19 4.000000 8.000000 125.000000 VP052
```

```

20 5.000000 8.000000 125.000000 VP053
21 6.000000 8.000000 125.000000 VP054
22 7.000000 8.000000 125.000000 VP055
23 8.000000 8.000000 125.000000 VP056
24 9.000000 8.000000 125.000000 VP057
25 10.000000 8.000000 125.000000 VP058
26 11.000000 8.000000 125.000000 VP059
27 12.000000 8.000000 125.000000 VP060
28 13.000000 8.000000 125.000000 VP061
    
```

The \*.nez file contains the (N,E,Z) coordinates and must include the SP and VP labels that match the shots.txt and phones.txt files. If the SEG2 data files were converted to BSEGY format with **EGG2SEG 3.1.6** we might have files 1001.seg through 1004.seg in our directory. We would then run setgeom with the following command:

```
setgeom shots.txt phones.txt samp0000.nez
```

where it is assumed that the \*.nez file is as shown here. The output BSEGY files would be setg1001.seg through setg1004.seg. The header dump using **BDUMP 4.0.1** of file setg1001.seg would look like this:

```

-----
Length = 1000 samples          | Shot Elevation = 100.0
Sample Interval = 0.00050 sec. | Shot Depth = 0.0
Delay Time = 0 msec.          | Up Hole Time = 0 msec
Low Cut Filter = 8 Hz.        | Shot X-COORD = 0.00
High Cut Filter = 500 Hz.     | Shot Y-COORD = 0.00
Line ID: 001~@               | Shot Date (year.moday) = 1994.1117
Shot Orientation:            | Shot Time (hr:min) = 11:46
Azimuth= 0 Deg. Vertical=180 Deg. | Charge Size (grams)= 0
-----
TRACE|SHOT| STATION | OFFSET|          RECEIVER          |VERT|1STBRK|K-GAIN|AZI|VER|
# |REC.|SHOT REC|      | ELEV. X-COORD  Y-COORD|FOLD|(SEC.)| (dB) |  |  |
-----|-----|-----|-----|-----|-----|-----|-----|-----|-----|-----|
 1 | 1| 001 001| 2.83| 100.00  2.00  2.00|1|0.0000| 0 | 0|180|
 2 | 1| 001 002| 3.61| 100.00  2.00  3.00|1|0.0000| 0 | 0|180|
 3 | 1| 001 003| 4.47| 100.00  2.00  4.00|1|0.0000| 20 | 0|180|
 4 | 1| 001 004| 5.39| 100.00  2.00  5.00|1|0.0000| 20 | 0|180|
 5 | 1| 001 005| 6.32| 100.00  2.00  6.00|1|0.0000| 20 | 0|180|
 6 | 1| 001 006| 7.28| 100.00  2.00  7.00|1|0.0000| 40 | 0|180|
 7 | 1| 001 007| 8.25| 100.00  2.00  8.00|1|0.0000| 40 | 0|180|
 8 | 1| 001 008| 9.22| 100.00  2.00  9.00|1|0.0000| 40 | 0|180|
 9 | 1| 001 009| 10.20| 100.00  2.00  10.00|1|0.0000| 40 | 0|180|
10 | 1| 001 010| 11.18| 100.00  2.00  11.00|1|0.0000| 40 | 0|180|
11 | 1| 001 011| 12.17| 100.00  2.00  12.00|1|0.0000| 40 | 0|180|
12 | 1| 001 012| 13.15| 100.00  2.00  13.00|1|0.0000| 40 | 0|180|
13 | 1| 001 050| 26.36| 125.00  8.00  2.00|1|0.0000| 40 | 0|180|
14 | 1| 001 051| 26.46| 125.00  8.00  3.00|1|0.0000| 40 | 0|180|
15 | 1| 001 052| 26.59| 125.00  8.00  4.00|1|0.0000| 40 | 0|180|
16 | 1| 001 053| 26.76| 125.00  8.00  5.00|1|0.0000| 40 | 0|180|
17 | 1| 001 054| 26.96| 125.00  8.00  6.00|1|0.0000| 40 | 0|180|
18 | 1| 001 055| 27.20| 125.00  8.00  7.00|1|0.0000| 40 | 0|180|
19 | 1| 001 056| 27.48| 125.00  8.00  8.00|1|0.0000| 40 | 0|180|
20 | 1| 001 057| 27.78| 125.00  8.00  9.00|1|0.0000| 40 | 0|180|
21 | 1| 001 058| 28.12| 125.00  8.00  10.00|1|0.0000| 40 | 0|180|
22 | 1| 001 059| 28.50| 125.00  8.00  11.00|1|0.0000| 40 | 0|180|
23 | 1| 001 060| 28.90| 125.00  8.00  12.00|1|0.0000| 40 | 0|180|
24 | 1| 001 061| 29.33| 125.00  8.00  13.00|1|0.0000| 40 | 0|180|
    
```

Note that the line ID has a binary zero. We would fix that by dumping the headers with **BHED 10.1.4**, then editing that zero out, replacing it with perhaps a space or some other valid ASCII character. This would be followed by an upload of the edited header file into the \*.seg data by a second run of **BHED**. Some renaming would be required. The flow would look like this:

```

bhed setg1001.seg 01.hed 1
(edit the file 01.hed, say with VI)
bhed setg1001.seg 01.hed 0
mv bhedsetg.seg 1001.seg
    
```

The final result would be over writing 1001.seg with the corrected header version.

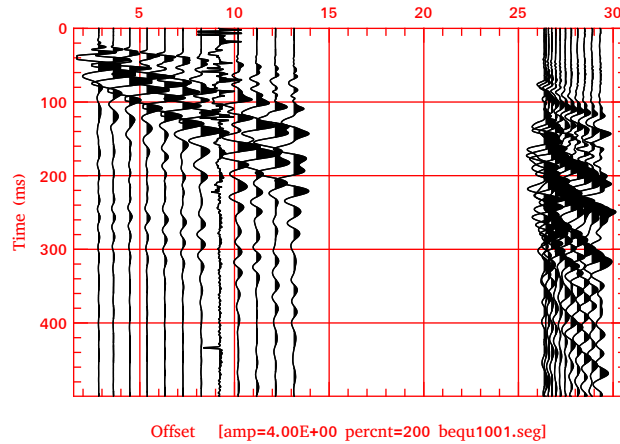


Figure 58: An example of what a plot by offset might look like, trace equalized with BEQU 12.0.9.

### 10.1.8 GENVSP

This is a Vertical Seismic Profile, VSP (down-hole) survey pattern generator. It is an interactive program that creates a NEZ file that can be used to assign geometry to a down-hole survey. Suitable for use with either Bison or SEG-2 file formats. The naming of the shot record files is used to determine the type of data. Bison file names are 8 character alpha numeric without a suffix, SEG-2 files are assumed to have names like 1000.DAT

The hole is assumed to be vertical, reference phone fixed. This code is hardwired for how the author acquires data. Note the initial Channel order switch assumption choice at the beginning of the program execution. See [Michaels \(1998\)](#).

Output files are:

- \*.nez (Northing Easting Elevation file) (10.1.8.1)
- geom (calls topcon program, unites \*.nez survey data with the data file, producing an \*.xyz header file) (10.1.8.2)
- geom2 (calls script go1 for each shot effort) (10.1.8.3)
- go1 (calls BHED program to unite \*.xyz header data with seismic data into BSEGY data formatted files.) (10.1.8.4)

NOTE: change the permissions on geom, geom2, and go1 files to executable. For example: `chmod +x geom`

EXAMPLE RUN:

```
Down-hole VSP Pattern Generator
For Setting Geometry
Handles both Bison and SEG-2 File Formats

Set Channel Order Switch
1=ascending 1,2,3=downhole 4,5,6=reference
-1=descending 6,5,4=downhole 3,2,1=reference
2=ascending 1,2,3=down 4,5,6=ref,7=load_cell
-2=descending 7=load_cell,6,5,4=down 3,2,1=ref

1

-----BOREHOLE-----
Enter 6 char. name for nez file (ex. STP001)
STP001
Enter 4 char. LINEID
0001
Enter Z-Datum: Casing Elevation
849.
```

```

BOREHOLE LOCATION:
Borehole is origin of the local coordinate system
Source and Reference phone locations are x,y
relative to borehole.

Following entries will shift every x,y input to
a final global coordinate system:
Enter Global x-coord. of borehole
1000
Enter Global y-coord. of borehole
1000
Enter number of sources
1
FOR THIS SOURCE:
Enter Shot Record Names 8char: First
STP30001
Enter Shot Record Names 8char: Last
STP30100
STP30001STP30100
Enter Source: x, y, z_sub_CE (positive down)
0,-1,-.5
Enter Source Polarization: azi, ver
0, 180

-----REFERENCE RECEIVER-----
Enter Reference: x, y, z_sub_CE (positive down)
0,+1,-.4
Enter Reference Polarizations: R-azi, T-azi
0,270

-----BOREHOLE PHONES-----
Enter Bulk Shift (Added To Geophone Depth ONLY)
.3
For Shot: SP01 AZI= 0 VER=180

Enter Station Spacing: dz
.25
Enter First Station Depth: zmax
20
Enter Last Station Depth: zmin
0.5
Number of receivers = 79
-----
CHECK DATA TYPE

Files like XXXX0001 detected, ID=BISON
Is above ID Correct, or override needed?
1=YES correct 0=NO incorrect
1

```

In the above example, the tool has 6 channels, so there will be 6 lines for each down-hole station. The first 3 lines are the down-hole components, 2 horizontal, 1 vertical. The next 3 lines are the reference phone (note the elevation column does not change for the reference phone since it is stationary).

**10.1.8.1 nez** The NEZ files starts like this:

```

0001      999.0000  1000.0000  849.5000  SP01
0001      1000.0000  1000.0000  828.7000  VP0001
0002      1000.0000  1000.0000  828.7000  VP0002
0003      1000.0000  1000.0000  828.7000  VP0003
0004      1001.0000  1000.0000  849.4000  VP0004
0005      1001.0000  1000.0000  849.4000  VP0005
0006      1001.0000  1000.0000  849.4000  VP0006
0007      1000.0000  1000.0000  828.9500  VP0007
0008      1000.0000  1000.0000  828.9500  VP0008
0009      1000.0000  1000.0000  828.9500  VP0009
0010      1001.0000  1000.0000  849.4000  VP0010
0011      1001.0000  1000.0000  849.4000  VP0011
0012      1001.0000  1000.0000  849.4000  VP0012
.
.
.
0469      1000.0000  1000.0000  848.2000  VP0469
0470      1000.0000  1000.0000  848.2000  VP0470
0471      1000.0000  1000.0000  848.2000  VP0471
0472      1001.0000  1000.0000  849.4000  VP0472
0473      1001.0000  1000.0000  849.4000  VP0473
0474      1001.0000  1000.0000  849.4000  VP0474

```

**10.1.8.2 geom** The bash script, **geom** file starts like this (for bison data in this instance, calls topcon [10.1.3](#)):

```

topcon STP001.nez STP30001 0001 0.0 1 6 0001 0006 1 0. 0 180 0 270
topcon STP001.nez STP30002 0001 0.0 1 6 0007 0012 2 0. 0 180 0 270
topcon STP001.nez STP30003 0001 0.0 1 6 0013 0018 3 0. 0 180 0 270
topcon STP001.nez STP30004 0001 0.0 1 6 0019 0024 4 0. 0 180 0 270
topcon STP001.nez STP30005 0001 0.0 1 6 0025 0030 5 0. 0 180 0 270
topcon STP001.nez STP30006 0001 0.0 1 6 0031 0036 6 0. 0 180 0 270
.
.
.
topcon STP001.nez STP30074 0001 0.0 1 6 0439 0444 74 0. 0 180 0 270
topcon STP001.nez STP30075 0001 0.0 1 6 0445 0450 75 0. 0 180 0 270
topcon STP001.nez STP30076 0001 0.0 1 6 0451 0456 76 0. 0 180 0 270
topcon STP001.nez STP30077 0001 0.0 1 6 0457 0462 77 0. 0 180 0 270
topcon STP001.nez STP30078 0001 0.0 1 6 0463 0468 78 0. 0 180 0 270
topcon STP001.nez STP30079 0001 0.0 1 6 0469 0474 79 0. 0 180 0 270

```

**10.1.8.3 geom2** The bash script **geom2** file starts like this:

```

go1 001
go1 002
go1 003
go1 004
go1 005
go1 006
.
.
.
go1 095
go1 096
go1 097
go1 098
go1 099
go1 100

```

**10.1.8.4 go1** For the instance of Bison data, the **go1** file is a bash script (calls bis2seg [3.1.3](#)):

```

bis2seg STP30$1
bhed STP30$1.seg STP30$1.xyz 0
mv bhedSTP3.seg S$1.seg
rm STP30$1.seg

```

### 10.1.9 GENBHOD

This is an program that generates bash script to conduct Principle Component Analysis (PCA) on down-hole data (Michaels, 2001b). A down-hole tool will rotate as it comes up the hole, and there is a need to determine the horizontal component orientations. This is an interactive program. The following is an example log of a run for a single station (normally, the last file will reflect many stations in a survey):

```

genbhod
|-----|
| Copyright (C) 2009 P. Michaels |
| All rights reserved |
see GNU General Public License

WARNING: !!
See Source Code, genbhod.f, or BSU documentation
(man pages and BSU user Guide)
before you use this program. It is hardwired for
a specific type of acquisition.

enter 1char_ALPHA PREFIX
c
enter FIRST FILE NUMBER (<=3digits)
for which source polarization is 270 deg.
009
enter LAST FILE NUMBER (<=3digits)
for which source polarization is 270 deg.
009
enter UP/DOWN SWITCH
-1= 90 Azimuth File Number 1 LESS than 270 Az
+1= 90 Azimuth File Number 1 MORE than 270 Az
1
enter azimuth of bowspring(R-comp)
180
OUTPUT====> Downhole: gobhodo
OUTPUT====> Reference: gobhodoR
OUTPUT====> Downhole: gorunbhod
OUTPUT====> Reference: gorunbhodR
-----
REMEMBER to change permissions on the
above files to execute.
-----
IF examining the Down-hole Phone

1. Run gobhodo in directory with 6 chan
records (3 down, 3 reference phones)

2. Run gorunbhod in directory with files
that are named hxxxyy.seg

-----
IF examining the Reference Phone

1. Run gobhodoR in the directory with
the 6 channel records.

2. Run gorunbhodR in the directory with
files that are named rxxxyy.seg
-----

```

After running the interactive program, change permissions of the generated scripts. For example, `chmod +x go*`. One can analyze either the fixed reference phone (scripts `gobhodoR` and `gorunbhodR`) or the down-hole phone (scripts `gobhodo` and `gorunbhod`). Because of a  $180^\circ$  ambiguity in the result of the PCA analysis, one must observe and record the tool orientation when it comes out of the hole. It is assumed that one starts logging the data with the first station at the deepest depth in the hole, pulling the tool up to the surface. For tools with a clamping bowspring, determine the orientation of the horizontal components relative to the bowspring and observe

the bowspring orientation when the tool is at the last station. This PCA procedure is for horizontal sources where there are two source efforts at each subsurface station. These are of opposite polarity and recorded in separate files which can be subtracted to enhance SH-waves (Michaels, 1998).

The procedure is to run the `gobhodo` script which will scale and then subtract the two source efforts. For example, at a station where the \*.seg files are `c009.seg` and `c010.seg`, the result will be a file, `h010009.seg`. For the single station example above, the `gobhodo` script is:

```
bscl c010.seg 1 1 3
bscl c009.seg 1 1 3
bsum bsclc010.seg bsclc009.seg -1.0
mv bsumbscl.seg h010009.seg
```

Next, the script `gorunbhod` is run. It consists of `BHOD` program commands like this:

```
bhod h010009.seg 2 3 50 90.0 180.0 +90.0
```

See program `BHOD 10.1.11` for more. The final result is a file `bhod.lst` which contains the horizontal component orientations that will be applied to data headers (`BTOR 12.2.2`) and later rotate the data as desired `BROT 12.2.4`. In addition to the `bhod.lst` file, the `gorunbhod` script calls to `BHOD` produces Postscript files showing the analysis results (see Figure 59).

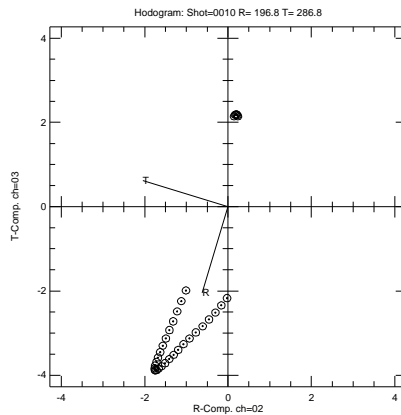


Figure 59: BHOD: plot produced showing PCA results for a geophone at about 19.39 meters depth. File `bhod.lst`: (00010 196.8 286.8) = (seq. R-azi, T-azi)

### 10.1.10 GENBHODV

Interactive program go generate bash scripts to determine down hole tool orientation by PCA with a VERTICAL IMPACT SOURCE. This is experimental, assumes Rayleigh waves generated by source. To run, type: `genbhodV`

**EXPERIMENTAL** program generates 4 bash script files which can be run to determine geophone orientations based on the large particle motion. The concept is experimental. In short, one uses the horizontal motion of the Rayleigh wave in the context of the experimental setup to determine horizontal tool orientation. It works well for the reference phone, but your mileage may vary down-hole (depending on the depth of penetration of the Rayleigh wave, and on the subsurface nodal pattern of the Rayleigh wave). Two of the scripts are for a surface, reference geophone, and two are for the down-hole geophone. The Principal Component Analysis (PCA) is actually done by the program, `BHOD 10.1.11`. There are many assumptions made in this code.

**10.1.10.1 Example Log** The following log is for illustration, and is for a single source effort by a vertical source recorded on a file, `c200.seg`. In practice one would have many files, and the number of files will be set by the first and last file numbers.

```
WARNING: !!
See Source Code, genbhodV.f, or BSU documentation
(man pages and BSU user Guide)
before you use this program. It is hardwired for
a specific type of acquisition.
```

```

enter 1char_ALPHA PREFIX
c
enter FIRST FILE NUMBER (<=3digits)
200
enter LAST FILE NUMBER (<=3digits)
200
enter azimuth of bowspring(R-comp)
180
OUTPUT====> Downhole: gobhodo
OUTPUT====> Reference: gobhodoR
OUTPUT====> Downhole: gorunbhod
OUTPUT====> Reference: gorunbhodR
-----
REMEMBER to change permissions on the
above files to execute.
-----
IF examining the Down-hole Phone

1. Run gobhodo in directory with 6 chan
records (3 down, 3 reference phones)

2. Run gorunbhod in directory with files
that are named hxxx.seg
-----
IF examining the Reference Phone

1. Run gobhodoR in the directory with
the 6 channel records.

2. Run gorunbhodR in the directory with
files that are named rxxx.seg
EXPERIMENTAL APPROACH on Rayleigh Wave
-----

```

The **gobhodo** script scales the data (**BSCL 12.0.10** by the maximum absolute value on traces 1 to 1 (ie. trace 1). The last option 3 sets the choice to maximum absolute value. The script is:

```

bscl c200.seg 1 1 3
mv bsclc200.seg h200.seg

```

The **gorunbhod** calls the **BHOD 10.1.11** program:

**bhod h200.seg 2 3 50 0.0 180.0 +90.0** The file **bhod.lst** contains the solution:

```
00200 263.0 353.0
```

(file number, R-azimuth, T-azimuth) The **bhod.lst** is input to program **BTOR 12.2.2** which rotates the data and updates the headers.

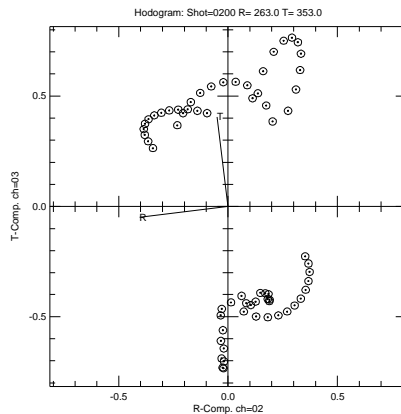


Figure 60: BHOD: plot produced showing PCA results for a geophone at about 11.68 meters depth.



### 10.1.11 BHOD

Hodogram analysis by Principal Component Analysis (PCA) can determine the orientation of the horizontal components of a 3-C geophone in a bore hole. Program **BHOD** does this analysis and outputs a file, **bhod.lst** that has rows of 3 numbers, (seq., R-azimuth, T-azimuth). The sequence number corresponds to the \*.seg files. File, **bhod.lst**, is then used by **BTOR 12.2.2** to update headers. Program **BROT 12.2.4** is then used to rotate the data to a desired orientation.

In the case of a horizontal impulse source, two opposite polarities will be struck for each geophone depth. Depending on which source blow azimuth is first, the **bhod.lst** sequence number will either be the first or second blow, and the result of PCA will be applied to both source efforts at the depth being analyzed. In a typical survey there will be twice as many \*.seg files as depth stations occupied. One surveys from the bottom to the top of the hole, and should make an **IMPORTANT** observation of the tool orientation at the surface to resolve the 180<sup>0</sup> ambiguity. Helper scripts are generated by **GENBHOD 10.1.9**.

In the case of a vertical impulse source, experimental helper scripts are generated by program **GENBHODV 10.1.10**. In this case, the procedure is designed to observe the large amplitude hodogram motion (which may be a Rayleigh wave). Rayleigh waves are a mix of P-SV motion. The P-motion is horizontal and may provide orientation information. Your mileage will vary depending how deep the Rayleigh waves motion penetrates.

The command line arguments to **BHOD** are:

```

bhod  infile chR chT ipct saz azctl tsw1

infile =input file name
chR    = channel with R-component (int)
chT    = channel with T-component (int)
ipct   = percent of max amplitude to include (int)
saz    = source azimuth (ie E-W, then 90 deg)
azctl  = desired direction for R- (bowspring)
        (azctl resolves 180 deg ambiguity PCA)
tsw1   = switch to set T-comp relative to R-comp
        T-comp Azimuth= R-Comp + tsw1
        Typically, tsw1= +90.(downhole) or -90.(ref)

```

EXAMPLE:

```

bhod h002001.seg 2 3 50 90.0 315.0 +90.0

```

**bhod.lst** file:

```

01002 258.6 348.6

```

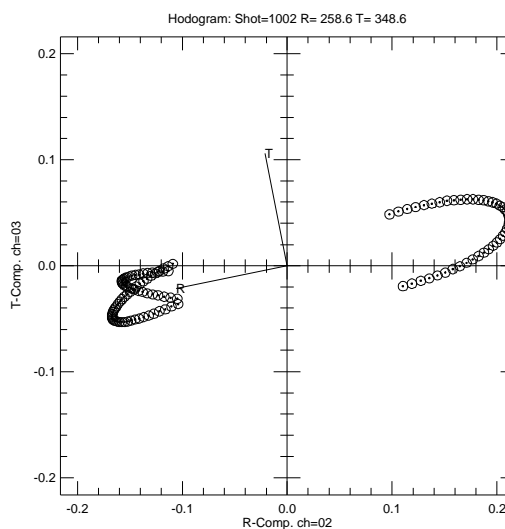


Figure 61: BHOD: plot produced showing PCA results for a geophone at about 20 meters depth.

**10.1.12 BNEZ**

Each row of an NEZ file provides the (Seq. Northing, Easting, Elevation, Tag). The Tag specifies either a source (SP) or geophone (VP, voltage point). One way of running the program is to run BNEZ twice to create shot.nez and phones.nez files which are then merged into a single NEZ file. Depending on if you have Bison or SEG-2 data recorded, use either program **TOPCON 10.1.3** or **TOPCON2 10.1.5** to generate \*.xyz files that can be used by the **BHED 10.1.4** program to set the geometry in the \*.seg file headers.

The command line arguments are:

```
bnez outfile n-points tag so yo xo zo ido dy dx dz did

outfile = output file name (ex. aaaa0001.nez

n-points = number of survey points to generate
tag      = 1 tag=VP
          = 2 tag=SP
so       = first value of sequence number
yo       = northing of first point
xo       = easting of first point
zo       = elevation of first point
ido      = initial ID number
dy       = spacing between points in north direction
dx       = spacing between points in east direction
dz       = spacing in elevation between points
did      = interval in ID between points
```

**10.1.12.1 Example, BNEZ** The commands for a single shot gather, Bison data, are:

```
bnez 000001.nez 1 2 1 0 0 0 01 1 1 1 1
bnez 000002.nez 12 1 2 140. 0. 0. 01 10. 0. 0. 1
cp 000001.nez LOG001.nez
cat 000002.nez >>LOG001.nez
cat LOG001.nez
```

|    |            |          |          |       |
|----|------------|----------|----------|-------|
| 1  | 0.000000   | 0.000000 | 0.000000 | SP001 |
| 2  | 140.000000 | 0.000000 | 0.000000 | VP001 |
| 3  | 150.000000 | 0.000000 | 0.000000 | VP002 |
| 4  | 160.000000 | 0.000000 | 0.000000 | VP003 |
| 5  | 170.000000 | 0.000000 | 0.000000 | VP004 |
| 6  | 180.000000 | 0.000000 | 0.000000 | VP005 |
| 7  | 190.000000 | 0.000000 | 0.000000 | VP006 |
| 8  | 200.000000 | 0.000000 | 0.000000 | VP007 |
| 9  | 210.000000 | 0.000000 | 0.000000 | VP008 |
| 10 | 220.000000 | 0.000000 | 0.000000 | VP009 |
| 11 | 230.000000 | 0.000000 | 0.000000 | VP010 |
| 12 | 240.000000 | 0.000000 | 0.000000 | VP011 |
| 13 | 250.000000 | 0.000000 | 0.000000 | VP012 |

The next step is to combine the Bison file headers with the NEZ and produce an output \*.xyz file:

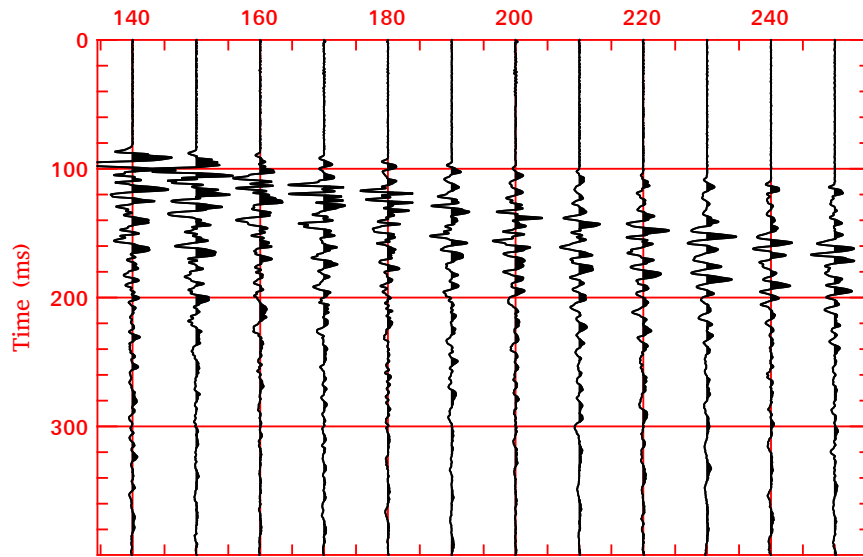
```
topcon LOG001.nez L0ST0001 0001 0.0 1 12 001 012 1 0. 0 0 0 0
cat L0ST001.xyz
```

```
&bhed
lowcut= 16, highct= 500, year=1992, day=0303,
line='0001', hour=17, minute=07,
sdepth= 0.0, uphole=0.000, phone='VERT', srec= 001,
&end
001 0.0000 001 0.000 0.000 0.000 001 0.000 140.000 0.000 60 000 000 000 000
002 0.0000 001 0.000 0.000 0.000 002 0.000 150.000 0.000 60 000 000 000 000
003 0.0000 001 0.000 0.000 0.000 003 0.000 160.000 0.000 60 000 000 000 000
004 0.0000 001 0.000 0.000 0.000 004 0.000 170.000 0.000 60 000 000 000 000
005 0.0000 001 0.000 0.000 0.000 005 0.000 180.000 0.000 60 000 000 000 000
006 0.0000 001 0.000 0.000 0.000 006 0.000 190.000 0.000 60 000 000 000 000
007 0.0000 001 0.000 0.000 0.000 007 0.000 200.000 0.000 60 000 000 000 000
008 0.0000 001 0.000 0.000 0.000 008 0.000 210.000 0.000 60 000 000 000 000
009 0.0000 001 0.000 0.000 0.000 009 0.000 220.000 0.000 60 000 000 000 000
```

```
010 0.0000 001 0.000 0.000 0.000 010 0.000 230.000 0.000 60 000 000 000 000
011 0.0000 001 0.000 0.000 0.000 011 0.000 240.000 0.000 60 000 000 000 000
012 0.0000 001 0.000 0.000 0.000 012 0.000 250.000 0.000 60 000 000 000 000
```

The final step is to use **BIS2SEG 3.1.3** to convert the Bison file to \*.seg, and then use **BHED 10.1.4** to apply the headers to the \*.seg file.

```
bis2seg LOST0001
bhed LOST0001.seg LOST0001.xyz 0
#plot the data by offset
bplt bhedLOST.seg 0 0 1 1 12 0 .5 1 4E-3 200
bdump bhedLOST.seg 0
cat bdump.lst
```



Offset [amp=4.00E-03 percent=200 bhedLOST.seg]

Figure 62: BNEZ: Plot of Bison file data with geometry added.

```
-----|
| PARTIAL SEG Y HEADER DUMP |
|                               |
|          bhedLOST.seg       |
```

```
-----|
Length = 2000 samples           | Shot Elevation = 0.0
Sample Interval = 0.00020 sec. | Shot Depth = 0.0
Delay Time = 0 msec.           | Up Hole Time = 0 msec
Low Cut Filter = 16 Hz.        | Shot X-COORD = 0.00
High Cut Filter = 500 Hz.     | Shot Y-COORD = 0.00
Line ID: 0001                  | Shot Date (year.moday) = 1992.0303
Shot Orientation:              | Shot Time (hr:min) = 17:07
Azimuth= 0 Deg. Vertical= 0 Deg. | Charge Size (grams)= 0
-----|
```

| TRACE # | SHOT REC. | STATION SHOT REC | OFFSET | ELEV. | RECEIVER X-COORD | Y-COORD | VERT FOLD | 1STBRK (SEC.) | K-GAIN (dB) | AZI | VER |
|---------|-----------|------------------|--------|-------|------------------|---------|-----------|---------------|-------------|-----|-----|
| 1       | 1 001     | 001              | 140.00 | 0.00  | 0.00             | 140.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 2       | 1 001     | 002              | 150.00 | 0.00  | 0.00             | 150.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 3       | 1 001     | 003              | 160.00 | 0.00  | 0.00             | 160.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 4       | 1 001     | 004              | 170.00 | 0.00  | 0.00             | 170.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 5       | 1 001     | 005              | 180.00 | 0.00  | 0.00             | 180.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 6       | 1 001     | 006              | 190.00 | 0.00  | 0.00             | 190.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 7       | 1 001     | 007              | 200.00 | 0.00  | 0.00             | 200.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 8       | 1 001     | 008              | 210.00 | 0.00  | 0.00             | 210.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 9       | 1 001     | 009              | 220.00 | 0.00  | 0.00             | 220.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 10      | 1 001     | 010              | 230.00 | 0.00  | 0.00             | 230.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 11      | 1 001     | 011              | 240.00 | 0.00  | 0.00             | 240.00  | 1 0.0000  | 60            | 0           | 0   |     |
| 12      | 1 001     | 012              | 250.00 | 0.00  | 0.00             | 250.00  | 1 0.0000  | 60            | 0           | 0   |     |

**10.1.13 TOP2NEZ**

Topcon is one of a number of Electronic Distance Measuring (EDM) instruments. It can be controlled with an FC4 module that stores measurements in an ASCII format assuming a Microsoft file convention. For example, consider a file survey.n:

```
00001
10000.00000
10000.00000
1000.00000
BP1
00003
10000.00000
10000.00000
1000.00000
C2-48
.
.
.
00266
10318.48928
10144.12327
1002.47977
SLEDGE7
00267
10320.68105
10136.84087
995.98539
1-SP5A
00268
10205.99591
10104.81427
1002.29261
SLEDGE6
```

There are tags, sequence numbers and (y,x,z) coordinates, one item per line. This program converts the file to a NEZ file format, all items in a single line corresponding to the tag. For example, if we issue the command `top2nez survey.n`, we have output file `survey.n.nez`:

```
00001      10000.00000 10000.00000 1000.00000  BP1
00003      10000.00000 10000.00000 1000.00000  C2-48
.
.
.
00266      10318.48928 10144.12327 1002.47977  SLEDGE7
00267      10320.68105 10136.84087 995.98539   1-SP5A
00268      10205.99591 10104.81427 1002.29261  SLEDGE6
```

The NEZ format is read by BSU programs. Program TOP2DXF 10.1.14 can be used to create a CAD file for making base maps.

#### 10.1.14 TOP2DXF

The command line arguments are:

```
*.nez file format(12x,f12.0,f12.0,f12.0,a)

top2dxf   infile   isw1   ilabel   txtsiz

infile   =*.nez input file name
isw1     =switch to control limits
         0=no limits header
         1=limits based on min and max values
ilabel   0=no printing of point labels
         1=print labels
txtsiz   =size of text in coord. units (float)
```

For an example, consider file **samp0000.nez**:

```
 1  0.000000  0.000000 100.000000 SP001
 2  0.000000  2.000000 101.000000 SP002
 3  0.000000  4.000000 102.000000 SP003
 4  0.000000  6.000000 103.000000 SP004
 5  2.000000  2.000000 100.000000 VP001
 6  3.000000  2.000000 100.000000 VP002
 7  4.000000  2.000000 100.000000 VP003
 8  5.000000  2.000000 100.000000 VP004
 9  6.000000  2.000000 100.000000 VP005
10  7.000000  2.000000 100.000000 VP006
11  8.000000  2.000000 100.000000 VP007
12  9.000000  2.000000 100.000000 VP008
13 10.000000  2.000000 100.000000 VP009
14 11.000000  2.000000 100.000000 VP010
15 12.000000  2.000000 100.000000 VP011
16 13.000000  2.000000 100.000000 VP012
17  2.000000  8.000000 125.000000 VP050
18  3.000000  8.000000 125.000000 VP051
19  4.000000  8.000000 125.000000 VP052
20  5.000000  8.000000 125.000000 VP053
21  6.000000  8.000000 125.000000 VP054
22  7.000000  8.000000 125.000000 VP055
23  8.000000  8.000000 125.000000 VP056
24  9.000000  8.000000 125.000000 VP057
25 10.000000  8.000000 125.000000 VP058
26 11.000000  8.000000 125.000000 VP059
27 12.000000  8.000000 125.000000 VP060
28 13.000000  8.000000 125.000000 VP061
```

In a terminal, we type the command:  
top2dxf samp0000.nez 0 1 .25

Figure 63 illustrates how the output file, **samp0000.dxf** can be read by a common CAD program (here Qcad). Other programs that can read Digital Exchange Format (DXF) files include Microstation and Autocad. Raw EDM files, like from Topcon FC4 controllers can be converted to the NEZ format using **TOP2NEZ 10.1.13**.

|       |       |
|-------|-------|
| VP012 | VP061 |
| VP011 | VP060 |
| VP010 | VP059 |
| VP009 | VP058 |
| VP008 | VP057 |
| VP007 | VP056 |
| VP006 | VP055 |
| VP005 | VP054 |
| VP004 | VP053 |
| VP003 | VP052 |
| VP002 | VP051 |
| VP001 | VP050 |

SP001 SP002 SP003 SP004

Figure 63: QCAD: Qcad used to read the file samp0000.dxf and exported to a PDF file. The point SP001 is at the origin, (0,0,0).

### 10.1.15 TOPBCRD

An NEZ file can be transformed by scale, shift, and rotations using this program. The program does the same thing as program **BCRD 10.1.16** (which operates on BSEGY, \*.seg, format data). The command line arguments are:

```
topbcrd  infile  theta  sfact  x0  y0  z0

infile  =input file name
theta   =angle (deg) from current to new x-axis
        (+)theta counterclockwise (-)theta clockwise
sfact   =units scale factor: (old)*(sfact)=(new)
x0      =X-offset in new coord. units
y0      =Y-offset in new coord. units
z0      =Y-offset in new coord. units
```

An example of modifying the **TOP2NEZ 10.1.13** example above follows. The transform is 45 degree counter-clockwise rotation, scale factor = 1, shift of (20,20,0) meters. The point **SP001** is at the origin, (0,0,0) in the original \*.nez file. After transform, the point, **SP001**, is at (20,20,0). The grid rotates counter-clockwise (or the survey points appear rotated clockwise).

```
topbcrd  samp0000.nez 45. 1. 20. 20. 0.
top2dxf  samp0000.mod 0 1 .25
```

Figure 64 shows the modified survey file after plotting with Qcad. The new \*.nez file is samp000.mod.

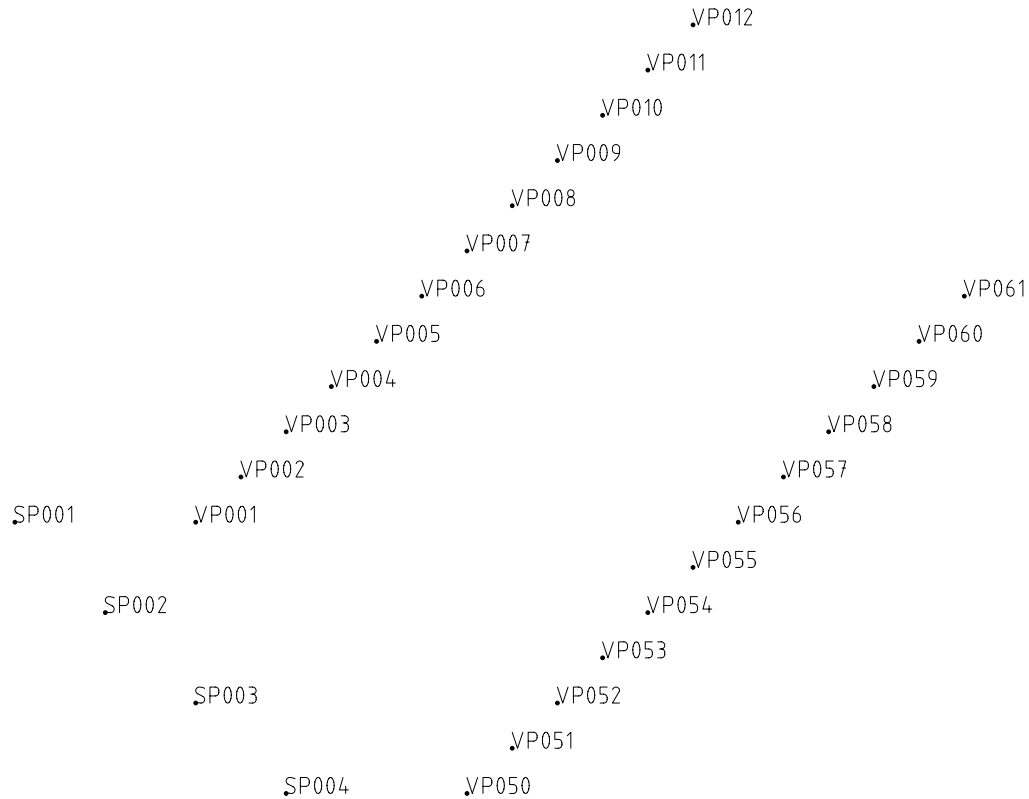


Figure 64: QCAD: Qcad plot of modified samp0000.nez file, samp0000.mod. Point SP001 is now at (20,20,0).

### 10.1.16 BCRD

This program does the same thing as **TOPBCRD 10.1.15**. The difference is that the input file is a BSEGY, \*.seg, file instead of a NEZ survey file. The command line arguments are:

```

bcd rd infile theta sfact x0 y0 z0

infile =input file name
theta  =angle (deg) from current to new x-axis
        (+)theta counterclockwise (-)theta clockwise
sfact  =units scale factor: (old)*(sfact)=(new)
x0     =X-offset in new coord. units
y0     =Y-offset in new coord. units
z0     =Z-offset in new coord. units

```

**NOTE:** When using **BCRD** or **TOPBCRD**, be aware that rotation and translation at the same time may not be the same as translation first, output a file, then rotation second on the translated file. The result of the translation and rotation operations may be viewed by running the program **BCAD 10.1.17** which produces a DXF file from the altered headers.

**10.1.17 BCAD**

Similar to **TOP2DXF 10.1.14**. Rather than reading an NEZ survey file, this program takes a BSEGY (\*.seg) file for input, and outputs a DXF file suitable to be read by a cad program. The command line arguments are:

```

bcad      infile  isw1  ilabel  txtsiz

infile   =*.seg input file name
isw1     =switch to control limits
         0=no limits header
         1=limits based on min and max values
ilabel   0=no printing of point labels
         1=print labels
txtsiz   =size of text in coord. units (float)
    
```

```

BCAD Example:
bcad c008.seg 1 1 1.0
qcad bcadc008.dxf
    
```

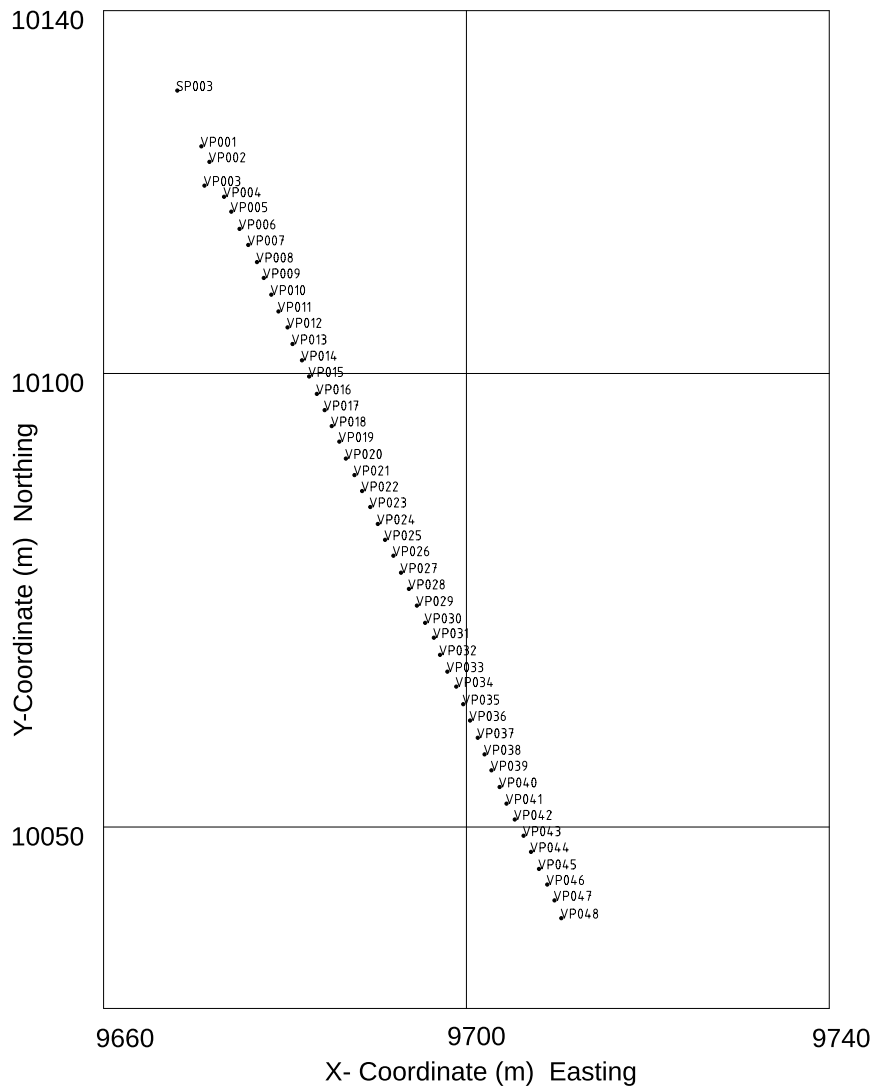


Figure 65: BCAD: DXF file edited, add some coordinates and labels. Editing DXF in QCAD, <http://qcad.org/en/>



## 11 Editing BSEGY Data

Once data are in the BSEGY format (ie. \*.seg files), they can be edited in a number of ways:

- **BMRG 11.0.1** Merge data from many files into one file
- **BEDT 11.0.2** Edit data by traces aperture, time aperture. Edit data by interpolation or decimation of samples (includes anti-alias option for decimation).
- **BRSP 11.0.3** Resample data. Interpolation by augmentation with zeros in frequency domain. Does not introduce any new frequencies.
- **BKIL 11.0.4** Remove or zero out traces by range or by list.
- **BEXT 11.0.5** Extract traces either by shot or receiver name, or by field record number.
- **BOFF 11.0.6** Compute offset header from coordinates of shot and receiver, insert into BSEGY header.
- **BWIN 11.0.7** Temporal window of BSEGY data. Tapers from a start time to full, extends to end time, tapers to zero.
- **BHED 10.1.4** Extract or upload headers for BSEGY data.

### 11.0.1 BMRG

Often data collected in surveys results in a number of files which are numbered sequentially. For example, in down-hole surveys, each file may relate to a down-hole station for a single source effort. There may be a number of components recorded at each station. This would also be the case in walk-a-way surface data collection. **BMRG** permits one to select a sequence of files, and specific traces in each file to output into a single file.

```
bmrq pfix iffile ilfile ifinc iftrc iltrc

pfix: =prefix for input file names
NOTE:pfix length is no. of invariant characters
  Ex. If file names run s001 to s090 then pfix=s0
  Ex. If file names run s001 to s132 then pfix=s
iffile =number of first file (suffix)
EXAMPLE: if file=s001.seg, iffile=001

ilfile =number of last file (suffix)
EXAMPLE: if file=s092.seg, ilfile=092

ifinc =increment for file number (suffix)
iftrc =first trace each file
iltrc =last trace each file
```

For example, consider a down-hole survey with files w001.seg through w166.seg. The file order in each file is:

| Channel | Component               |
|---------|-------------------------|
| 1       | Vertical (down-hole)    |
| 2       | Radial (down-hole)      |
| 3       | Transverse (down-hole)  |
| 4       | Vertical (ref. phone)   |
| 5       | Radial (ref. phone)     |
| 6       | Transverse (ref. phone) |

The reference phone is fixed at the surface and the down hole phone is logged from bottom to surface. We want to collect the transverse down-hole phone, channel 3, and output that to a single file. The command:

```
bmrq w 001 166 2 3 3
```

If we want every source effort, the command is:

```
bmrq w 001 166 1 3 3
```

Figure 66 shows both cases. Since the source blow is 135 degrees from the vertical, the horizontal T-component

will show different polarity of source effort (every trace has the checkered look, peaks against troughs, plotted by elevation).

The data have not been rotated to a standard orientation. The T-component in this example drifts from 313 to 288 degrees azimuth as determined by PCA analysis (see **BHOD 10.1.11**). Program **BTOR 12.2.2** applies the PCA results to data headers, and program **BROT 12.2.4** actually does the rotation to a standard orientation (with respect to the source axis for horizontal component hammer blows).

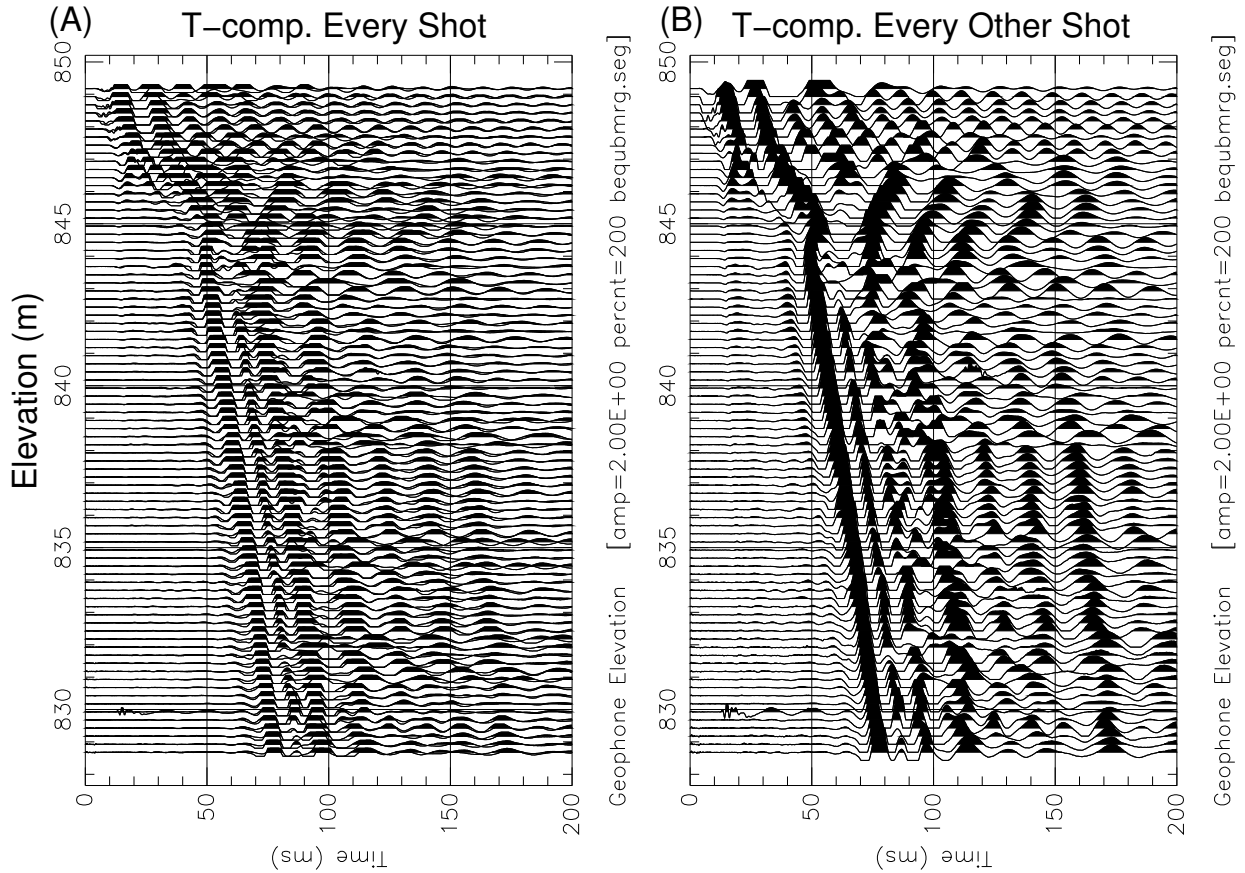


Figure 66: BMRG: A) is plot of all shot efforts (166 traces) and B) is plot of only very other shot (83 traces). NOTE: data are not rotated to a standard orientation, azimuth of T-component drifts up the hole.

### 11.0.2 BEDT

The command line arguments are:

```
bedt infil tmin tmax ifirst ilast idecm iantia

infil  =input file name to edit
tmin   =minimum time to extract data
tmax   =maximum time to extract data
ifirst =first trace to extract (<0 pads left)
ilast  =last trace to extract (>ntraces pads right)
idecm  =decimation factor (idecm>0)
       =interpolation factor (idecm<0)
EXAMPLES:
idecm=1 keep same sample interval
idecm=2 output every other sample
idecm=-2 output samples between originals

iantia =0 no anti-alias filter for resample
       =1 use anti-alias filter for resample
```

Figure 67 shows an example where a data set is resampled to include only 0 to 200 msec. of data, only first 6 traces, and interpolated to .00025 seconds per sample. The command:

```
bedt c008.seg 0 .2 1 6 -2
```

Sinc interpolation does not add any additional frequencies beyond the original Nyquist.

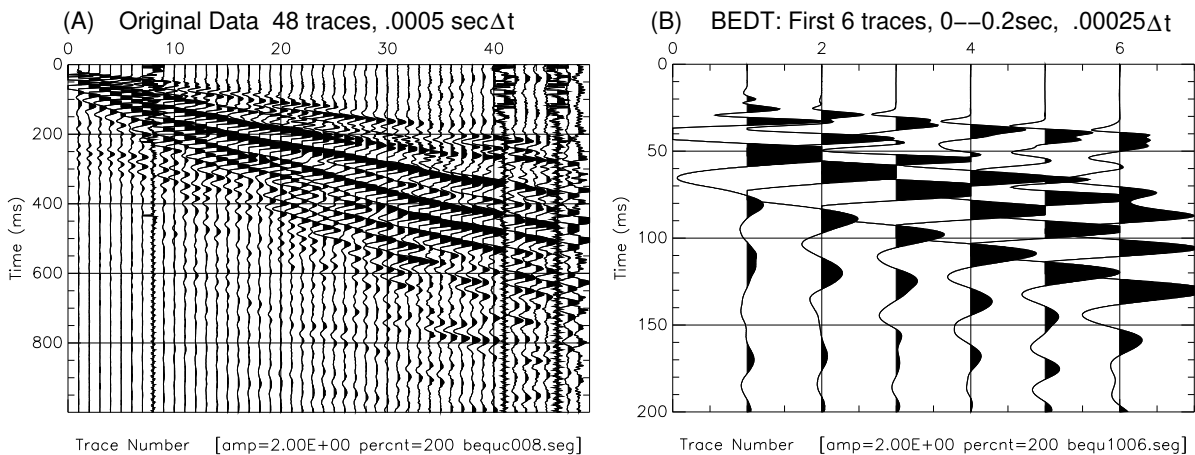


Figure 67: BEDT: (A) Original data, 48 traces, 0-1 seconds, .0005 second sample interval. (B) Edited to only first 6 traces, 0-0.2 seconds, interpolated to .00025 second sample interval.

### 11.0.3 BRSP

Interpolation only resampling. Done in frequency domain by augmentation with zeros. No new frequencies introduced. **Number of sample increases rapidly!**

```
brsp infile ifact tmax
infile =input file name
ifact  = resample factor ifact>=1
       = 1 sample interval halved.
       = 2 sample interval 1/4 of original
       = n sample interval 1/(2**n) of original
       (note: trace size increases by 2**ifact)
tmax   =max time of output trace (float) sec
NOTE: This is a radix 2 algorithm, so trace may be
with zeros before computing new number of samples
BSEGY number of samples header is 16 bit
Maximum number of samples limited to 32,767
```

### 11.0.4 BKIL

The command line arguments are:

```

INDIVIDUAL OPTION-----
bkil infile iopt1 iopt2 ntrc itr itr ...itr

infile:  =input file name
iopt1:   =option 0=kill 1=zero traces
iopt2:   =specify traces 0=individual 1=by range
ntrc    =number of traces to kill or zero
itr..... =trace numbers to kill or zero
RANGE OPTION-----
bkil infile iopt1 iopt2 iftr iltr

infile:  =input file name
iopt1:   =option 0=kill 1=zero traces
iopt2:   =specify traces 0=individual 1=by range
iftr     =first trace
iltr     =last trace

```

Example: Zero noisy traces 8, 41, 46 of data in Figure 67 (A).

```
bkil c008.seg 1 0 3 8 41 46
```

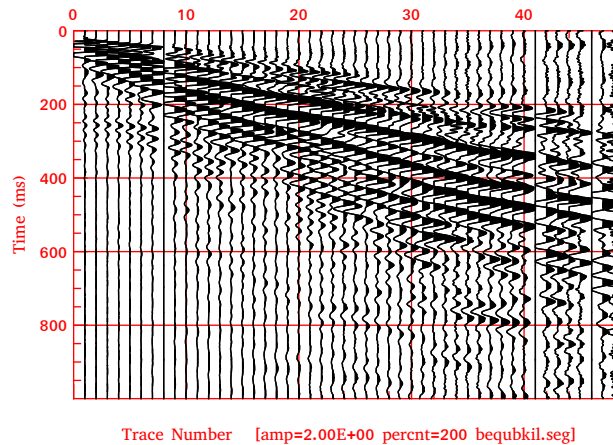


Figure 68: BKIL: Zero noisy traces 8, 41, 46 of data shown in Figure 67 (A).

### 11.0.5 BEXT

Traces can be extracted by either shot or receiver name in the BSEGY headers. Alternatively, field record number can also be used. This is useful when more than one shot record is in a larger file. The command line arguments are:

```

bext  infile  extsw  value

infile  =  input file name
extsw   =  extraction switch (1 char)
        s=  shot name
        r=  receiver name
        f=  field record number
value   =  shot or rec name (4 char)
        or
        field record number (int)

```

WARNING: leading blanks are important

(enclose 4 char string in quotes if on command line)

Use `bdump` program to find names of shots or receivers

For example, consider extracting the traces with receiver label 30 in a file with two shots.

```
bext merged.seg r "030"
```

A partial dump of the headers for `merged.seg` is:

```
30 | 8|8001 030| 148.53| 1000.50 9938.79 9800.50| 1|0.0514| 40 | 0| 0|
78 | 9|9048 030| 85.15| 1000.50 9938.79 9800.50| 1|0.0386| 40 | 0| 0|
```

This shows that traces 30 and 78 are at receiver name "030" and have the same (x,y,z) coordinates.

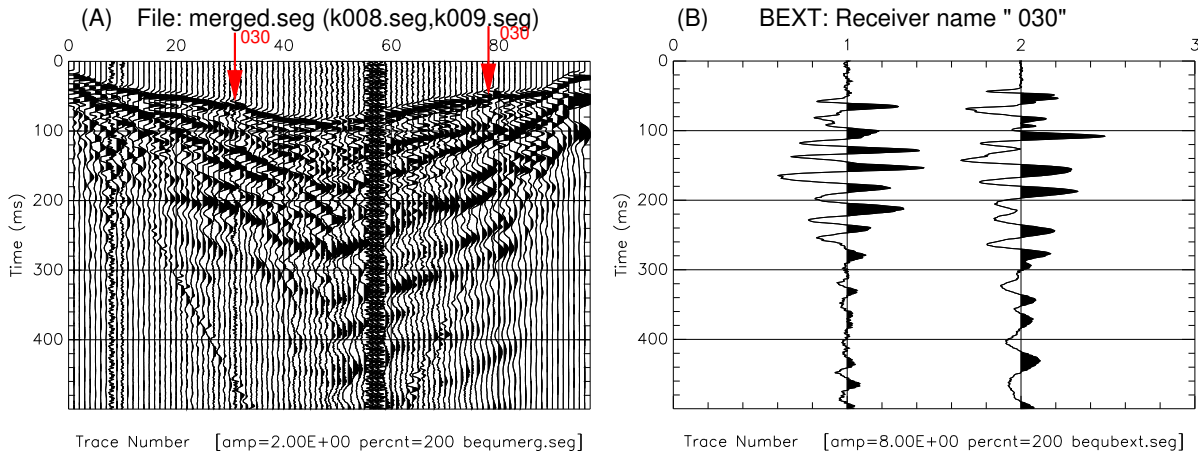


Figure 69: BEXT: Extracted traces from receiver location "030". In the merged file (A) red arrows show receiver "030" and these are replotted in (B). Note the receiver name is 4 characters, "blank,zero,three,zero".

### 11.0.6 BOFF

Some programs need a header value for the shot to receiver offset. This program computes that offset in case it is not in the headers, and then inserts the value in the header of the output file, `boff****.seg`. The only command line argument is the input file name. If the offset BSEGY header value has not been set, it will contain garbage. In the following example, we look at the first trace offset (74 m) in a test file. The commands are:

```
#!/bin/bash
# adds offset to a header value
# This header is not really needed for BSU programs,
# but is useful when converting to Seismic Unix codes
# that require it.
echo " PC linux is little endian"
echo "offset is 4 byte integer in header at hex bytes 0x24 0x25 0x26 0x27"
hexdump -C k007.seg |grep 00000020
echo "00000020      | | | | "
echo "hexdump k007.seg: starting at hexbyte 0x20, List shows 45 00 00 00"
echo "this is garbage"
# BOFF computes offset header, for trace 1 this is 74 meters
boff k007.seg >/dev/null
hexdump -C boffk007.seg |grep 00000020
echo "00000020      | | | | "
echo "hexdump boffk007.seg: shows 4a 00 00 00 "
echo "garbage replaced with 0x4a = 74, correct value "
```

The output when run is:

```
PC linux is little endian
offset is 4 byte integer in header at hex bytes 0x24 0x25 0x26 0x27
```

```

00000020 00 00 01 00 45 00 00 00 61 90 01 00 ff 85 01 00 |....E...a.....|
00000020          | | | |
hexdump k007.seg: starting at hexbyte 0x20, List shows 45 00 00 00
this is garbage
00000020 00 00 01 00 4a 00 00 00 61 90 01 00 ff 85 01 00 |....J...a.....|
00000020          | | | |
hexdump boffk007.seg: shows 4a 00 00 00
garbage replaced with 0x4a = 74, correct value

```

### 11.0.7 BWIN

The command line arguments are:

```
bwin infile tw1 tw2 tw3 tw4
```

```

infile =input file name
tw1    =time of start taper on, amp=0
tw2    =time of taper off,      amp=1
tw3    =time of start taper off amp=1
tw4    =time of taper off,      amp=0

```

For example, `bwin c008.seg .10 .15 .3 .6`

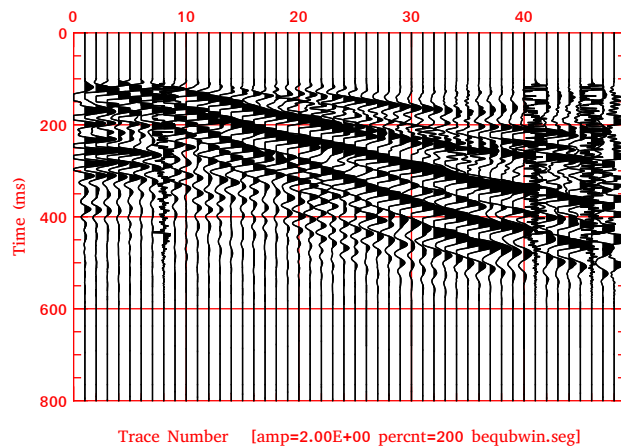


Figure 70: BWIN: Data zeroed outside of the tapered window.

## 12 Signal Processing

The focus here is signal processing of BSEGY data, similar to the section on editing (sec 11), but with more emphasis on altering the data by traditional signal processing methods. Programs include:

- **BREV 12.0.1** Reverse (a) channel order OR (b) polarity
- **BABS 12.0.2** Rectify the data (absolute value).
- **BSDC 12.0.3** Compute DC levels of each trace and show.
- **BRDC 12.0.4** Remove DC levels of each trace.
- **BINT 12.0.5** Integrate BSEGY data.
- **BSRT 12.0.6** Sorts data by offset
- **BRPT 12.0.7** Remove pre-trigger in header and shift data.
- **BDIF 12.0.8** Differentiate BSEGY data.
- **BEQU 12.0.9** Trace equalization of amplitude, BSEGY data.
- **BSCL 12.0.10** Scale data in a profile by a determined or provided factor.
- **BGAR 12.0.11** Exponential gain recovery by source to receiver offset.
- **BGAZ 12.0.12** Exponential gain recover by depth gate for down-hole data.
- **BAGC 12.0.13** Automatic Gain Control (AGC). Choice of single pole exponential envelope or zero-phase box-car envelope.
- **BBAL 12.0.14** Balance amplitudes between two BSEGY data files such that they both have the same MAV  
= (MAV1 + MAV2)/2
- **BSTK 12.0.15** Stacking data in a BSEGY file.
- **BXCR 12.0.16** Auto- or Cross-correlation computed from a file or between two files.
- **BNOS 12.0.17** Computes a band limited noise profile to match the aperture of a template case.
- **BSHF 12.1.1** Static shift BSEGY data by headers or by a file of times, plus a bulk static shift.
- **BSHP 12.2.1** Wiener Least Square Shaping filter. Can apply to data or an alternate file.
- **BTOR 12.2.2** Apply PCA analysis (see **GENBHOD 10.1.9**) to headers of all the \*.seg files in the **bhod.lst** file.
- **GENBROT 12.2.3** Generates a bash script which will run the **BROT** program. That program will rotate the horizontal components of the down-hole data to a desired relationship to the source polarization.
- **BROT 12.2.4** Rotates data based on horizontal component headers or a user supplied value.
- **BFXT 12.3.1** Compute frequency-distance (FX) transform of a shot gather, or compute an inverse transform of amplitude, phase data sets.
- **BCAR 12.3.2** Box car filter, both low- and high-pass options. Fast and specified by a moving average filter duration.
- **BFIL 12.3.3** ARMA filter, Low-pass, Band-pass, or High-pass filters, minimum phase or zero phase, by Bilinear transform.
- **BDCN 12.3.4** Minimum phase deconvolution.
- **BFTR 12.3.5** Filter data with \*.seg file or namelist file.
- **BWHT 12.3.6** Non-linear whitening using AGC in overlapping band-pass filters.

### 12.0.1 BREV

One or more channels may be reversed in polarity, or all channel order reversed.

```
brev infil iop1 nflip ch1 ch2 ...
```

```
infil: =name of input file
```

```
iop1      0=reverse channel order
```

```
or
```

```
1=reverse data polarity
```

```
nflip =number of channels to be polarity reversed
```

```
ch1   =number of channel reverse data polarity
```

```
ch2   =number of channel reverse data polarity
```

```
ch3   =number of channel reverse data polarity
```

```
... ch_nflip. NOTE: if nflip=number of channels,
    then all channels will be reversed in pol. )
    (no need to input ch1, ch2,...)
```

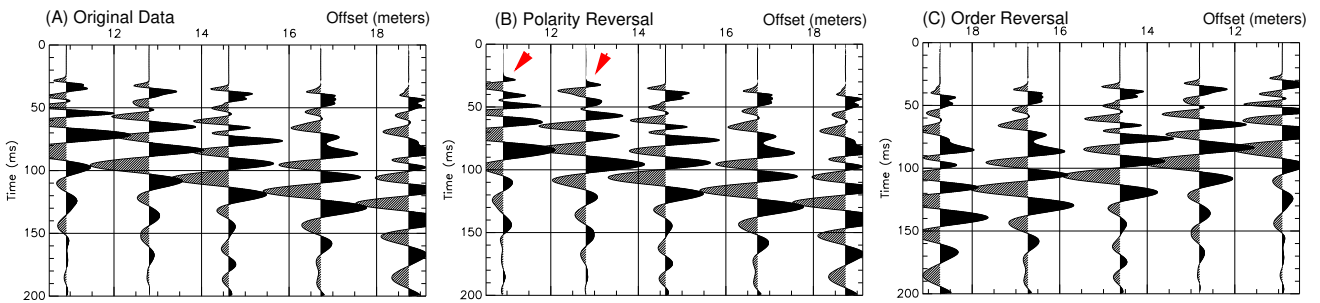


Figure 71: BREV: (A) original data, (B) reverse polarity first 2 channels, (C) reverse channel order. Data plotted by offset.

### 12.0.2 BABS

Takes the absolute value. The only command line argument is the input file name. Figure 72 shows an example.

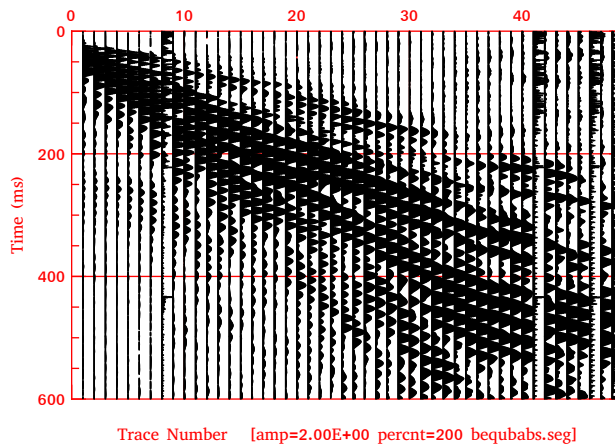


Figure 72: BABS: Rectify data (take absolute value).



### 12.0.3 BSDC

The only command line argument is the input BSEGY file. The output is to a file, **bsdccxxx.lst** where xxxx.seg is the input file name. Example of partial output:

```

Program bsdcc
Input File: c008.seg
Output File: bsdcc008.lst
Number of traces= 48
Parameters: none
  Trace      DC %(MAV)      MAV
-----
   1         0.23      0.4653721E+04
   2         0.43      0.4331486E+04
   3         0.13      0.3471209E+04
   4         0.08      0.2875869E+04
   5         0.11      0.2294398E+04
   6         0.06      0.1858094E+04
   7         0.10      0.1950509E+04
   8         1.44      0.5034012E+03
   9         0.11      0.1754470E+04
  10         0.15      0.1428828E+04
  11         0.06      0.1383624E+04
  12         0.12      0.1140977E+04

```

### 12.0.4 BRDC

BRDC removes a DC level from each trace, or a linear trend. The command line arguments are:

```

brdc infile iswdc usrdc

infile: name of input file to remove DC
iswdc:  0=user supplied DC level to remove
        1=determine dc level from each trace
        2=remove linear trend measured from trace
usrdc:  = user supplied value of DC to remove

```

Example of a partial output by **BSDC 12.0.3** after running this command:

```
brdc c008.seg 1
```

```

Program bsdcc
Input File: brdcc008.seg
Output File: bsdccbrdc.lst
Number of traces= 48
Parameters: none
  Trace      DC %(MAV)      MAV
-----
   1         0.00      0.4653861E+04
   2         0.00      0.4330840E+04
   3         0.00      0.3471210E+04
   4        -0.00      0.2875866E+04
   5        -0.00      0.2294333E+04
   6         0.00      0.1858083E+04
   7         0.00      0.1950462E+04
   8         0.00      0.5032268E+03
   9        -0.00      0.1754434E+04
  10         0.00      0.1428763E+04
  11        -0.00      0.1383601E+04
  12         0.00      0.1140975E+04

```

### 12.0.5 BINT

Integration of seismic traces. For example, if traces are in particle velocity, output will be in displacement. If traces are in acceleration units, output will be particle velocity. If data are clipped, integration will reveal a DC level by trace drift. Figure 73 shows a plot after integration with the command:

```
bint c008.seg
```

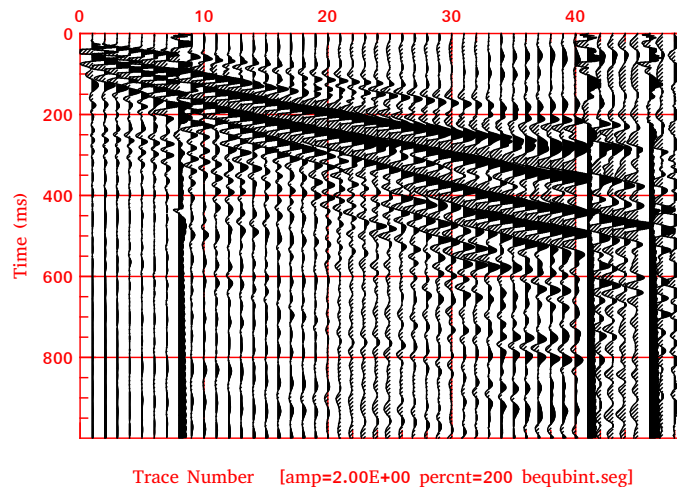


Figure 73: BINT: Integration of traces, plotted trace equalized with BEQU 12.0.9. Negative values grey, positive. DC levels are revealed by drift in either the positive or negative direction.

### 12.0.6 BSRT

Sorts data traces by offset.

```
bsrt infile isort

infile = input file name
isort +1= up by offset
      -1= down by offset
```

### 12.0.7 BRPT

Pre-trigger refers to the time before a trigger signal is received. Engineering seismographs can retain data continuously sampled before the trigger signal. This program shifts the data and resets the delaytime header in the shot header section. The only argument is the input file. The delay time header value is shown in this sample bdump.lst of a file with a pre-trigger (-10 ms pre-trigger).

```
-----
Length = 2000 samples          | Shot Elevation = 849.2
Sample Interval = 0.00025 sec. | Shot Depth = 0.0
Delay Time = -10 msec.         | Up Hole Time = 0 msec
Low Cut Filter = 0 Hz.         | Shot X-COORD = 9963.19
High Cut Filter = 1000 Hz.     | Shot Y-COORD = 10022.41
Line ID: 00X5                  | Shot Date (year.moday) = 2001.0417
Shot Orientation:              | Shot Time (hr:min) = 10:32
Azimuth= 0 Deg. Vertical=180 Deg. | Charge Size (grams)= 0
-----
TRACE|SHOT| STATION | OFFSET| RECEIVER          |VERT|1STBRK|K-GAIN|AZI|VER|
# |REC.|SHOT REC| | ELEV. X-COORD Y-COORD|FOLD|(SEC.)| (dB) | | |
```

**12.0.8 BDIF**

Data are differentiated with BDIF. Thus, if the data are in units of particle velocity, then the output will be in units of acceleration. The code uses a Bi-linear Transform to compute the derivative.

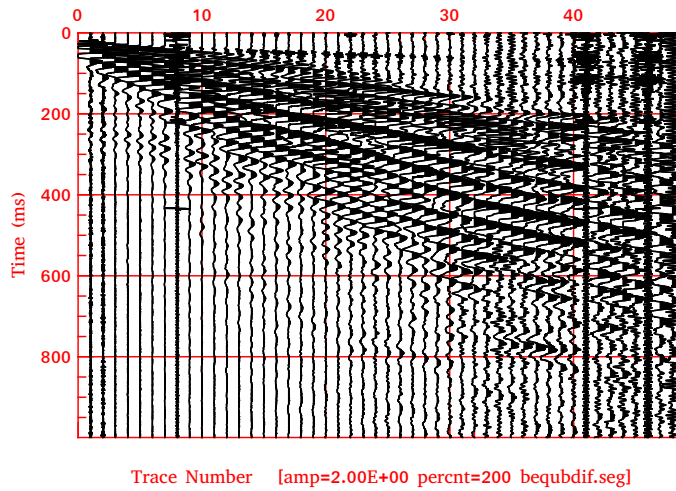


Figure 74: BDIF: Differentiation of BSEGY data, plot trace equalized with BEQU 12.0.9.

**12.0.9 BEQU**

Data amplitudes are rescaled by the L2 norm or the peak absolute value. Command line arguments are:

```

bequ infile tmin tmax normsel
  infile = input file name
  tmin   = gate: minimum time (s)
  tmax   = gate: maximum time (s)
  normsel = select normalization
          2= L2 Norm
          0= Peak abs(Value)
    
```

NOTE: Default is normsel=2  
 No interactive prompt for normsel  
 (Must be specified on command line)

Figure 75 illustrates how BEQU helps compensate for the wide range of amplitudes in seismic data.

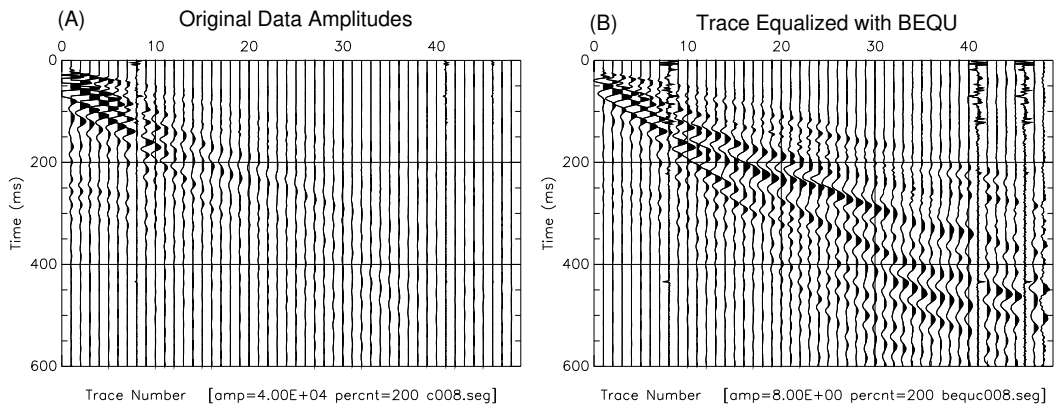


Figure 75: BEQU: (A) original scaling of data, (B) trace equalized with L2 norm. The scale factors for plotting are 40000 for (A) and 8 for (B).

### 12.0.10 BSCL

The program can scale a data set by a user provided value, 1/L2 norm, (trace,amplitude) pairs in a file, ampfil, or by the maximum absolute value in the file. The scale factor is found by scanning a limited number of traces defined by itr1 and itrn.

```
bscl infil itr1 itrn isw1 [ scaler | ampfil ]
```

```
infil:  =input file name
itr1:   =starting trace for determination window
itrn:   =number of traces to include in window
```

```
isw1:   0=user supplied scale factor
        1=scale factor from 1/L2 norm
        2=input file with (trace,ampfactor)
        3=scale factor from Max Abs Value
```

```
scaler: =user supplied scale factor (ONLY isw1=0)
```

```
ampfil: file name for option (ONLY isw1=2)
```

```
-----
| Converting microvolts to m/s particle velocity |
|
|           VSP Ref.Phone 28Hz Oyo
| scalar=4.0978E-8 28Hz Oyo SMC 28-720
| scalar=5.6497E-8 14Hz Oyo Phone
| scalar=5.0761E-8 10Hz Oyo GS-20DM Phone
| scalar=3.2787E-8 10Hz Mark L10-A Phone
| scalar=3.2034E-8 08Hz Mark L10-A Phone
|
| strain=(part.vel.)/(wave phase vel.)
|
|-----
```

Example, use the first 5 traces closest to the source and determine the maximum absolute value, compute a scale factor so that sample with the MAV has a value of unity (1), then apply to all the traces.

```
bscl c008.seg 1 5 3
```

Figure 76 shows the first 10 traces for clarity.

```
bplt bsclc008.seg 4 0 0 1 10 0 .6 1 1 200
```

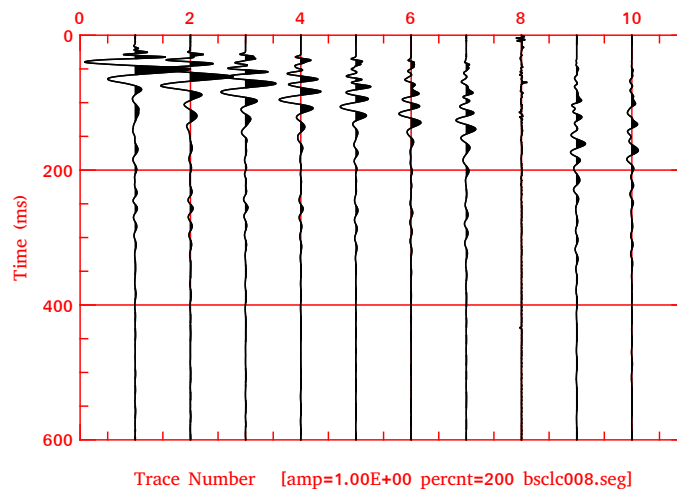


Figure 76: BSCL: Scale all traces by the maximum absolute value (MAV) found in the first 5 traces.

```
Peak Absolute Value=      192440.8750000000
```

```
-----
Maximum Value= 0.1924E+06   Trace #= 1
                          Sample #= 101
Minimum Value= -.1766E+06   Trace #= 1
                          Sample #= 80
```

```
-----
Scale Factor= 0.5196401E-05
```

**12.0.11 BGAR**

Computes an amplitude decay envelope over a user provided **RANGE** interval. The envelope is corrected for spherical divergence, converted to decibels, and a linear fit performed. The user may then apply the recommended gain correction, or over ride it with their own choice. The spherical divergence and exponential gain corrections are applied to the entire data set, (not just the interval of analysis). The data are not filtered before hand, so the decay measurements are a single result for the entire available bandwidth. If you want to measure inelastic decay as a function of frequency, then use program **BAMP 8.2.6**. This program simply provides a broad-band view of amplitude decay as sensed by the summed absolute value amplitude of each trace on a survey. A PostScript plot of the linear regression is output (requires **PLPLOT** package be installed)

```

bgar infile rmin rmax dbu

infile = input file name
rmin   = min. range design gate
rmax   = max. range design gate
dbu    = gain correction to apply (dB/m)
    
```

**NOTE:**

No prompt for dbu until after gain assessment.  
 However, you may specify dbu on the command line  
 if you already have a value you wish to use.

Example, bgar c008.seg 6. 100. .03

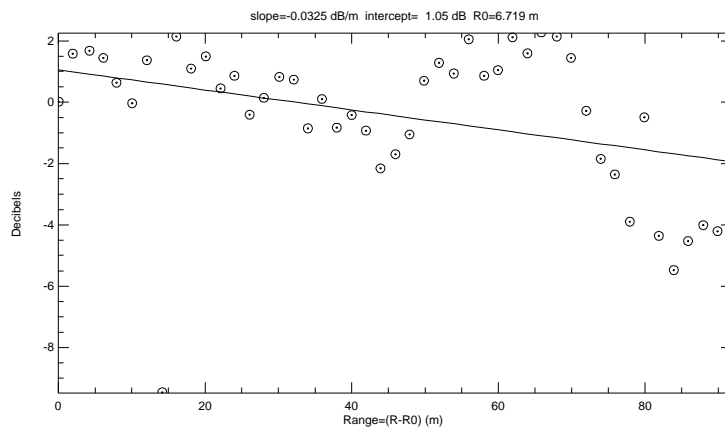


Figure 77: BGAR: Broadband scale by spherical divergence and exponential decay. Range from 6 to 100 meters.

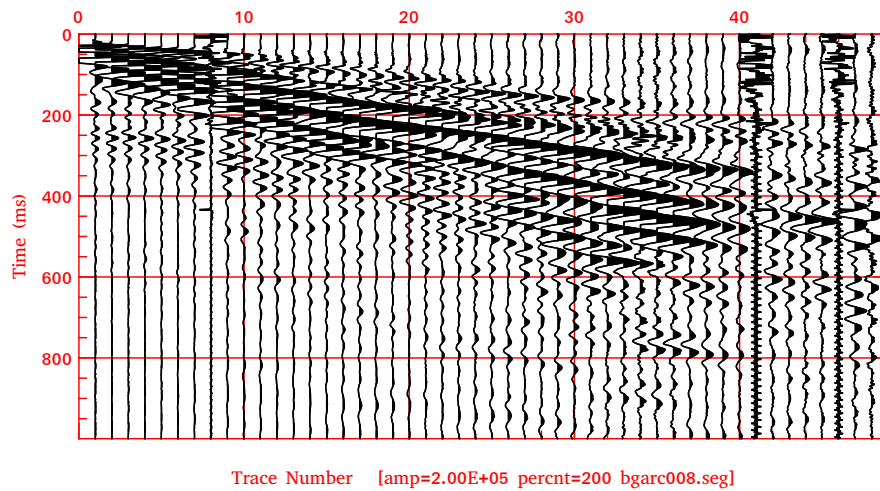


Figure 78: BGAR: Broadband scale by spherical divergence and exponential decay. Specified .03 dB/m for inelastic decay.

**12.0.12 BGAZ**

Computes an amplitude decay envelope over a user provided **depth interval**. The envelope is corrected for spherical divergence, converted to decibels, and a linear fit performed. The user may then apply the recommended gain correction, or over ride it with their own choice. The spherical divergence and exponential gain corrections are applied to the entire data set, (not just the interval of analysis). The data are not filtered before hand, so the decay measurements are a single result for the entire available bandwidth. If you want to measure inelastic decay as a function of frequency, then use program BAMP 8.2.6. This program simply provides a broad-band view of amplitude decay as sensed by the peak-peak amplitude of the direct arrival on a down-hole survey. A PostScript plot of the linear regression is output (requires PLPLOT package be installed).

```

bgaz infile zmin zmax dbu

infile = input file name
zmin   = min. depth design gate
zmax   = max. depth design gate
dbu    = gain correction to apply (dB/m)
    
```

**NOTE:**

No prompt for dbu until after gain assessment.  
 However, you may specify dbu on the command line  
 if you already have a value you wish to use.

Example, bgaz twave.seg 2. 20. 1.43

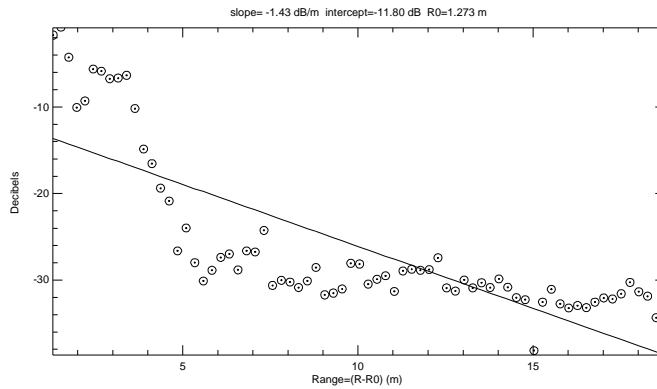


Figure 79: BGAZ: Broadband scale by spherical divergence and exponential decay. Depth range from 2 to 20 meters.

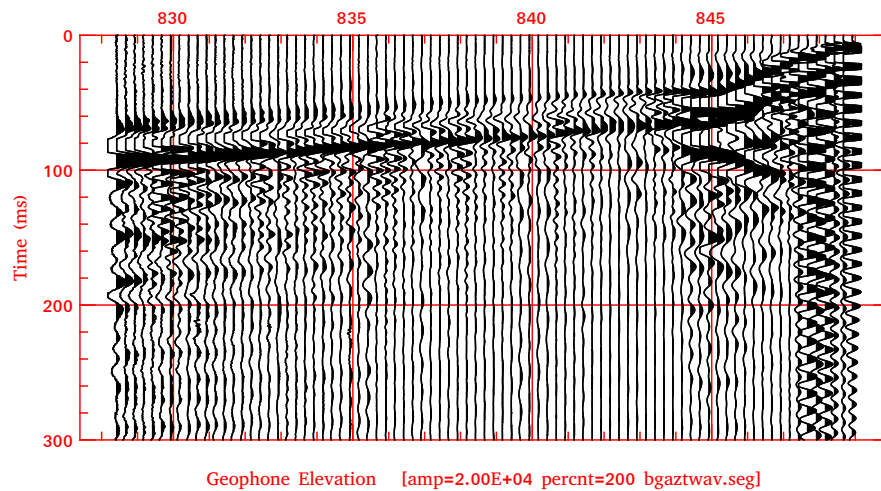


Figure 80: BGAZ: Broadband scale by spherical divergence and exponential decay. Specified 1.43 dB/m for inelastic decay. Elevations are down the bore-hole.

### 12.0.13 BAGC

Performs Automatic Gain recovery (both in space and time). One may choose to smooth the energy envelope with either a zero phase box car operator (which then gives an anticipatory component to the gain recovery), or one may choose to use the minimum phase (single pole on the real axis in the z-plane) filter. Output can be either the gain recovered data, or the smoothed gain recovery envelopes,  $\sqrt{\text{smoothed energy}}$ . First sample set to zero to avoid noise spike.

```

bagc infile twide itype
  infile = input file name
  twide  = width moving energy window (s)
  itype  = envelope smoother and output
          0= ARMA one-pole exp. decay
          1= zero-phase box car
          2= output ARMA envelope
          3= output BOXCAR envelope

```

Example zero-phase box car, bagc c008.seg .3 1

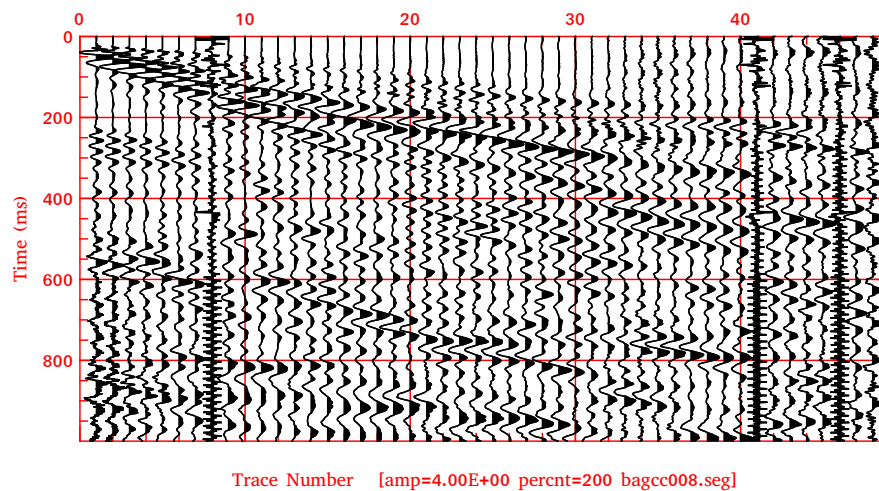


Figure 81: BAGC: Zero-phase boxcar 0.3 seconds.

Example single pole, bagc c008.seg .04 0

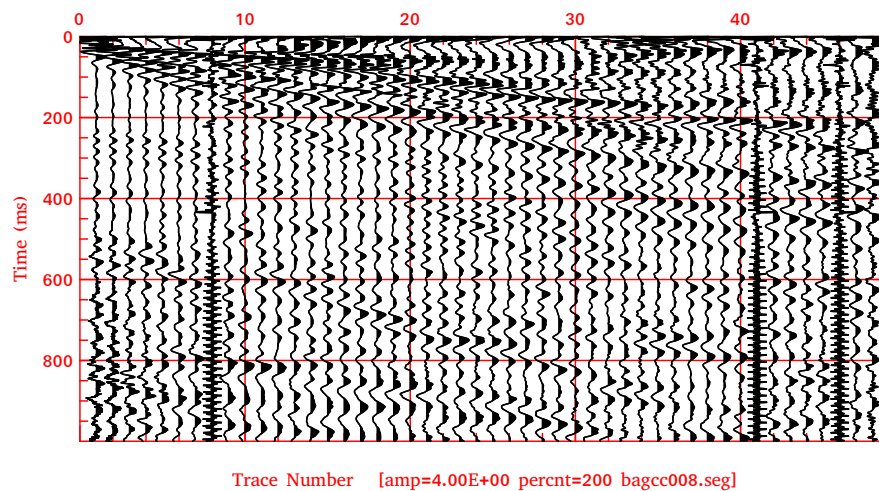


Figure 82: BAGC: Single pole AGC envelope .04 seconds.

### 12.0.14 BBAL

Balances two BSEGY files so that their Mean Absolute Values (MAV) are the same. Can be executed in either a trace or profile balancing mode.

```
bbal  infile1  infile2  iopt

infile1 = input file_1 name
infile2 = input file_2 name
iopt 0= profile mode
iopt 1= trace mode
```

Example that illustrates the concept is taken from down-hole data. There are two 3 component geophones, one at a depth of 19.39 meters, the other is a reference phone fixed at the surface elevation of the hammer.

```
-----
Length = 2500 samples          | Shot Elevation = 820.0
Sample Interval = 0.00020 sec. | Shot Depth = 0.0
Delay Time = 0 msec.          | Up Hole Time = 0 msec
Low Cut Filter = 4 Hz.        | Shot X-COORD = 9897.04
High Cut Filter = 1000 Hz.    | Shot Y-COORD = 10066.29
Line ID: 18A_                 | Shot Date (year.moday) = 1996.0604
Shot Orientation:             | Shot Time (hr:min) = 10:52
Azimuth= 90 Deg. Vertical= 90 Deg. | Charge Size (grams)= 0
-----
```

| TRACE # | SHOT REC. | STATION REC. | OFFSET | RECEIVER ELEV. | X-COORD | Y-COORD  | VERT FOLD | 1STBRK (SEC.) | K-GAIN (dB) | AZI | VER |
|---------|-----------|--------------|--------|----------------|---------|----------|-----------|---------------|-------------|-----|-----|
| 1       | 10        | 002 517      | 19.39  | 800.72         | 9897.04 | 10067.79 | 10        | 0.0000        | 60          | 0   | 0   |
| 2       | 10        | 002 518      | 19.39  | 800.72         | 9897.04 | 10067.79 | 10        | 0.0000        | 60          | 189 | 90  |
| 3       | 10        | 002 519      | 19.39  | 800.72         | 9897.04 | 10067.79 | 10        | 0.0000        | 60          | 279 | 90  |
| 4       | 10        | 002 520      | 1.59   | 819.96         | 9897.04 | 10064.70 | 10        | 0.0000        | 20          | 0   | 0   |
| 5       | 10        | 002 521      | 1.59   | 819.96         | 9897.04 | 10064.70 | 10        | 0.0000        | 20          | 0   | 90  |
| 6       | 10        | 002 522      | 1.59   | 819.96         | 9897.04 | 10064.70 | 10        | 0.0000        | 20          | 270 | 90  |

The first 3 traces are separated to a new file, as are the last 3 traces. In this example, we then do a trace balance between the down-hole and the reference phone traces and recombine as in Figure 83. The commands are:

```
bplt c010.seg 2 0 0 1 7 0 .25 1 2e4 200
mv bplt.fig c010.fig
bedt c010.seg 0 .5 1 3 1 0
mv bedtc010.seg down.seg
bedt c010.seg 0 .5 4 6 1 0
mv bedtc010.seg refn.seg
bbal down.seg refn.seg 1
cp bbaldown.seg BBAL.seg
cat bbalrefn.seg >> BBAL.seg
bplt BBAL.seg 2 0 0 1 7 0 .25 1 2e4 200
mv bplt.fig BBAL.fig
```

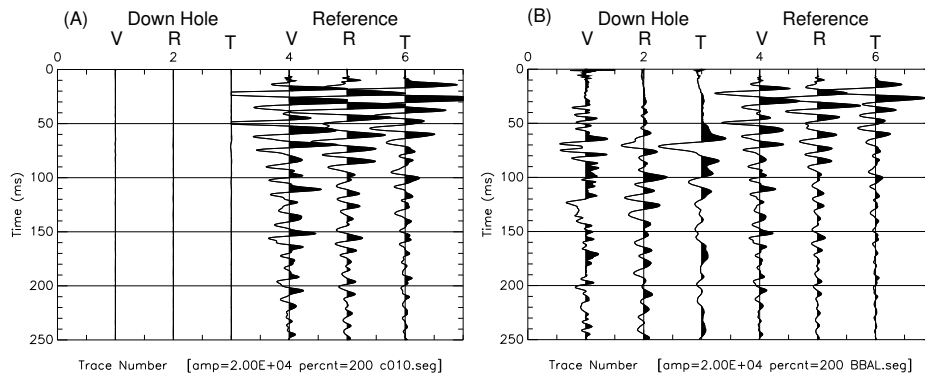


Figure 83: BBAL: (A) Original data (down-hole barely visible) (B) data after splitting the data into two files, running BBAL, then combining into a second file.



### 12.0.15 BSTK

Stacks all the traces in a gather, outputs the same trace repeatedly, number in = number out. Command line argument is the input file name. Figure 84 (A) shows the reference phone recording for each source effort of a down-hole survey. While repeatable, there is some variation as the source compacts the ground. (B) shows the average of all the source efforts estimated by the sum of all the traces in (A). The summing is called a stack.

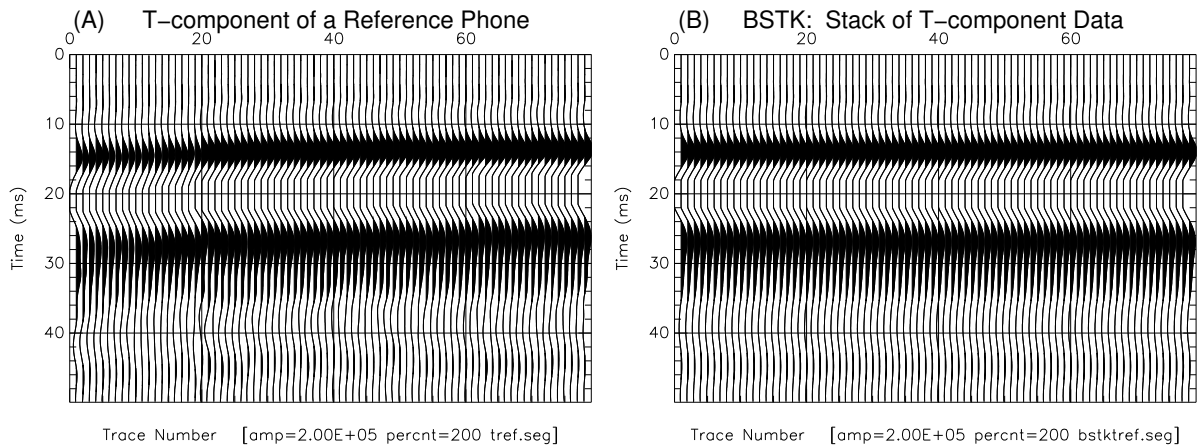


Figure 84: BSTK: (A) Original data T-component data (B) Stack of the T-component data (all traces replicas of the stack result).

### 12.0.16 BXCR

The command line arguments are:

```
bxcr infile1 infile2 t1 t2 tlagmx

infile1: name of input file #1
infile2: name of input file #2
t1:      =start time of cross correlation gate (sec.)
t2:      =end time of cross correlation gate (sec.)
tlagmx: =maximum cross correlation lag time (sec.)
```

If the second input file is the same as the first, the result will be an auto correlation. If the input files are different, then the result is a cross-correlation between the two, and the order of the file names is important when looking at relative time shifts.

Example: Auto correlation is computed for data shown in Figure 75.

```
bxcr c008.seg c008.seg 0 1.2 .25
```

Both the auto correlation and the stack of the auto correlation are shown in Figure 85. The stack presents an average of the auto correlations at each offset. In (A) of Figure 85 we see that the near offset data (on left of the figure) present a broader bandwidth than at the further offsets. The spectral computation of the stack will provide an average spectrum, while spectral computations of the simple auto correlation in (A) will show the change in bandwidth with offset.

The zero lag sample is at the middle time. In this case, sample time of 125 msec. corresponds to zero lag (125 msec is 1/2 of 250 msec). The command above took all 1.2 seconds of data and computed the auto correlation out to  $\pm 125$  msec.

To compute an all pole spectrum, see **OCTAVE YULEWALKER** in section 6.0.7.

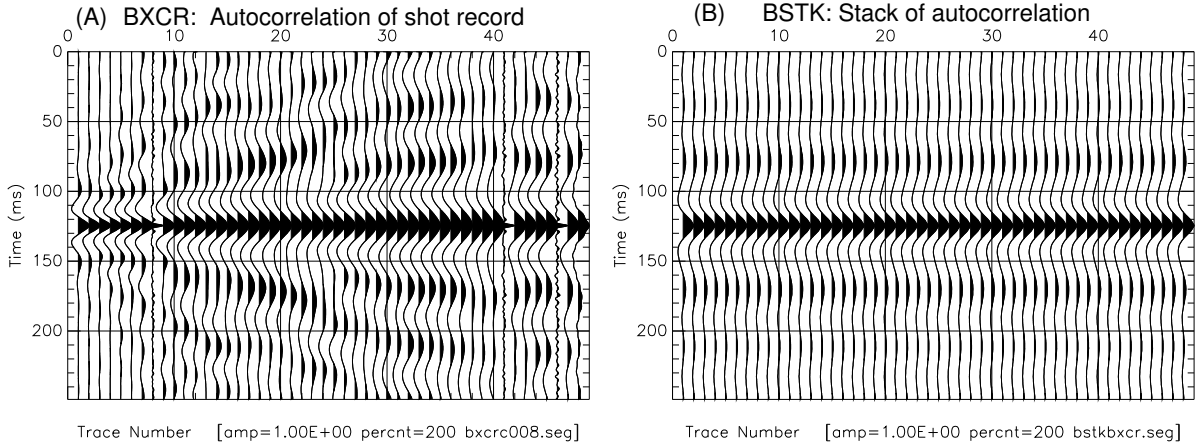


Figure 85: BXCR: (A) Auto correlation of data shown in Figure 75 (B) Stack of the auto correlation (all traces replicas of the stack result).

12.0.17 BNOS

Computes band-limited random noise which copies headers from a template \*.seg file. Noise can be added back into the template file with **BSUM** 12.1.4.

```
bnos infil seed F_Low F_High F_Roll

infil  =name of a file to match: npts,ntraces...
seed   =seed for random noise
       (positive number less than 1)
F_Low  =low-cut frequency for random noise, Hz
F_High =high-cut frequency for random noise, Hz
F_Roll =roll off in Hz (trapezoidal pass band)
```

Example: bnos k007.seg .91827364 10.0 100. 5.0  
See Figure 86.

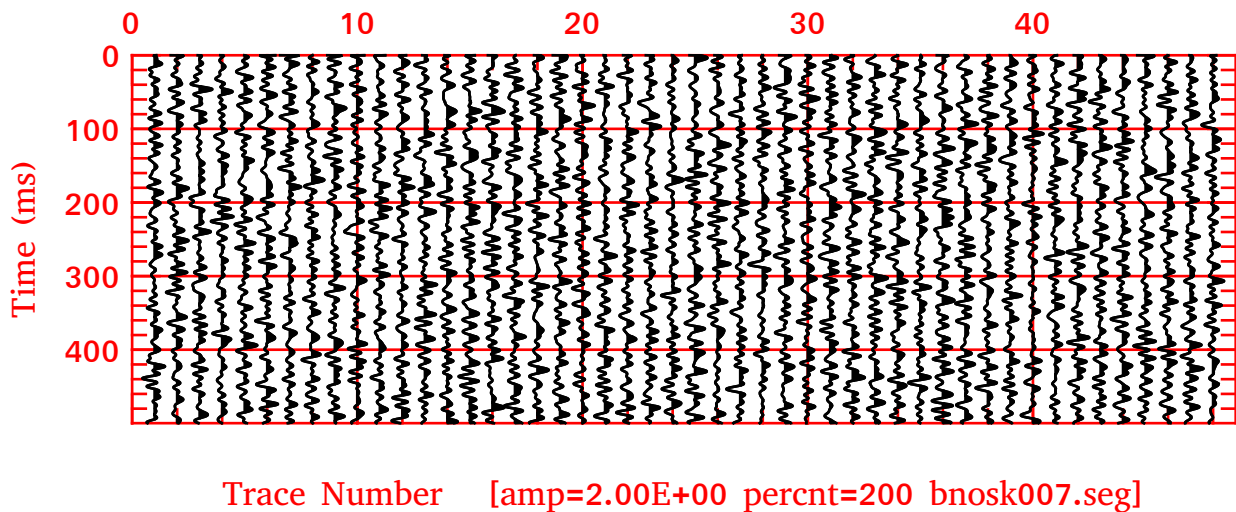


Figure 86: BNOS: Band-limited noise, 10-100 Hz.

## 12.1 Down-hole VSP Processing for Reflections

The following signal processing applications are included as a set to illustrate how a Vertical Seismic Profile (VSP) may be processed for reflections. The steps are:

1. **BGAZ 12.0.12** gain correction:  
bgaz twave.seg 5 20 0.6
2. **BSHF 12.1.1** flatten data on direct arrivals + 20 ms.  
mv bgaztwav.seg 00X5.seg  
bshf 00X5.seg 0 1 .02
3. **BMED 12.1.2** median mix of flattened data to extract down going wave.  
bmed bshf00X5.seg 15
4. **BSUM 12.1.4** subtract direct wave from total wavefield.  
bsum bshf00X5 bmedbshf.seg -1.0
5. **BSHF 12.1.1** restore to 1-way time.  
bshf bsumbshf 0 0 -.02
6. **BSHF 12.1.1** shift to 2-way time to flatten reflections (adding the direct arrival times again)  
bshf bshfbsum 0 0 0.

### 12.1.1 BSHF

All samples move in time by a constant shift. The shift in seconds is either in the header for first break pick, or in a separate file. See Figure 87 for the example.

```
BSHF infile ipic ishf tshift picfil

infile:  =input file name
ipic:    1=static shift data by picks pick file
         0=static shift data by picks in headers
ishf:    0=add static shifts
         1=subtract static shifts
tshift:  =bulk static to add to picks
picfil:  =file name with picks (for ipic=0)
```

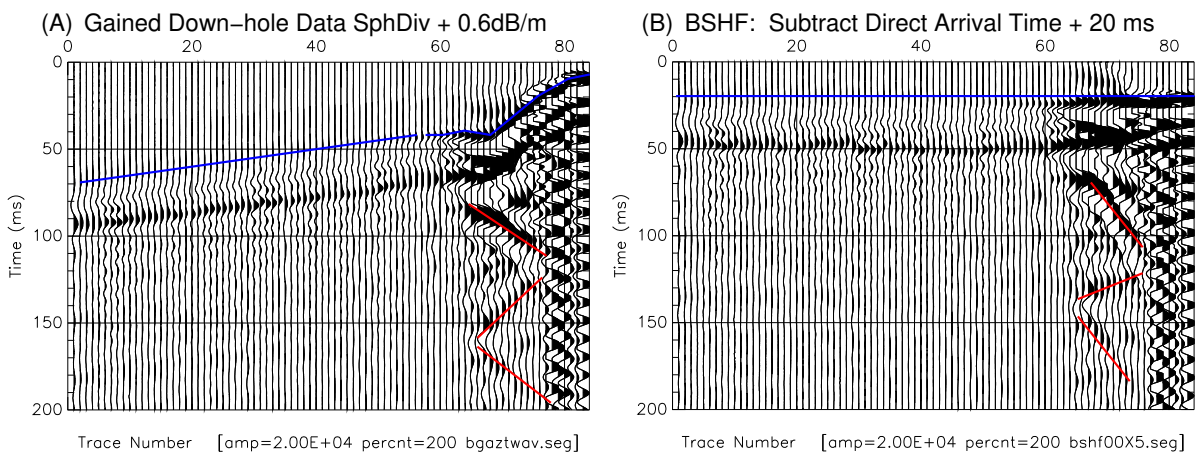


Figure 87: BGAZ: (A) Gained down-hole data, blue=direct wave, red=reverberating reflections (B) BSHF: Data flattened on down-going wave.

### 12.1.2 BMED

A median mix is usually preferred since it is less likely to smear large amplitude, often noise spikes. As an alternative, one could use the mean mix, **BMIX** 12.1.3 program. One should use an odd number of traces in the mix.

```
bmed infile mix

infile =input file name
mix    =mix width <21
```

### 12.1.3 BMIX

Mean mix, only difference between median and mean mix in wave field separation is which value (mean or median) is used in the moving average operator. The mean is not used in this example.

### 12.1.4 BSUM

The median mix is an estimate of the down-going wave in this example. When subtracted from the total wave-field data, the result should be up-going waves. See Figure 88 (B) .

```
bsum  infile1  infile2  scalef
      infile1 = first input file name
      infile2 = second input file name
      scalef  = scale factor
      output  = input1 + scalef*input2
```

The up-going wave field estimate is then shifted back to 1-way time. Then the data are shifted again by the direct wave down-going times (this time without any bulk shift, to adjust the data to 2-way time. The reflections should be flattened in 2-way time (see Figure 89).

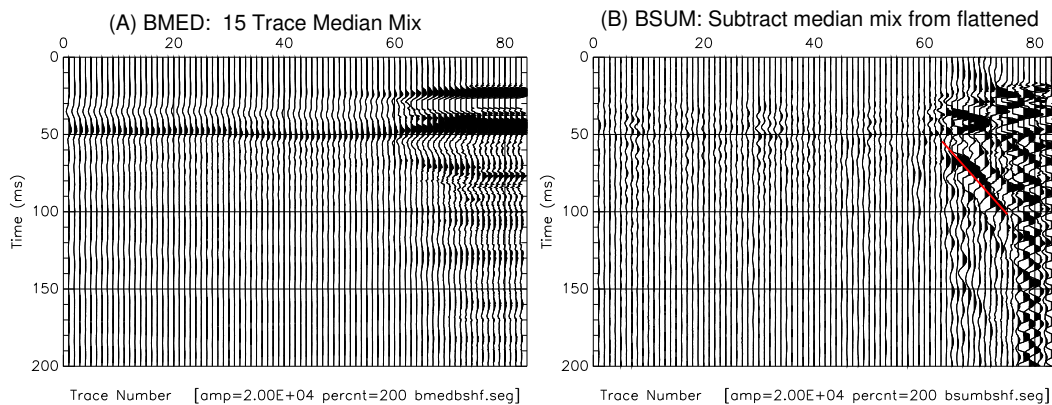


Figure 88: BMED: (A) median mix of the direct wave (see figure 87 B) (B) BSUM: direct down-going wave estimate subtracted from total wave field.

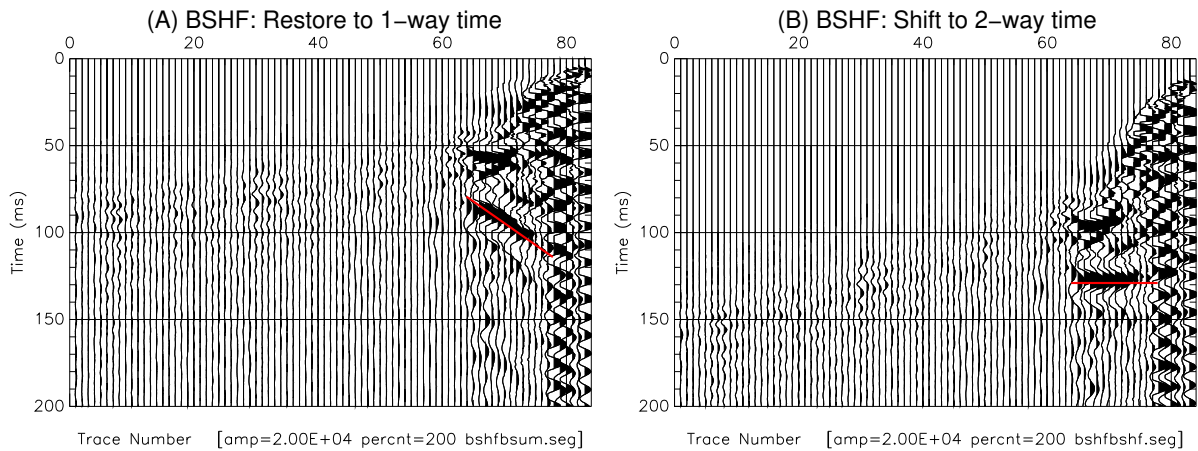


Figure 89: BSHF: (A) median mix of the direct wave (see figure 87 B) (B) BSUM: direct down-going wave estimate subtracted from total wave field. Data in 2-way time.

## 12.2 Additional Down-hole Processing

The following signal processing programs are often used in processing down-hole data, as well as in other circumstances.

- **BSHP** 12.2.1 Wiener least square shaping filter.
- **BTOR** 12.2.2 Applies PCA analysis to headers.
- **GENBROT** 12.2.3 Generates a bash script for **BROT**.
- **BROT** 12.2.4 Actually rotates horizontal component data in a down-hole survey.

### 12.2.1 BSHP

Wiener least squares shaping filter design and application to standardize an embedded wavelet. When used in a down-hole survey, it can be used to remove variations in triggering and bandwidth from repeated source efforts from the down-hole data. The design is to find a filter which matches each source effort to a target trace on the reference phone. This way, we remove variations in the source effort from appearing in the down-hole data. This removal occurs when we re-apply the designed filters on the reference phone to the down-hole data, thus standardizing the embedded wavelet.

```
bshp infil infil2 iswopt iswch tmin tmax npf sf infil3
```

```
infil :   =name of input file #1
         (file1*filter=file2)
infil2:   =name of input file #2
iswopt:   0=profile mode
         1=trace mode
iswch:    1= (file#1*filter)=output
         0= (file#3*filter)=output
tmin:     =start time of design gate in seconds
tmax:     =end time of design gate in seconds
npf:      =number of point in filter
sf:       =stability factor (.001 typical)
infil3:   =name of input file #3 (iswch=0 only)
```

Example: If **tref.seg** is the T-component of the reference phone, and if **targ.seg** is a target trace (perhaps the last of the source efforts as recorded on that same component, the a shaping filter can be found that matches each source effort to that last effort. Rational is that the last source effort is stable due to the compaction of soil below a hammer source. The command to match each source effort to the target trace might be:

```
bshp tref.seg targ.seg 1 1 0 0.1 360 .0001
```

The output file would be **bshptref.seg** and should be plotted to asses the degree of success and the chosen command line arguments. Then application of the filter designed above to the down-hole data might be done with this command:

```
bshp tref.seg targ.seg 1 0 0. 0.1 360 .0001 twav.seg
```

where **twav.seg** is the file with the original down-hole data which are contaminated by variations in trigger source timing and embedded source wavelet. What happens is that the shaping filters are recomputed with the same design input, but applied this time to the file listed as the last argument on the command line. The shaped down-hole data would be the output file, **bshptwav.seg**.

The degree to which shaping is helpful depends on how repeatable the source efforts are. With a highly stable and repeatable source, there will not be much difference in the result from shaping. However, with a source that produces variation in triggering or wavelet radiated, the result may be very helpful. Shaping will not hurt unless significantly poor choices are made in the command line parameters.

### 12.2.2 BTOR

Applies azimuth and vertical angles to geophone trace headers from a **bhod.lst** file. The command line arguments are:

```
btor lstfil, prfx, isw1 maxtr

lstfil  =input list file name (ex. bhod.lst)
prfx    =*.seg file prefix (one character)
isw1    =up/down switch
        -1=apply to *.lst file and one less
        +1=apply to *.lst file and one more
        0=IF VERTICAL IMPACT source
maxtr   =maximum number of traces in shot record
        6= 3 components down-hole, 3 ref-phone
        7= 3 down, 3 ref-phone, 1 load cell
```

#### EXAMPLES:

```
[if file number (col. 1 of bhod.lst) is 005
and isw1=-1, the azimuth and vertical angle
also applied to file 004]
```

```
[if file is 005 and isw1=+1, azi and vert
applied to 006 also]
```

A summary of the flow is this:

- **GENBHOD 10.1.9** creates bash scripts to run on down-hole data acquired from a horizontal component source, two blow orientations per subsurface station. The output is a file, **bhod.lst**.
- **BTOR** Reads the **bhod.lst** file and applies the determined phone orientations to the headers.
- **GENBROT 12.2.3** creates scripts to run **BROT**.
- **BROT 12.2.4** runs the script to actually rotate the data to a standard orientation.

**12.2.2.1 Example of BTOR** Consider a single depth station for illustration. There may actually be 100 or more depth stations in a single down-hole survey. There are two files in this example:

- **c009.seg** Source orientation is azimuth 270 degrees, 90 degrees from vertical (ie. horizontal blow West).
- **c010.seg** Source orientatio is azimuth 90 degrees, 90 degrees from the vertical (ie. horizontal blow East).

The steps are:

1. **gobhodo** This generates the difference between scaled versions of the two source efforts. The scaling is done on the vertical component of the down-hole phone (ch 1 on the author's wiring). File 9 is subtracted from 10. The difference file is renamed as h010009.seg
2. **gorunbhod** Program bhod is run to analyze file h010009.seg and produces files: h0010.plt.ps, bhod.lst  
These are the hodogram plot and a file with the determined phone orientations (R and T downhole)  
The command in the script for this depth is:  
bhod h010009.seg 2 3 50 90.0 180.0 +90.0  
Ch 2 is R and Ch 3 is T component downhole. 50 percent max amplitudes used in analysis 90 deg is source azimuth (ie E-W) and bowspring, R-phone observation is close to 180 degrees. The downhole phone is wired for +90 degrees between R and T components
3. **BTOR** This program inserts the orientations of the phone azimuths and vertical orientations into the headers. The command in the script for this depth is:  
btor bhod.lst c -1 6  
*This command will process ALL the cxxx.seg files in the directory (subject to be included in the **bhod.lst** file).*

## Renaming btorxxxx.seg files to xxxx.seg

A script to rename the **BTOR** files in a directory is as follows:

```
#!/bin/sh
#Script to rename files after btor process
#overwrite pxxx.seg files, p=prefix
# Author: P. Michaels    Date:April 2002    See GNU License

if test "$1" = ''
then
    echo 'Enter 1 character prefix'
    echo 'Example: w'
    echo ' for files btorw001.seg, btorw002.seg, etc...'
    read PRFX
else
    PRFX=$1
fi

find -name "$PRFX*.seg" | \
sed s/'\.\./'/' '/g | \
gawk '{print "mv","btor"$1,$1}' \
>go-rename
chmod +x go-rename
./go-rename
echo "btor files renamed"
```

### 12.2.3 GENBROT

Once the \*.seg files have had their headers updated with the geophone orientations, we can rotate the data so that the horizontal components face in a standard direction. In down-hole surveys, as the tool is dragged up the hole, it can slowly rotate. In some cases, the tool may become stuck, have to be unclamped and then reclamped, resulting in tool spin. This program is interactive and generates a bash script to apply a rotation of the data so that one component is parallel to the source azimuth (assuming an SH-wave source is used). An example log of a run follows:

Enter alpha prefix (char) of \*.seg data to be rotated

EXAMPLE: if enter 1, then files 1001.seg to 1010.seg  
would be processed if sequence  
numbers 1 and 10 entered next

L

```
...L
Enter first file number to process
1
Enter last file number to process
146
Output in file==>gobrot
```

The generated script file, **gobrot**, will then look like this:

```
brot L001.seg 2 3 1
brot L002.seg 2 3 1
brot L003.seg 2 3 1
.
.
.
brot L144.seg 2 3 1
brot L145.seg 2 3 1
brot L146.seg 2 3 1
```

Of course, one must then make the **gobrot** file executable:  
chmod +x gobrot



### 12.2.4 BROT

One runs the **gobrot** bash script and this produces files **brotL001.seg** through **brotL146.seg** in this example. The command line arguments are:

```

brot infil itrx itry iopt iangle

infil:  =input file name
itrx:   =trace number for x-axis (example, 2)
itry:   =trace number for y-axis (example, 3)
iopt:   =rotation option
        0=user supplied rotation angle
        1=Transverse Alignment: chan 3 aligned to 6
        2=Radial Alignment:   chan 2 aligned to 5
iangle: =user supplied rotation angle from x-axis

```

One only needs to add an **iangle** parameter with the **iopt=0** option. For each \*.seg file rotated, there will also be a \*.lst file output. The \*.lst file shows what **iangle** value has been used based on the headers for options 1 or 2.

**NOTE:** BSU codes like this one assumes that the channel order in down-hole surveys matches those of the author. See section 6.7.2 of the BSU Users Guide ([bsu-user-guide3-1.pdf](#)) for more on this topic. Figure 90 illustrates the author's notation and wiring, and is taken from the BSU user guide. A discussion on Principal Component Analysis (PCA) is found in the literature ([Michaels, 2001b](#)).

Once the **gobrot** script is run, the rotated data will have names **brotL001.seg** through **brotL146.seg** in this example. A good practice is to create a child directory, **brot** and move the **brotxxxx.seg** files to that directory before doing further analysis. This will preserve clarity on which files have been rotated, and which files are as recorded.

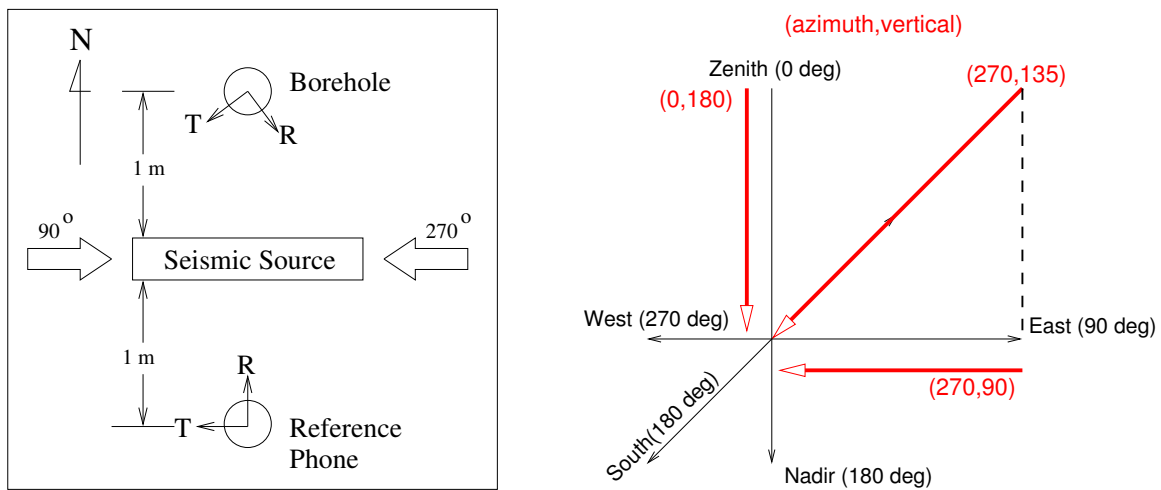


Figure 90: The author's orientations and notation for down-hole surveys. Note that the reference and bore-hole phones are wired differently (in terms of R- and T-component wiring).

## 12.3 FILTER Codes

### 12.3.1 BFXT

The Frequency-Distance (FX) transform may be computed for a shot gather. The output files are **bfxtamp1.seg** and **bfxtphaz.seg** if a forward transform is computed. The command line arguments are:

```
bfxt infile1 iopt infile2
```

infile1 =name of X-T input file (space,time)

```
iopt    1=forward transform (unwrap phase)
        2=forward transform (no unwrap phase)
       -1=inverse transform (two input files)
```

infile1 =name of input file AMPLITUDES(iopt=-1 Only)

infile2 =name of input file PHASE(iopt=-1 Only)

NOTE: Forward transform requires only one input file  
Inverse transform requires two input files

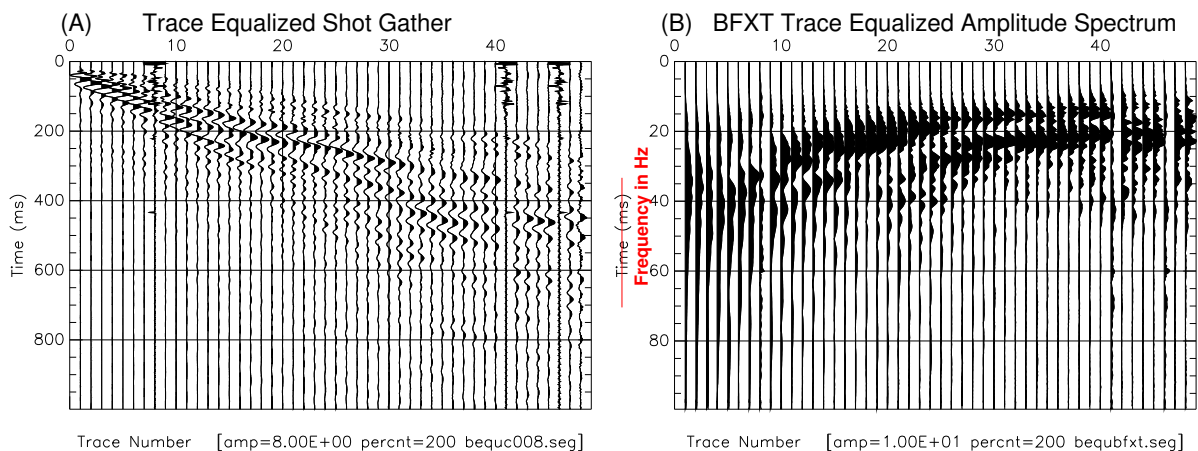


Figure 91: BFXT: (A) trace equalized shot gather using BEQU 12.0.9 (B) the amplitude spectrum after equalization with BEQU. Not shown is the phase transform.

The example shown in Figure 91A shot gather has a sample interval of  $\Delta t = .0005 \text{ sec}$  and 2500 samples per trace. The code uses a Radix 2 FFT, and in this case the sample interval is modified and there are 2048 samples per trace. The Figure 91B plot has to be relabeled since we are using BSU plot program **BPLT** here, and that program is limited to assuming all data are in time. A frequency axis replaces the time axis, and frequencies run from zero to the Nyquist. It appears to the headers as if the maximum sample is at a time of 1.0. In actual fact, the maximum sample is at 1000 Hz. Some scaling is going on to make plotting easier.

So Figure 91B is plotted to a maximum of 0.1 which turns out to be 100 Hz. So what is going on? BFXT calls a subroutine, **nrad2.f** which computes the first power of two larger than the number of samples in the shot gather, call it  $N2$ . A frequency domain sample interval is computed on this larger number of samples (the code pads with zeros to fill it out). Thus,  $\Delta f = 1/(N2 \cdot \Delta t)$ . But because we are dealing with time domain codes for other things we might do, we scale the sample interval, dividing it by 1000. Thus a Nyquist of 1000 Hz (maximum sample frequency in FX domain) becomes 1.0, as if it were 1.0 seconds. When going back into the time domain (TX), all this is reversed.

### 12.3.2 BCAR

This is a high-speed filter based on a moving average box car operator. It can do smoothing (ie. low pass) or high-pass filtering by subtracting a low pass result from the original data. For most applications, BSU has better filters (parameters in frequency rather than time), but this is quick and dirty, and is specified in time duration. Auto-Regressive-Moving-Average (ARMA) filtering can be done with BFIL [12.3.3](#). The command line arguments for BCAR are:

```

bcar infile isign ntimes twide

infile   =input file name
isign    -1=low-pass box car
          0=high-pass (1-sincntimes)
          +1=high-pass (1-sinc)ntimes
ntimes   =number of times to apply boxcar
twide    =width of box car in seconds

```

An example is shown in Figure [92](#) where both a low and high-pass filter are demonstrated. The commands were: `bcar c008.seg -1 1 .1` for the low-pass filter and `bcar c008.seg 1 1 .1` for the high-pass filter

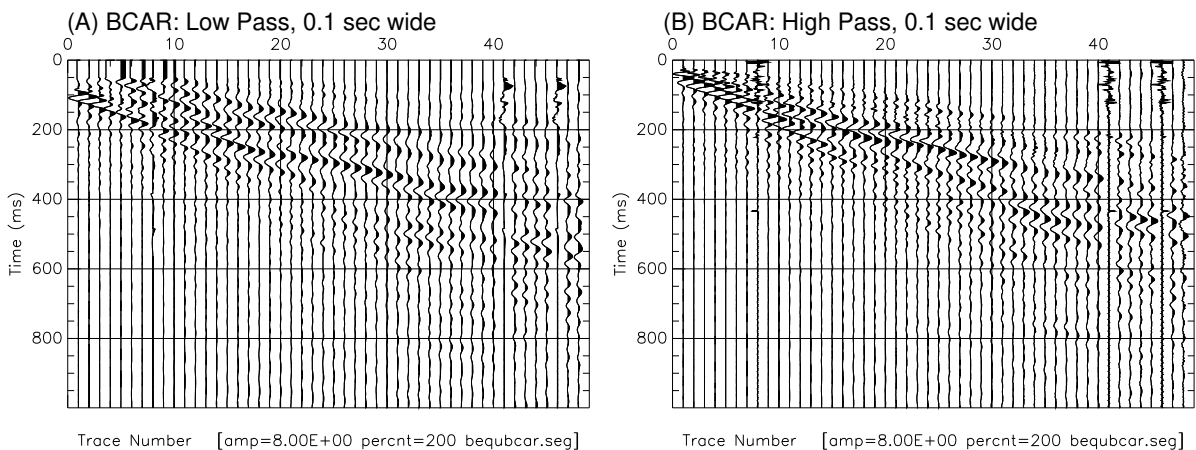


Figure 92: BCAR: (A) low-pass filter, trace equalized with BEQU [12.0.9](#) (B) high-pass filter by subtracting low-pass from original data, also trace equalized. Input data are same as in Figure [91A](#).

### 12.3.3 BFIL

BFIL uses a bilinear transform to perform ARMA filtering. Zero phase filtering is done by two passes of minimum phase filtering in opposite temporal directions. The command line parameters are:

```

bfil infile itype npoles fcenter bwidth ifaz

infile   =input file name
itype    0=low-pass filter, cut off freq= fc (-3dBv)
          1=band-pass filter, center frequency= fc
          2=high-pass filter, cut off freq= fc (-3dBv)
npoles   =number of poles in filter
          (6dB/octave)/(pair of poles)
fcenter  =center frequency Hz
bwidth   =band-pass filter bandwidth (-3dB) Hz
ifaz     1=minimum phase 0=zero phase filter

```

Examples of filtering with BFIL are shown in Figure [93](#):

- (A) **Low-Pass** Minimum phase, 12 Hz cut-off, 4 poles  
bfil c008.seg 0 4 12. 1
- (B) **High-Pass** Minimum phase, 48 Hz cut-off, 4 poles  
bfil c008.seg 2 4 48. 1
- (C) **Band-Pass** Minimum phase, 24 Hz center, 24 Hz band-width, 4 poles  
bfil c008.seg 1 4 24. 24. 1

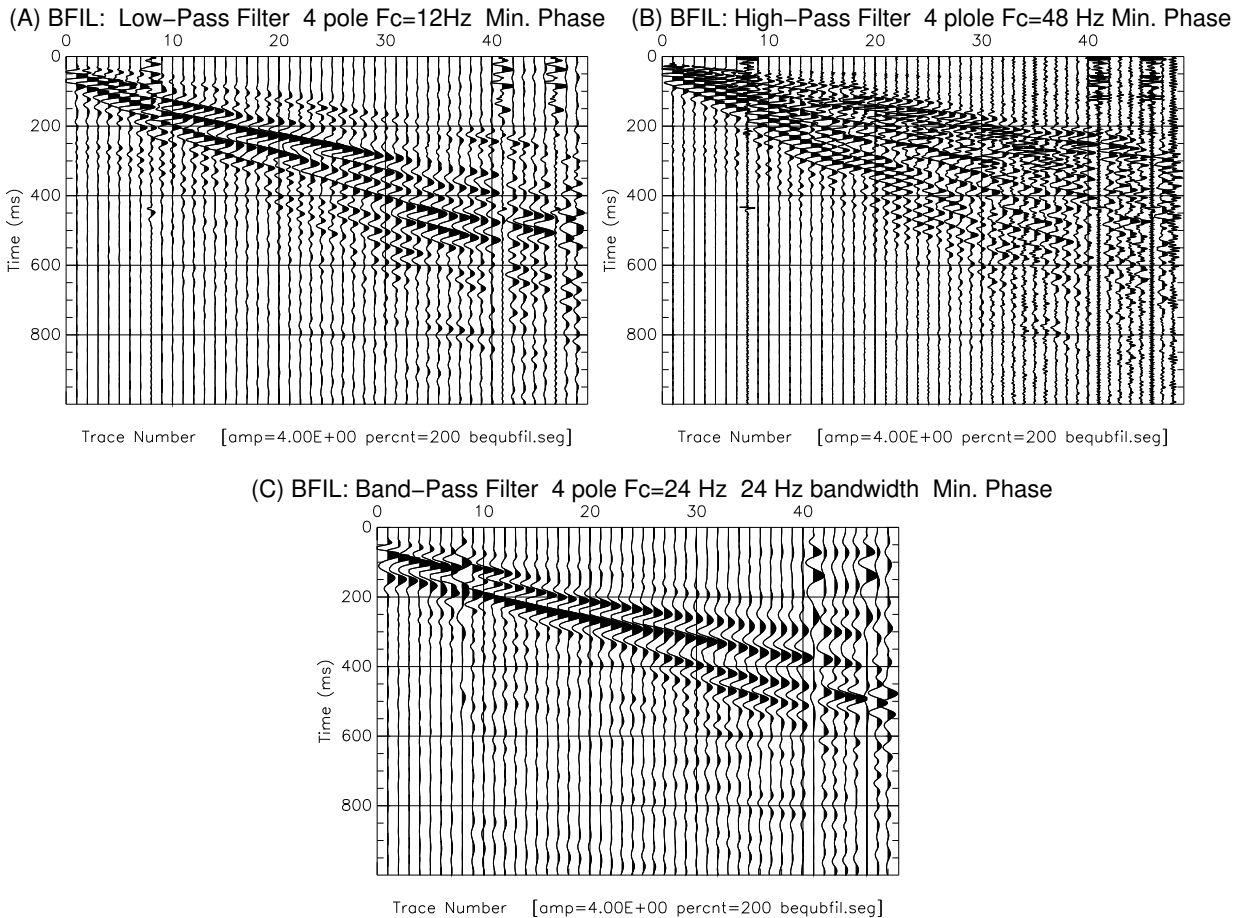


Figure 93: BFIL: Input data are same as in Figure 91A.

Another way to band-pass filter is to run the data twice, once through a low-pass, and then through a high-pass filter, choosing cut-off frequencies to produce a band-pass.

### 12.3.4 BDCN

Classic minimum phase spiking decon when the prediction error option is chosen. Intended for reflection data with random reflections and minimum phase wavelet, can be run in trace or profile mode. However, it can be run on other data as a whitening operator, your mileage will vary. Command line arguments are:

```

bdcn infile tmin tmax mpts stabf iprof imode
infile =input file name
tmin   =Autocorrelation Gate: START
tmax   =Autocorrelation Gate: END
mpts   =Length of Decon Operator
stabf  =Stability Factor (ex: 0.01)
iprof  1=profile mode 0=trace mode
imode  1=Prediction 0=Prediction error
        Choose 0 for spiking decon

```

Examples of BDCN are shown in Figure 94.

- **(A) Prediction Gate:**[0-1.2 sec] 30 sample (15 ms  $\Delta t = .0005 \text{ sec}$ ) operator, 0.1 stab factor, trace mode:  
`bdcn c008.seg 0 1.2 30 .1 0 1`
- **(B) Prediction Error, Spiking Gate:**[0-1.2 sec] 30 sample (15 ms  $\Delta t = .0005 \text{ sec}$ ) operator, 0.1 stab factor, trace mode:  
`bdcn c008.seg 0 1.2 30 .1 0 0`

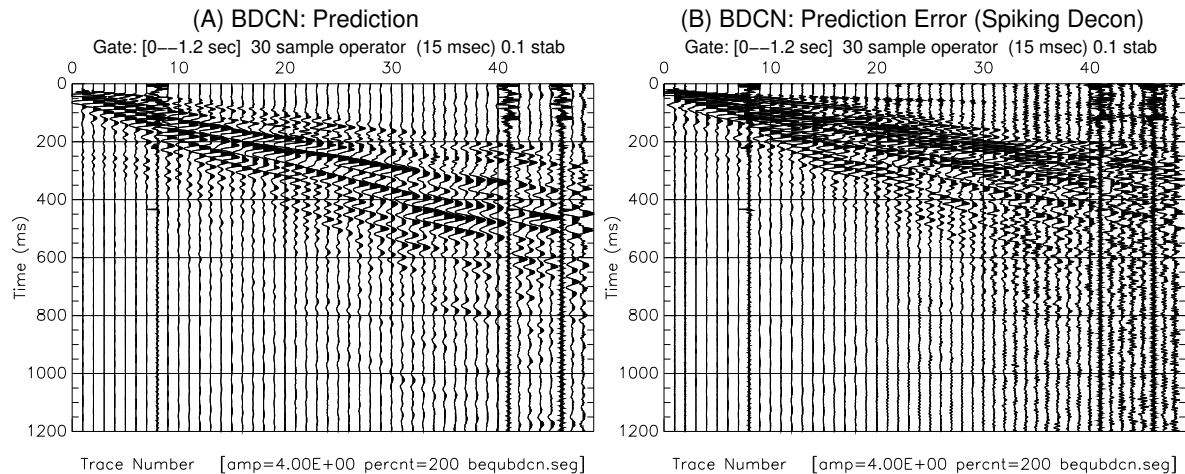


Figure 94: BDCN: Input data are same as in Figure 91A.

### 12.3.5 BFTR

One can either filter one data set with another, or filter using a namelist file (like the one produce with TRAPLT 6.0.1). The command line arguments are:

```
bftr  infil  iswf  filef

infil:  =input file to be filtered
iswf:   =filter source switch
        0=filters in *.SEG data set
        (one filter trace for each trace in infil)
        1=single filter specified in namelist file

filef:  =name of filter data set

-----Namelist Definitions-----
&FILTER  npf=number of points in filter
         nzph=sample for zero reference
         fil=f1,f2,f3,f4,...
         (values of filter samples)
&end
```

NOTE: if you get a core dump, you may have forgotten the &end at the end of the file

The following example shows how to generate filter traces and apply them.

```
# filter with low pass, 4 pole 12 Hz cut off minimum phase
bdum c008.seg .10
bfil bdumc008.seg 0 4 12. 1
bplt bfilbdum.seg 2 0 0 1 100 0.0 .4 1 2E-2 200

# apply low passed by convolving bfilbdum.seg with c008.seg
bftr c008.seg 0 bfilbdum.seg
bequ bftrc008.seg 0 1.
bplt bequbftr.seg 2 0 0 1 100 0.0 1. 1 2 200
```

The procedure:

1. BDUM creates a file with an impulse at 0.1 seconds, the template is the field data file, **c008.seg**. The output is **bdumc008.seg**.
2. BFIL filter the impulse file with a low pass filter, 4 pole, 12 Hz cutoff, minimum phase. Output is **bfilbdum.seg** Figure 95A.
3. BFTR filter **c008.seg** with the file, **bfilbdum.seg** Figure 95B.

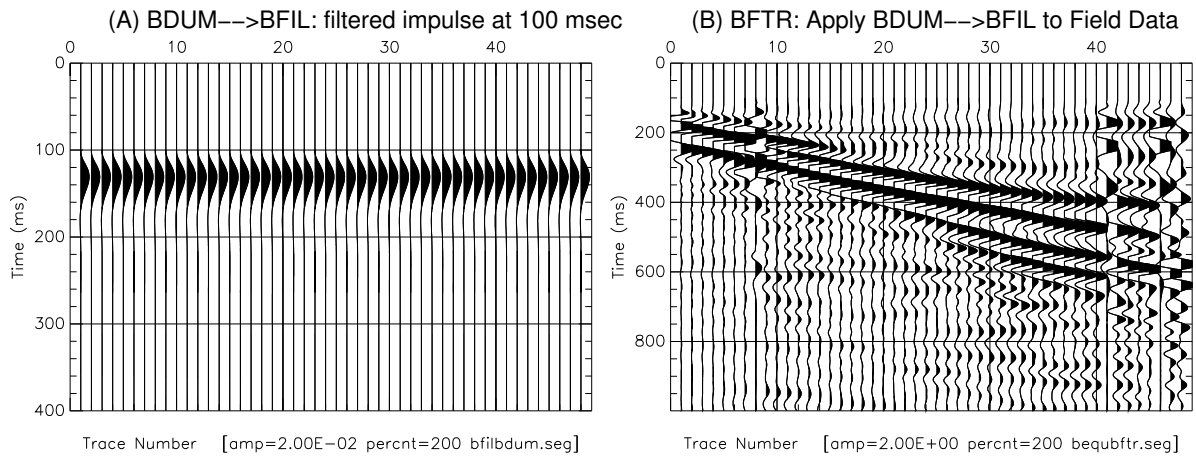


Figure 95: (A) BDUM-->BFIL: Filtered file of impulses. (B) BFTR: Filter field data with filtered impulse file. Input data **c008.seg** are same as in Figure 91A.

The other alternative would be to run **TRAPLT** on file **bfilbdum.seg** and copy the namelist section to a file, call it **filter.dat**. We must add a **&end** statement not provided by **TRAPLT**. The **filter.dat** file will look like this:

```
&filter
npf= 221, nzph= 1
fil=
0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000,
0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0000,
0.0000, 0.0000, 0.0000, 0.0000, 0.0000, 0.0001, 0.0003, 0.0007,
0.0015, 0.0027, 0.0044, 0.0067, 0.0095, 0.0130, 0.0172, 0.0220,
0.0276, 0.0338, 0.0406, 0.0481, 0.0563, 0.0650, 0.0743, 0.0842,
0.0945, 0.1054, 0.1166, 0.1282, 0.1401, 0.1524, 0.1648, 0.1775,
0.1903, 0.2032, 0.2161, 0.2290, 0.2419, 0.2548, 0.2675, 0.2800,
0.2923, 0.3044, 0.3162, 0.3278, 0.3389, 0.3498, 0.3602, 0.3702,
0.3798, 0.3889, 0.3975, 0.4056, 0.4132, 0.4203, 0.4269, 0.4329,
0.4384, 0.4433, 0.4477, 0.4515, 0.4547, 0.4574, 0.4596, 0.4612,
0.4622, 0.4628, 0.4627, 0.4622, 0.4612, 0.4597, 0.4577, 0.4552,
0.4523, 0.4489, 0.4451, 0.4409, 0.4363, 0.4313, 0.4260, 0.4203,
0.4143, 0.4080, 0.4015, 0.3946, 0.3875, 0.3801, 0.3726, 0.3648,
0.3569, 0.3488, 0.3405, 0.3322, 0.3237, 0.3151, 0.3064, 0.2977,
0.2889, 0.2801, 0.2712, 0.2624, 0.2535, 0.2447, 0.2359, 0.2271,
0.2184, 0.2098, 0.2013, 0.1928, 0.1844, 0.1761, 0.1680, 0.1599,
0.1520, 0.1442, 0.1366, 0.1291, 0.1218, 0.1146, 0.1076, 0.1007,
0.0940, 0.0875, 0.0811, 0.0750, 0.0690, 0.0632, 0.0575, 0.0521,
0.0468, 0.0417, 0.0368, 0.0321, 0.0275, 0.0232, 0.0190, 0.0150,
0.0112, 0.0075, 0.0040, 0.0007, -0.0025, -0.0054, -0.0083, -0.0109,
-0.0135, -0.0158, -0.0181, -0.0201, -0.0221, -0.0239, -0.0255, -0.0271,
-0.0285, -0.0298, -0.0309, -0.0320, -0.0330, -0.0338, -0.0345, -0.0352,
-0.0357, -0.0362, -0.0366, -0.0368, -0.0371, -0.0372, -0.0372, -0.0372,
-0.0372, -0.0370, -0.0368, -0.0366, -0.0363, -0.0359, -0.0356, -0.0351,
-0.0347, -0.0341, -0.0336, -0.0330, -0.0324, -0.0318, -0.0311, -0.0305,
-0.0298, -0.0291, -0.0284, -0.0276, -0.0269, -0.0261, -0.0254, -0.0246,
-0.0239, -0.0231, -0.0223, -0.0216, -0.0208, -0.0200, -0.0193, -0.0185,
-0.0178, -0.0171, -0.0164, -0.0156, -0.0149,
&end
```

The commands for the alternative would be:

```
traplt bfilbdum.seg 0.09 .2 1 0 1
bftr c008.seg 1 filter.dat
bequ bftrc008.seg 0 1.
bplt bequbftr.seg 2 0 0 1 100 0.0 1. 1 2 200
```

NOTE: The TRAPLT command above does not start listing at 0.0 seconds, but at .09 seconds. This produces a namelist file with less delay, and this is evident comparing the two different approaches.

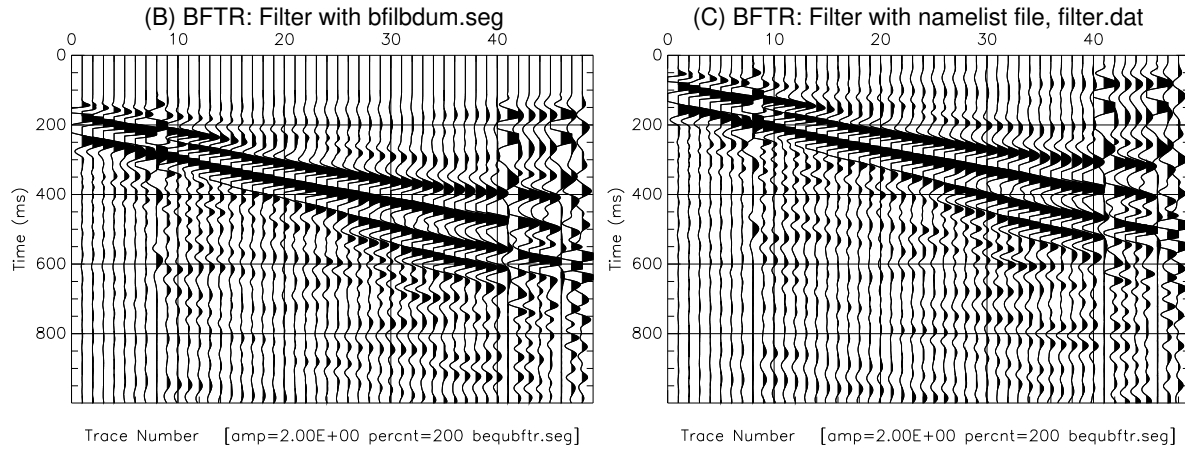


Figure 96: (B) BFTR: same as in Figure 95B. (C) BFTR: Filter field data with namelist file, **filter.dat**. Input data **c008.seg** are same as in Figure 91A. Note the different delay in time.

### 12.3.6 BWHT

Data are whitened (increased bandwidth). The user defines a number of overlapping frequency bands which are individually subjected to Automatic Gain Control (AGC), and then reassembled into a whitened product. Highly nonlinear, but may reveal details in the data by overcoming dynamic range limitations in traditional plots of data. The command line arguments:

```
bwht infil twide fcent bwdth froll
```

```
infil:   =input file to be filtered
twide:   =AGC window length in sec.
fcent:   =center frequency (Hz)
bwdth:   =bandwidth (Hz)
froll:   =roll off (Hz)
```

EXAMPLE: `bwht c008.seg .4 50. 80. 10.`

We can view the **bwhtc008.lst** file to see the filter details. See Figure 97.

```
Parameters:
twide=    0.40
freqc=    50.00
bwdth=    80.00
deltf=    10.00
nfilt=    9
number of points in filter=    201
  J    F_Center (Hz)
  1    10.00
  2    20.00
  3    30.00
  4    40.00
  5    50.00
  6    60.00
  7    70.00
  8    80.00
  9    90.00
```

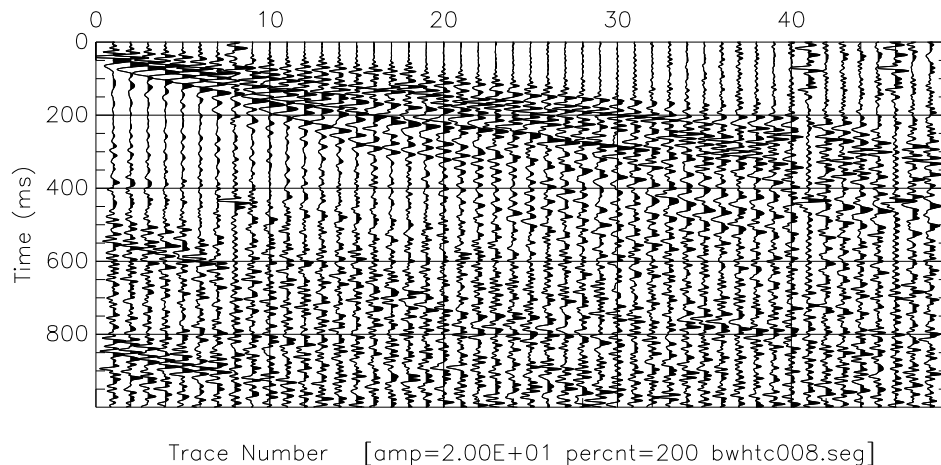


Figure 97: BWHT: 0.4 second AGC window, 50 Hz center, 80 Hz bandwidth, 10 Hz rolloff. Input data **c008.seg** are same as in Figure 91A.

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```
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under certain conditions; type 'show c' for details.
```

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